



Status and Perspectives of the Gaseous Photon Detectors Technologies

Fulvio Tessarotto (INFN - Trieste)



Gaseous Photon Detectors

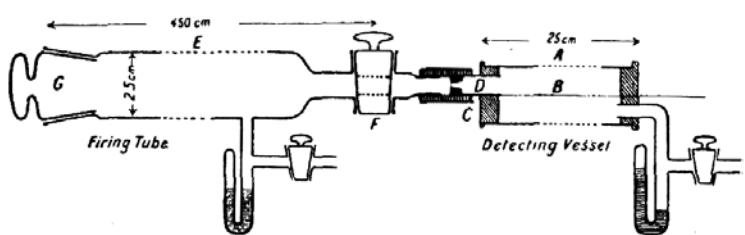
- Historical overview of gaseous detectors
- MWPCs with CsI Photocathodes
- GEM-based PDs
- THGEM-based PDs
- Other architectures
- Gaseous detectors for visible light
- Cryogenic gaseous photon detectors
- Large area coverage



Glorious tradition: 100 years of gaseous detector developments



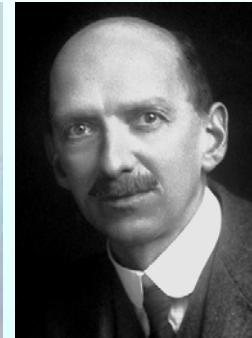
1908: FIRST WIRE COUNTER
USED BY RUTHERFORD IN THE STUDY OF NATURAL RADIOACTIVITY



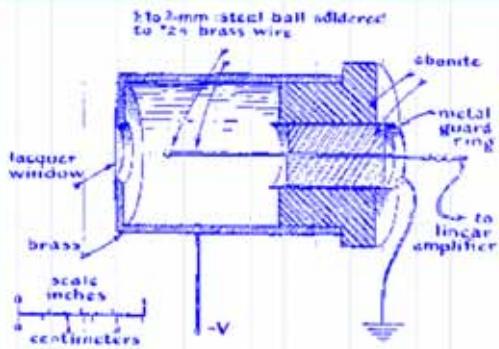
E. Rutherford and H. Geiger,
Proc. Royal Soc. A81 (1908) 141



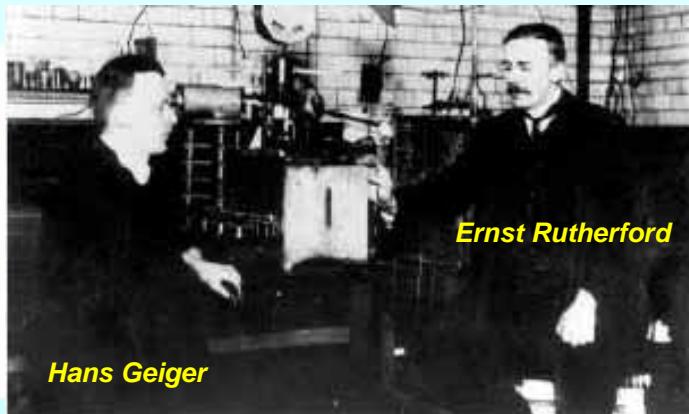
1911: CLOUD CHAMBER



1928: GEIGER COUNTER
SINGLE ELECTRON SENSITIVITY



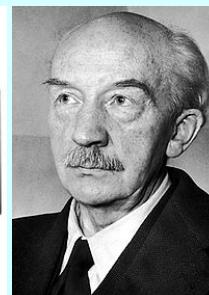
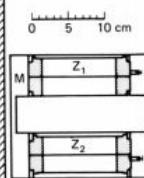
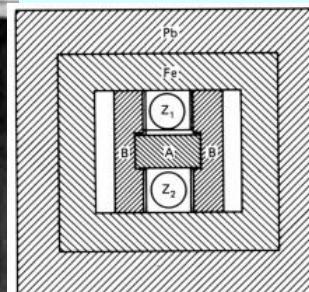
H. Geiger and W. Müller,
Phys. Zeits. 29 (1928) 839



Ernst Rutherford

Hans Geiger

COINCIDENCE METHOD

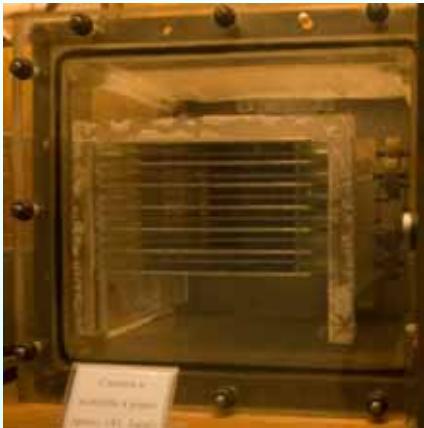


Walther Bothe
Nobel Prize in 1954



Glorious tradition: 100 years of gaseous detector developments

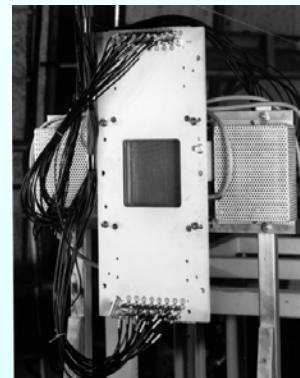
SPARK CHAMBER



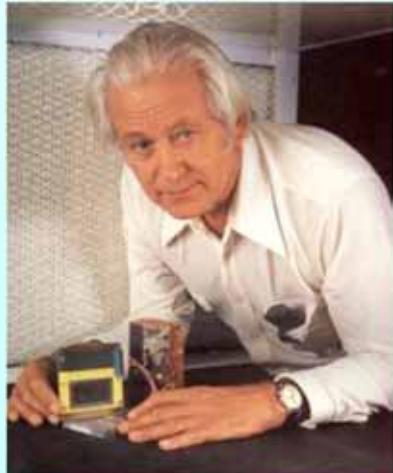
1952: BUBBLE CHAMBER



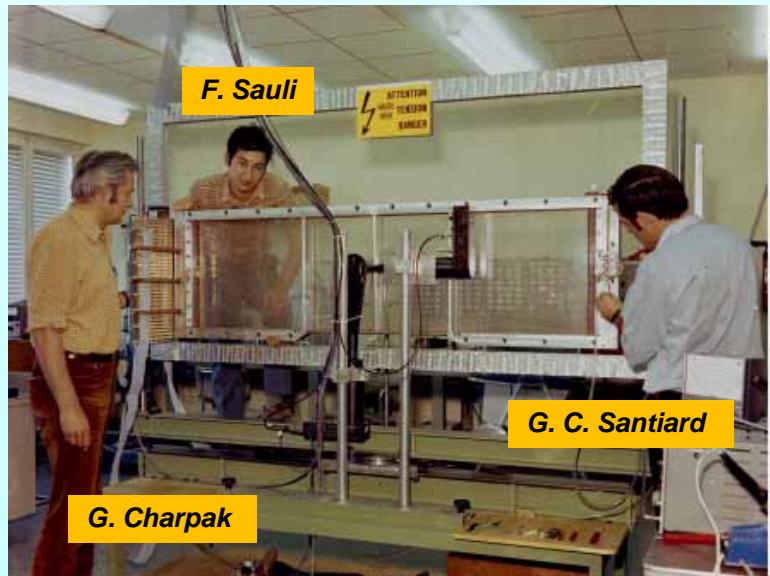
1968: MULTIWIRE PROPORTIONAL CHAMBER



*Donald A. Glaser
Nobel Prize in 1992*



*George Charpak
Nobel Prize in 1992*

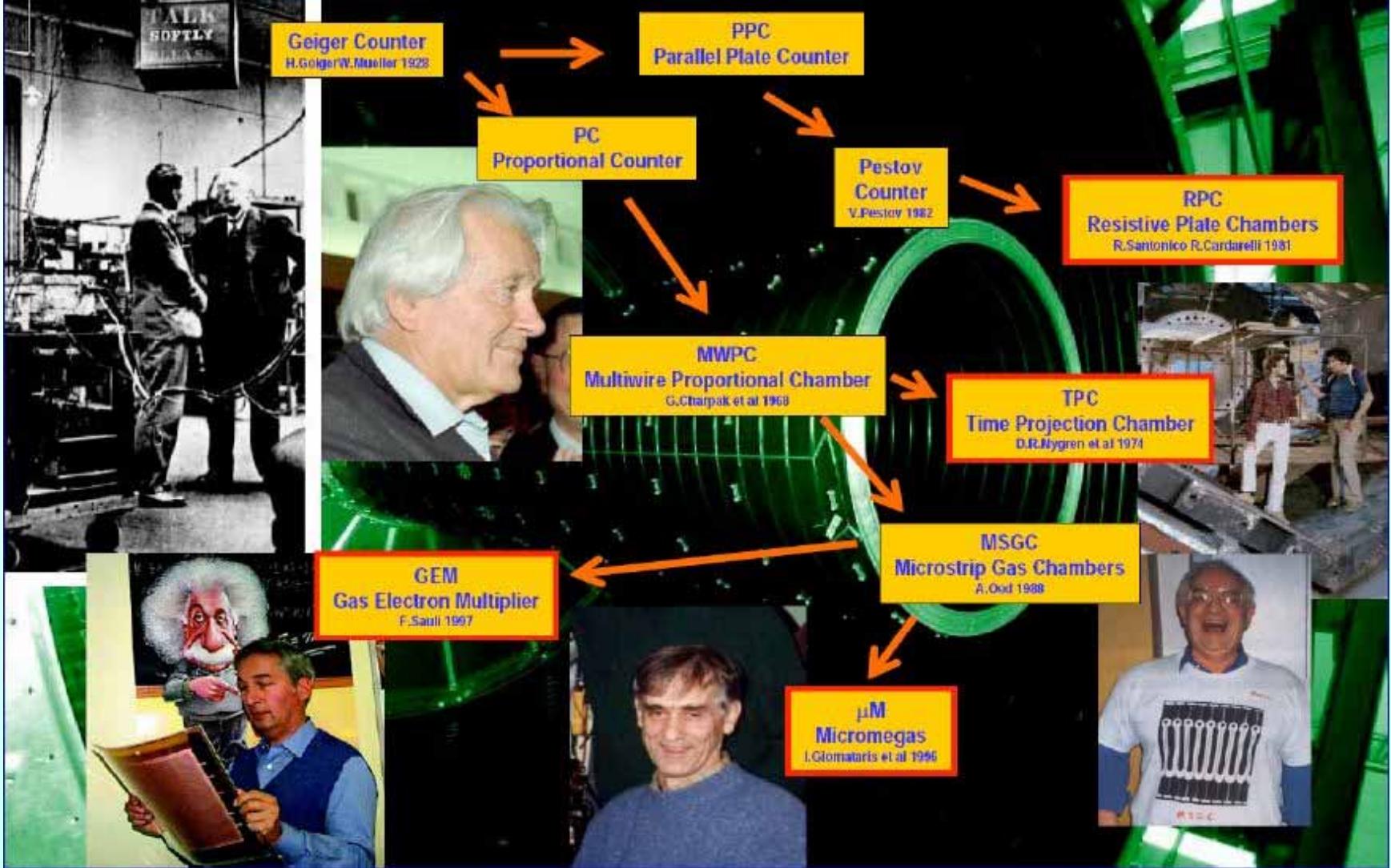


F. Sauli

G. C. Santiard

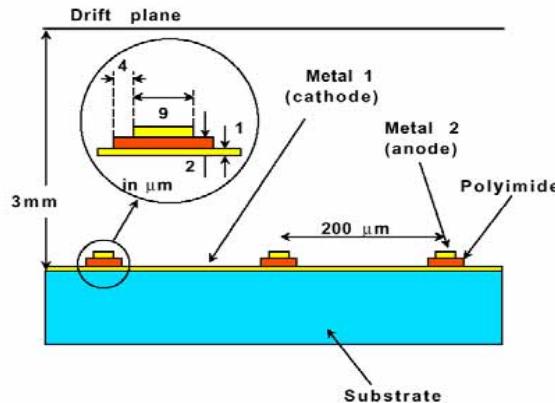
G. Charpak

Glorious tradition: 100 years of gaseous detector developments

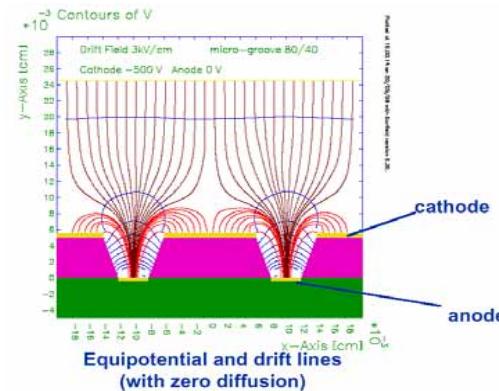


Many different MPGDs have been developed

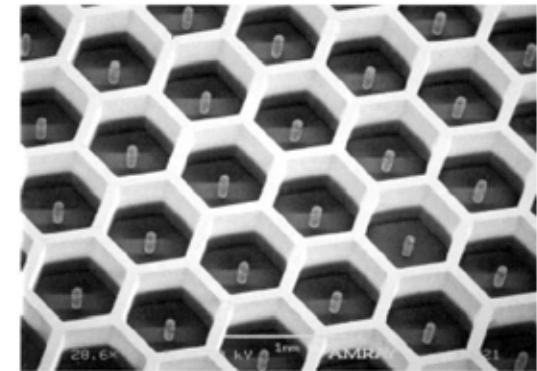
MICRO-GAP CHAMBER



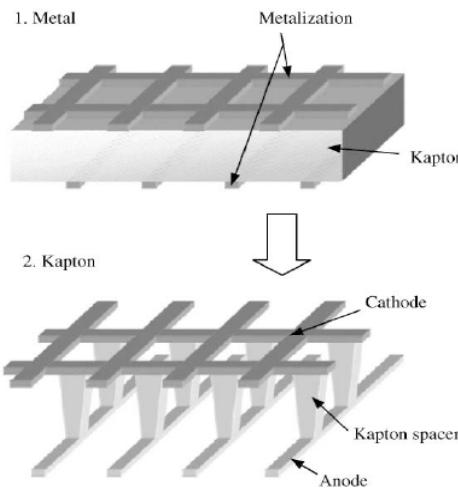
MICRO-GROOVE CHAMBER



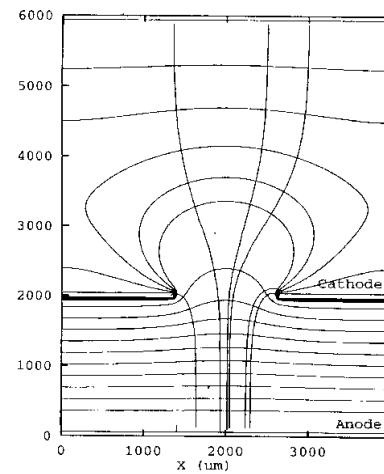
MICRO-PIN ARRAY



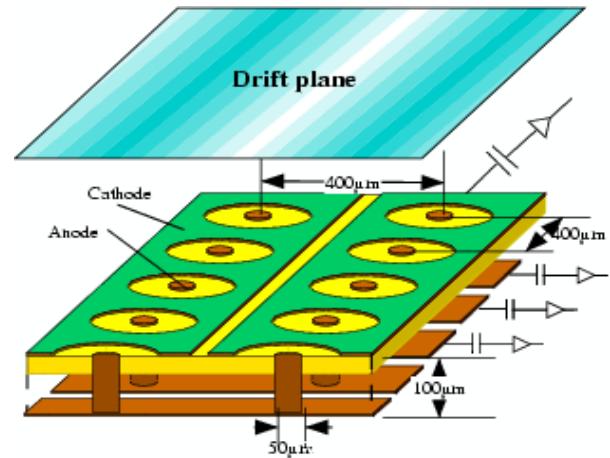
MICRO-WIRE CHAMBER



COMPTEUR A TROUS



MICRO-PIXEL CHAMBER

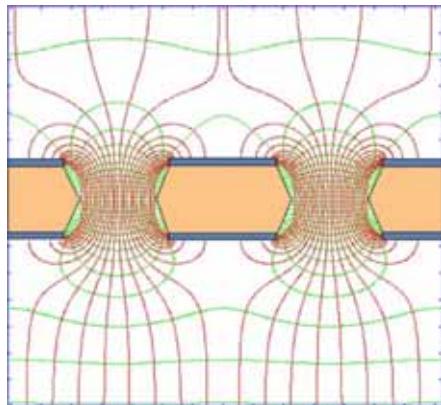


2 very solid, fully established technologies:

MICRO MEsh GAseous Structure (MICROMEGAS)

Thin gap Parallel Plate Chamber: micromesh stretched over readout electrode.

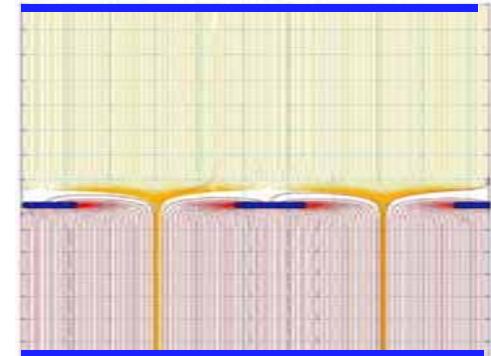
Y. Giomataris et al., Nucl. Instr. and Meth. A376(1996)29



GAS ELECTRON MULTIPLIER (GEM)

Thin, metal-coated polymer foil with high density of holes, each hole acting as a proportional counter.

F. Sauli, Nucl. Instrum. Methods A386(1997)531

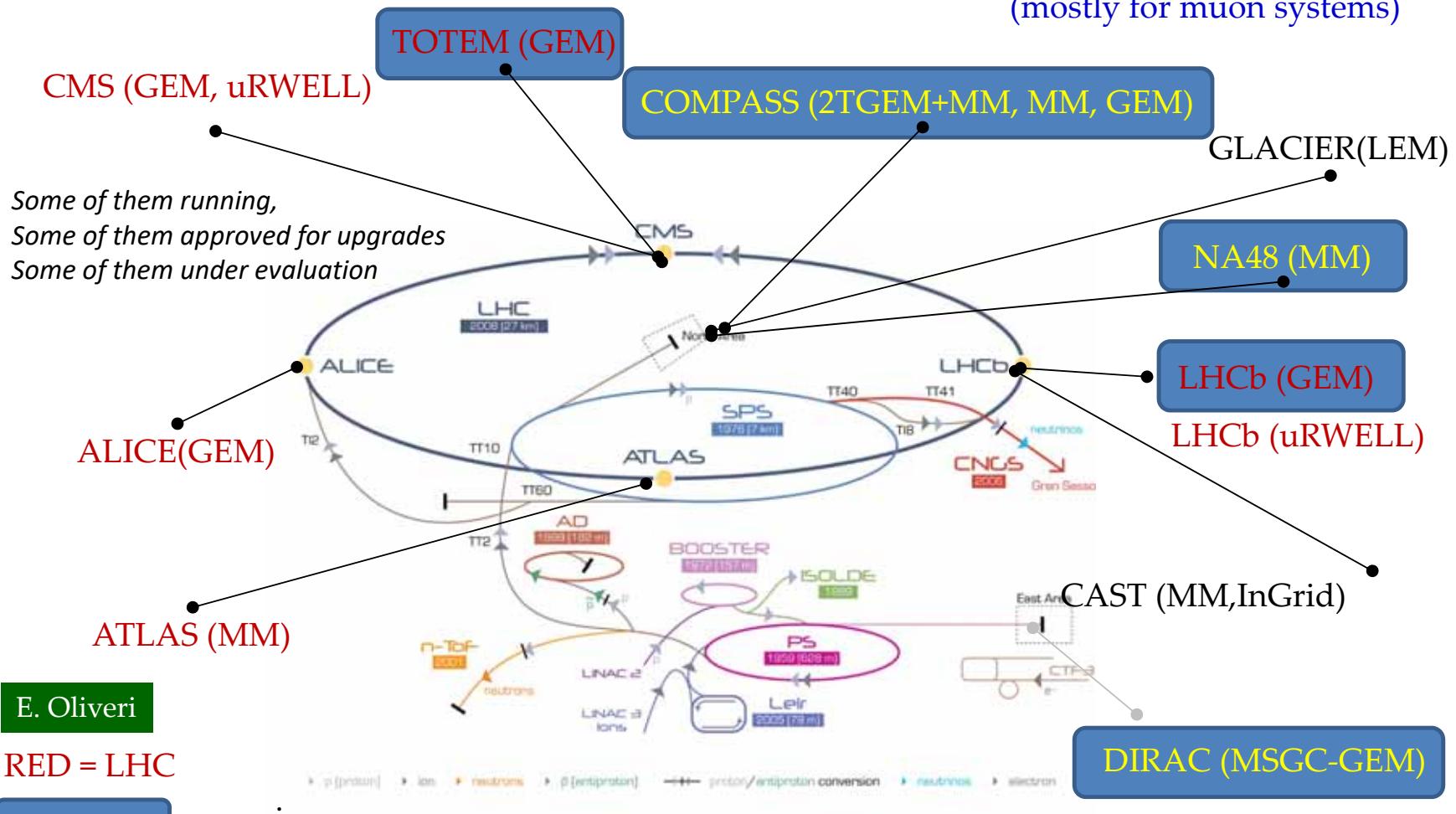


HIGHER LUMINOSITIES, HIGHER PRECISION EXPERIMENTS

- 1. MPGDs allow for**
 - High rates (granularity & occupancy, signal formation time)**
 - Fine space resolution**
- 2. Technological maturity and accurate engineering FUNDAMENTAL for successful MPGDs**

MPGDs at CERN Experiments

Despite original MSGC discharge problems, MPGDs have been chosen for all LHC upgrades
(mostly for muon systems)



MPGDs and RD51

- **MPGDs exist thanks to a few genial inventors**
- **Great progress due to engineering for the use in experiments**
- **Recently, a fundamental boost thanks to RD51:**
“RD51, aims at facilitating the development of advanced gas-avalanche detector technologies and associated electronic-readout systems, for applications in basic and applied research.”

RD51 serves as an access point to MPGD “know-how” for the world-wide community

Unique in providing:

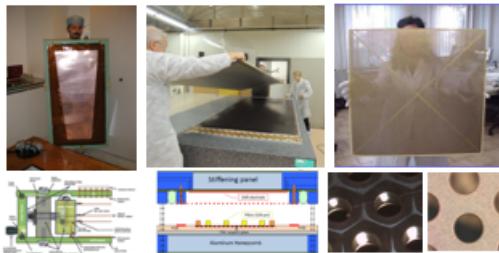
- Space and resources for non – project related R&D
- tools for the word-wide MPGD community AND BEYOND



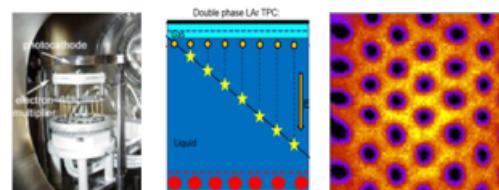
~ 80 Institutes from 4 continents: Europe, Nord and South America, Asia, Africa
~ 450 physicists

RD51 Working groups

Technological Aspects and Development of New Detector Structures



Common Characterization and Physics Issues



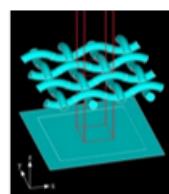
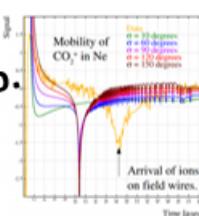
Academia-Industry Matching Events, Training, Education



WG3/NEW WG:



Simulations and Software Tools



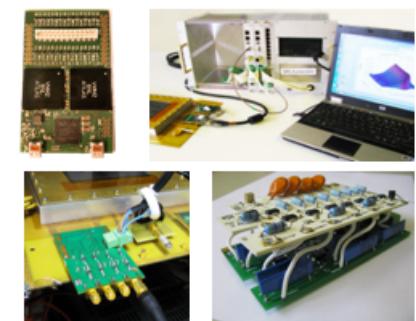
Common Facilities : Test Beam and Laboratory



Production, quality control, industrialization



MPGD Related Electronics



photon conversion and Cherenkov light



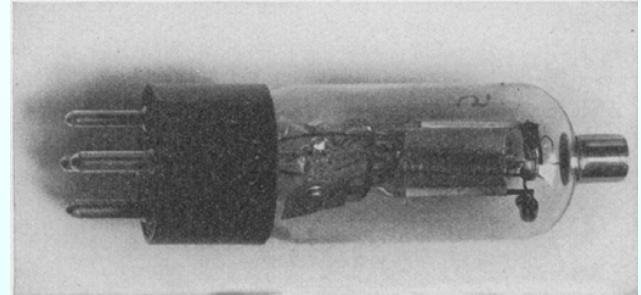
John Sealy Townsend



Heinrich Rudolf Hertz
photoelectric effect, 1887



A. Einstein, Nobel Prize in 1921



Iams, H. E. and B. Salzberg, "The secondary emission phototube," Proc. IRE **23**, 55 (1935).



Pavel Cherenkov 1904-1990



Ilya Frank and



Igor Tamm



Arthur Roberts 1912-2004



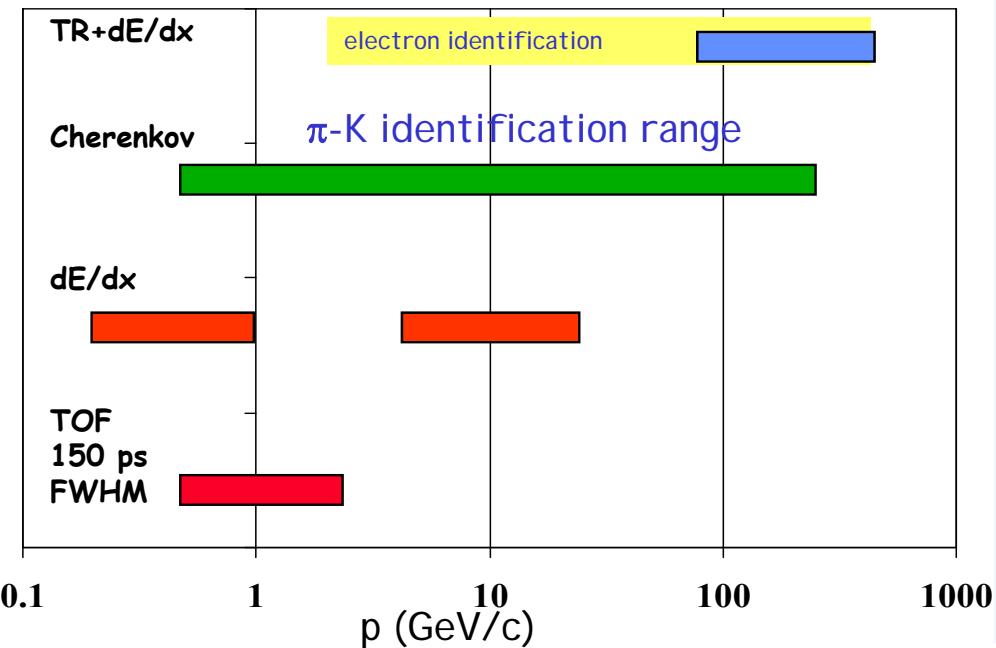
Tom Ypsilantis 1928-2000

Nobel Prize in 1958

Motivation

- need for π -K identification from HEP Experiments
- Large momentum acceptance → Cherenkov angle measurement technique
- Large angular acceptance → large area of efficient single photon detection

Particle Identification Techniques:

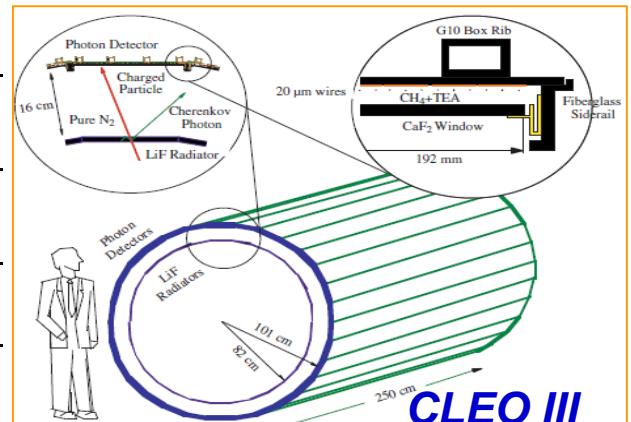
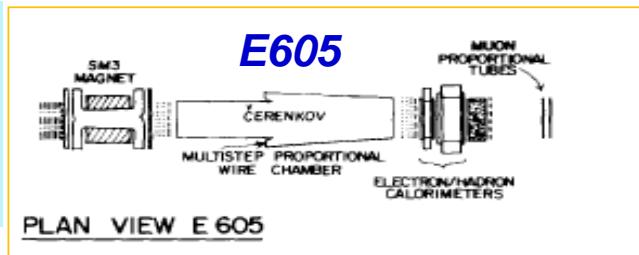
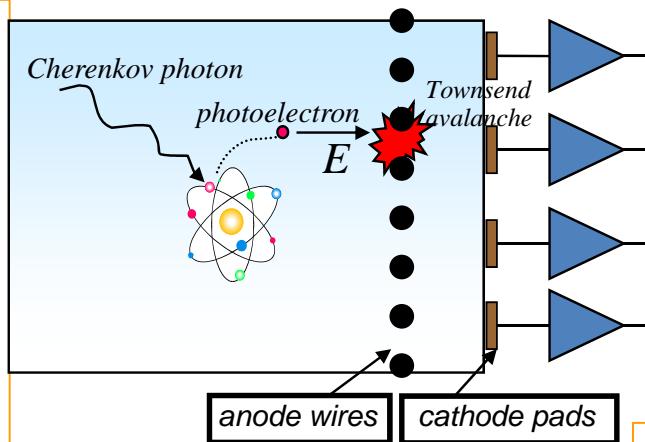
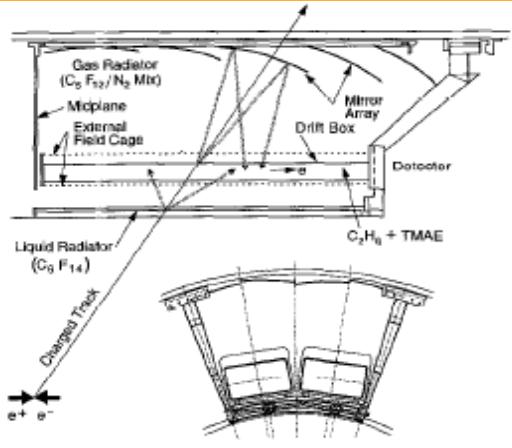


- 1970s: large area position sensitive gaseous detectors available
- Suitable photo-ionizing agent:
benzene: Seguinot-Ypsilantis NIM 142 (1977) 377,
TEA (7.6 eV) NIM 173 (1980) 283,
TMAE (5.3 eV) NIM 178 (1980) 125.
- a gas gain high enough to detect single photoelectrons
- conflicting requirements because of the copious UV emission by the multiplication avalanche.
- solution: multistep avalanche chamber
(Charpak-Sauli Phys. Lett. B 142 (1977) 377) or TPC

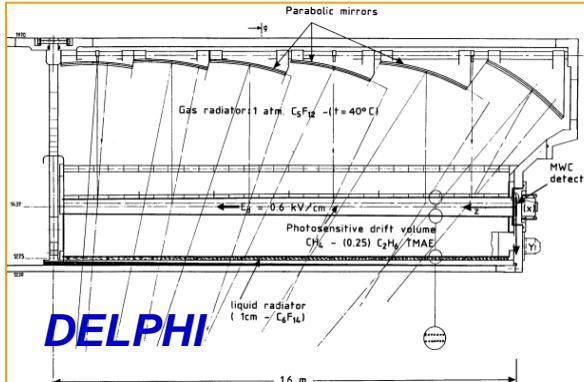
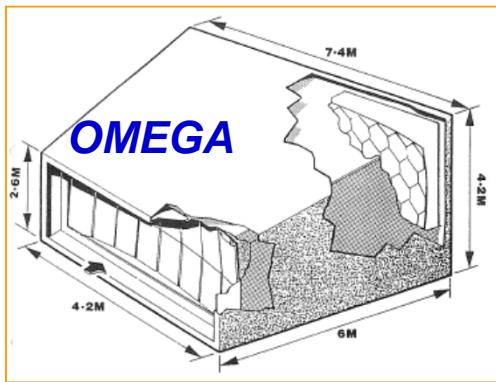
Gaseous detectors: 1) cheap, 2) magnetic insensitive, 3) low material budget

RICH with large area gaseous PD's 1st generation: photoconverting vapours

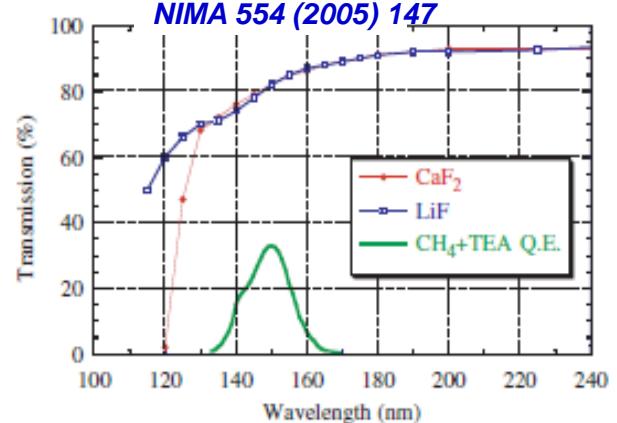
SLD - CRID

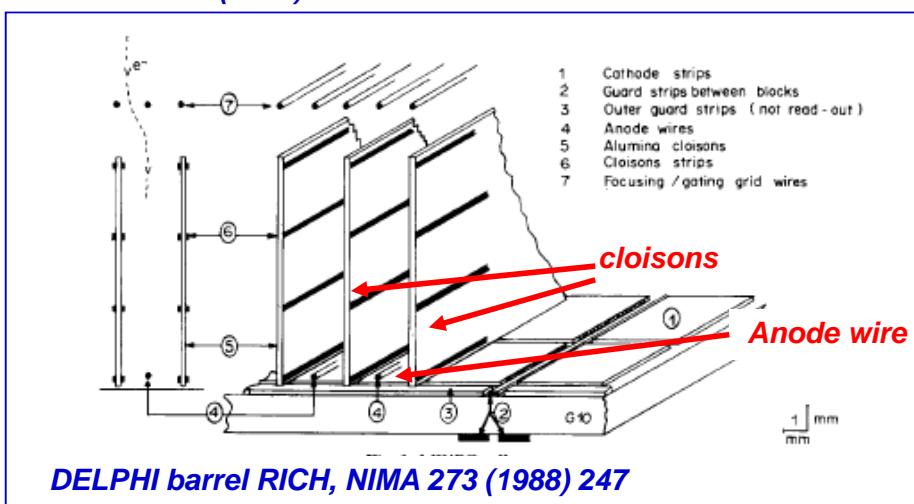
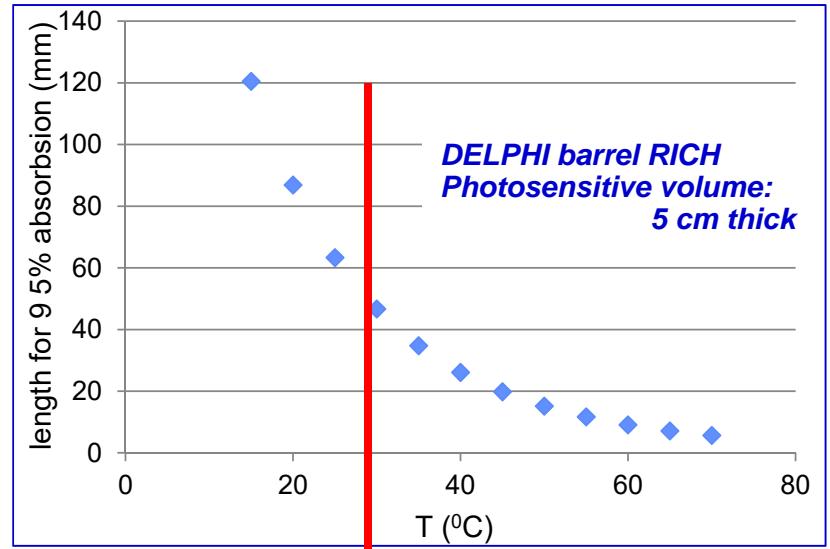
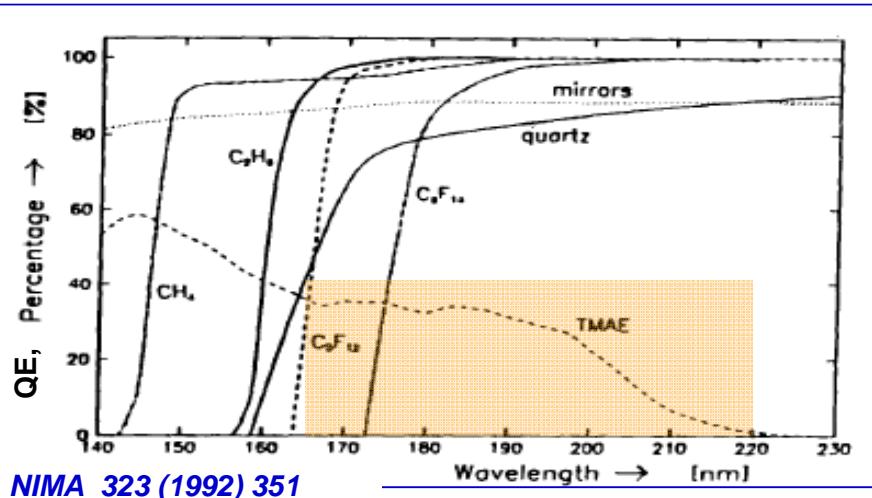


TEA(Tri-Ethyl-Amine): $E_f=7.6$ eV



TMAE(Tetrakis-Dimethylamine-Ethylene): $E_f=5.3$ eV





- thick photosensitive volume (slow photon detectors, parallax error)
- heating and temperature control ($T_{\text{bubbling}} < T_{\text{operation}}$)
- photon feed-back from amplification region (protections)
- chemically extremely reactive



Thin CsI film provides UV photoconversion



1956:
CsI layer has large
QE for photons
with $h\nu > 6$ eV
(Philipp and Taft)

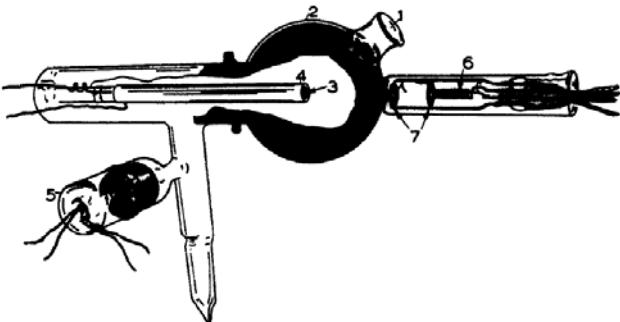


FIG. 1. Cutaway sketch of phototube; (1) 9741 glass bubble window, (2) graphite coated collector sphere 4 inches in diameter, (3) $\frac{1}{2}$ inch glass tube, platinum painted, (4) nickel sleeve insulated from tube by glass beads, (5) ion gauge, (6) evaporating cylinder and helical platinum heater, (7) collimating shields.

J. Phys. Chem. Solids. Pergamon Press 1956. Vol. 1. pp. 159–163.

PHOTOELECTRIC EMISSION FROM THE VALENCE
BAND OF CESIUM IODIDE

H. R. PHILIPP AND J. E. A. TAFT

General Electric Research Laboratory, Box 1088, Schenectady, New York

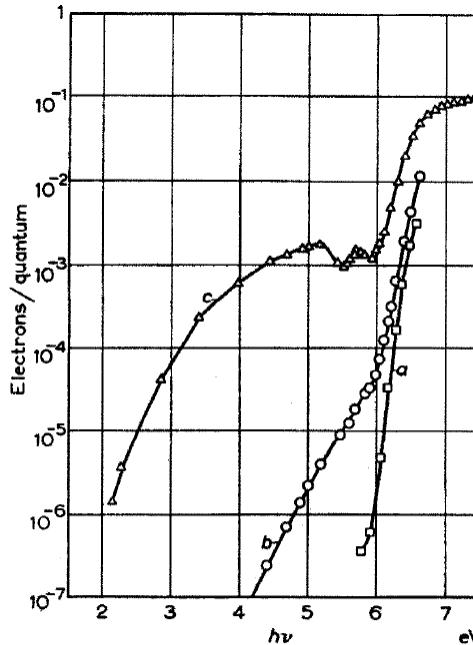


FIG. 2. Spectral distribution of the photoelectric yield for CsI surfaces: (a) thick film, (b) single crystal, (c) thin film evaporated in presence of excess Cs.

vacuum operated
photocathodes
were developed for space
astronomy:

G.R.Carruthers,
Appl. Opt. 8 (1969) 633
Appl. Opt. 12 (1973) 2501
Appl. Opt. 14 (1975) 1667

The first position sensitive gas
detectors with CsI
photocathodes were developed
at the end of the 80s:

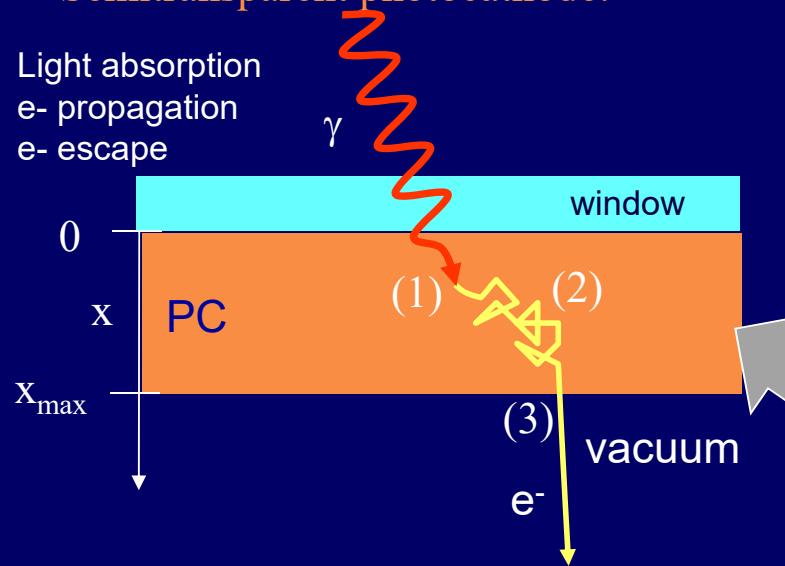
G.Charpak et al., Proceedings of
Symposium on Particle
Identification at
High Luminosity Hadron
Colliders, Fermilab, Batavia, IL,
1989, p. 295.

J.Séguinot et al., Nucl. Instr. and
Meth. A 297 (1990), p. 133

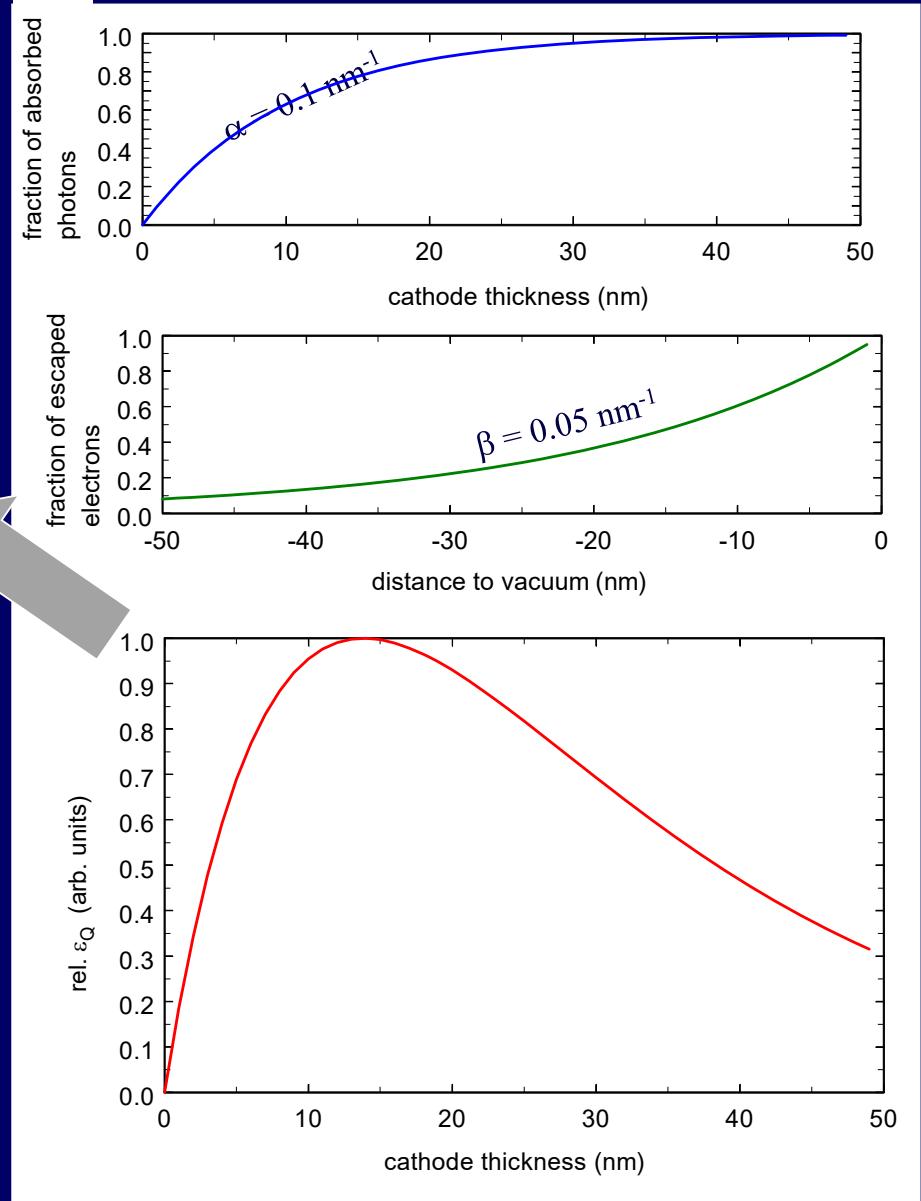
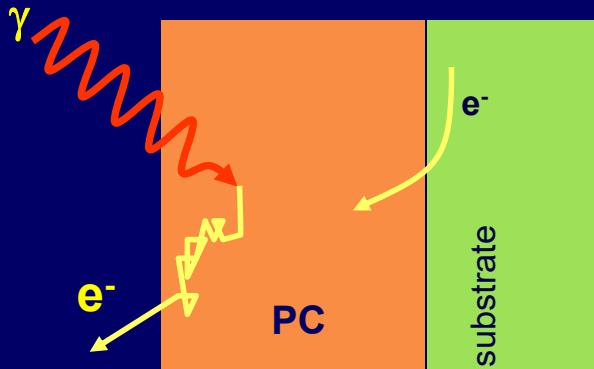
Thin film photo cathodes

Semitransparent photocathode:

- (1) Light absorption
- (2) e- propagation
- (3) e- escape



Reflective photocathode:



Reflective CsI photocathodes

70

J. Séguinot et al./Nucl. Instr. and Meth. in Phys. Res. A 371 (1996) 64–78

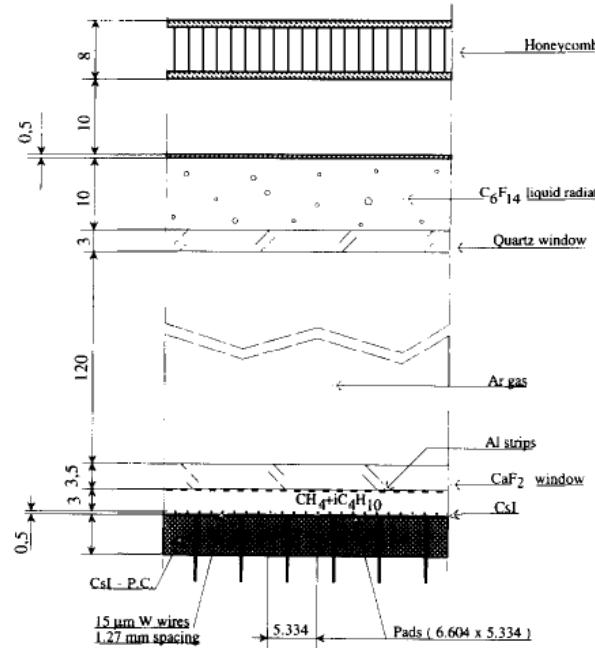
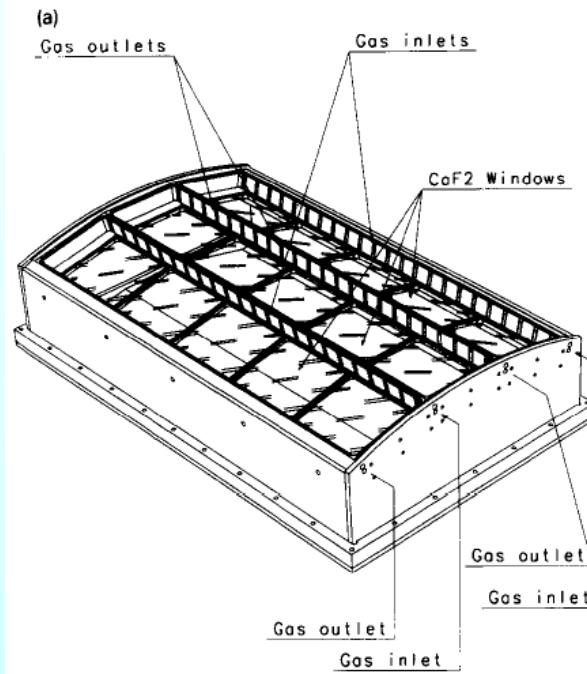
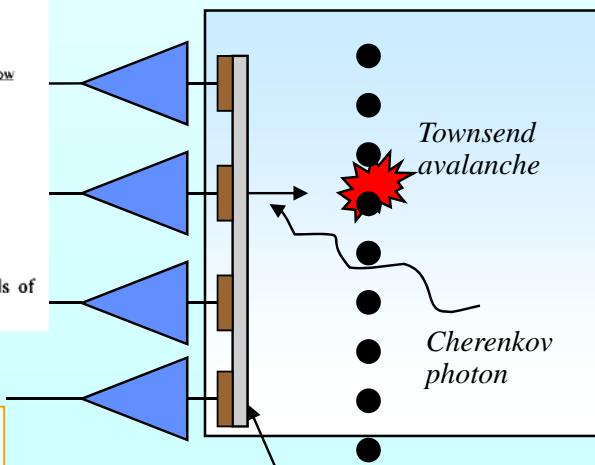


Fig. 9. Cross section on the Fast-RICH Prototype showing the details of radiator and detector geometry.

J. Séguinot et al., Nucl. Instr. and Meth. A 371 (1996), p. 64.

This technique was pioneered by T. Ypsilantis and J. Séguinot, who built an extraordinary prototype and correctly concluded the technology was not adequate for the extreme time performance they needed

CsI is highly reactive with moisture:
it took many years to develop appropriate substrate preparation, deposition method, handling technology for high QE gaseous PDs





RD26: the technology of MWPCs + CsI

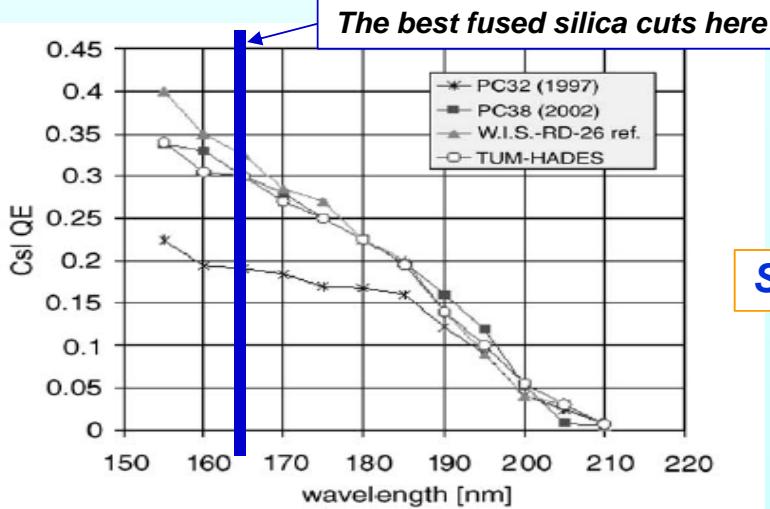


Fig. 1. The QE of CsI PCs produced at CERN for ALICE and at TUM for HADES, compared to that measured at the W.I.S. on small samples (reference for RD-26). PC32 is one of the four PCs equipping the ALICE-RICH prototype used in STAR at BNL.

A. Di Mauro, NIM A 525 (2004) 173.

1992, F. Piuz et al. Development of large area advanced fast-RICH detector for particle identification at LHC operated with heavy ions

TO ACHIEVE HIGH CsI QE:

Substrate preparation:

Cu clad PCB coated by Ni (7 μm) and Au(0.5 μm), surface cleaning in ultrasonic bath, outgassing at 60 °C for 1 day

Slow deposition of 300 nm CsI film:

1 nm/s (by thermal evaporation or e⁻-gun) at a vacuum of $\sim 10^{-7}$ mbar, monitoring of residual gas composition

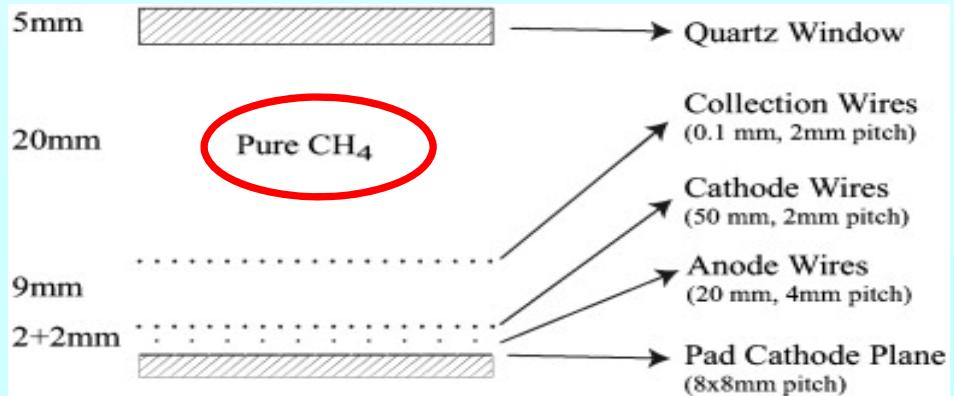
Thermal treatment:

after deposition at 60 °C for 8 h

Careful Handling:

measurement of PC response, encapsulation under dry Ar, mounting by glove-box.

Schematic structure of the COMPASS Photon Detector:

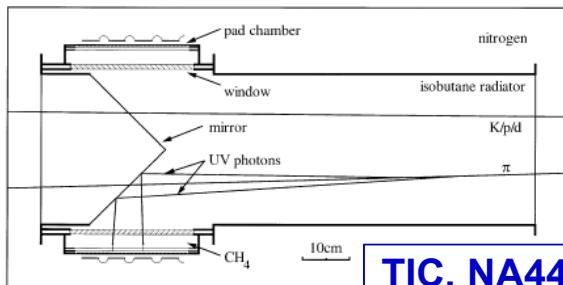




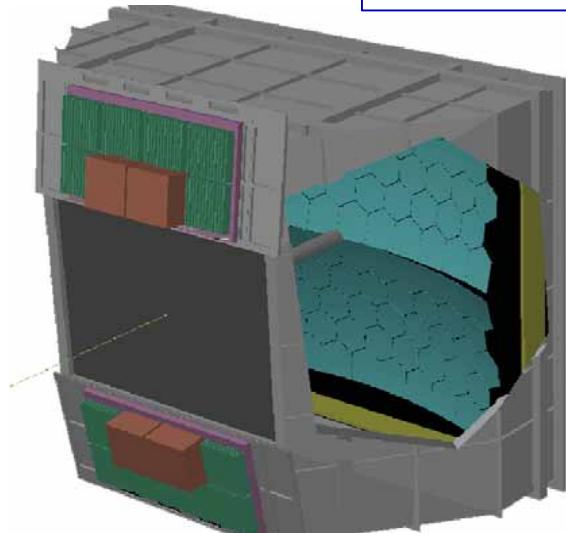
RICH with large area gaseous PD's

2nd generation: MWPC's + CsI

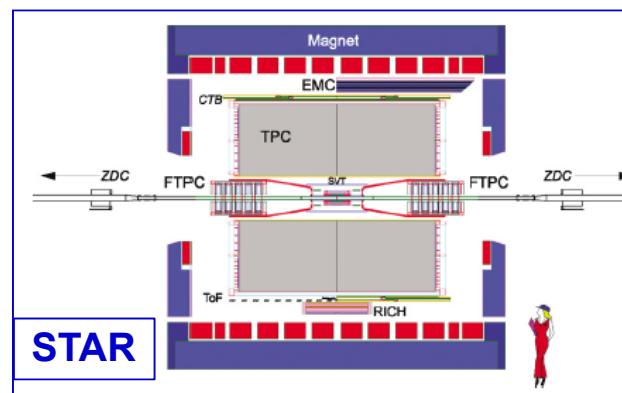
- MWPCs with solid state photocathode (the RD26 effort)



TIC, NA44



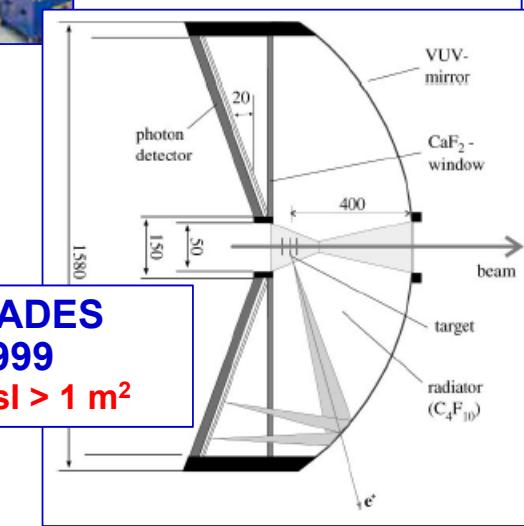
COMPASS RICH-1 2002
CsI > 5 m²



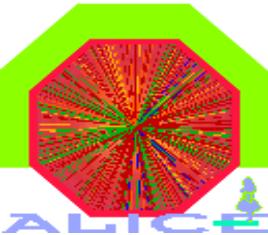
STAR



JLAB-HALL A



HADES
1999
CsI > 1 m²



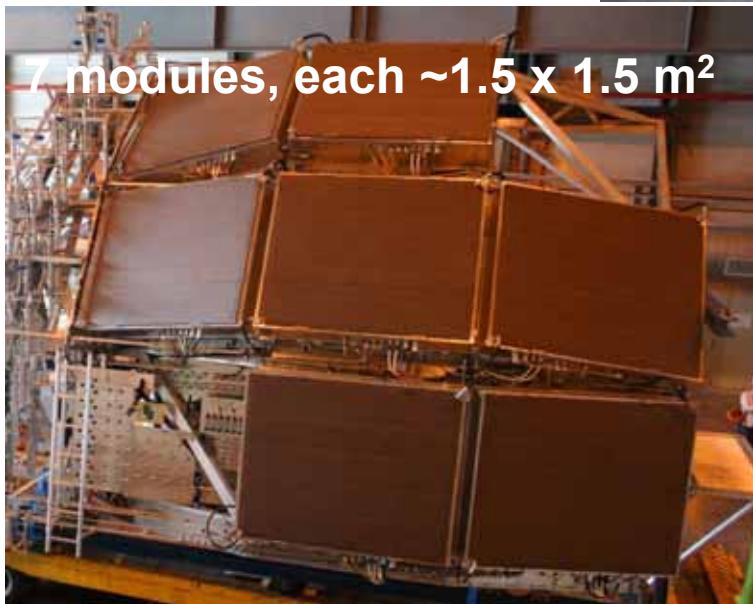
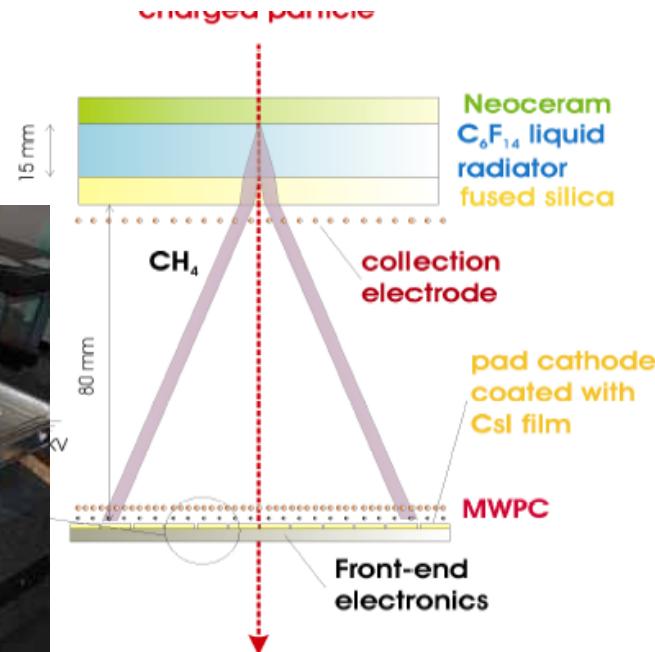
ALICE HMPID

RADIATOR: 15 mm liquid C_6F_{14} ,
 $n \sim 1.2989$ @ 175nm, $\beta_{th} = 0.77$

PHOTON CONVERTER: Reflective
layer of CsI (QE $\sim 25\%$ @ 175 nm)

PHOTOELECTRON DETECTOR:
MWPC with CH_4 at atmospheric
pressure (4 mm gap) **HV = 2050 V.**

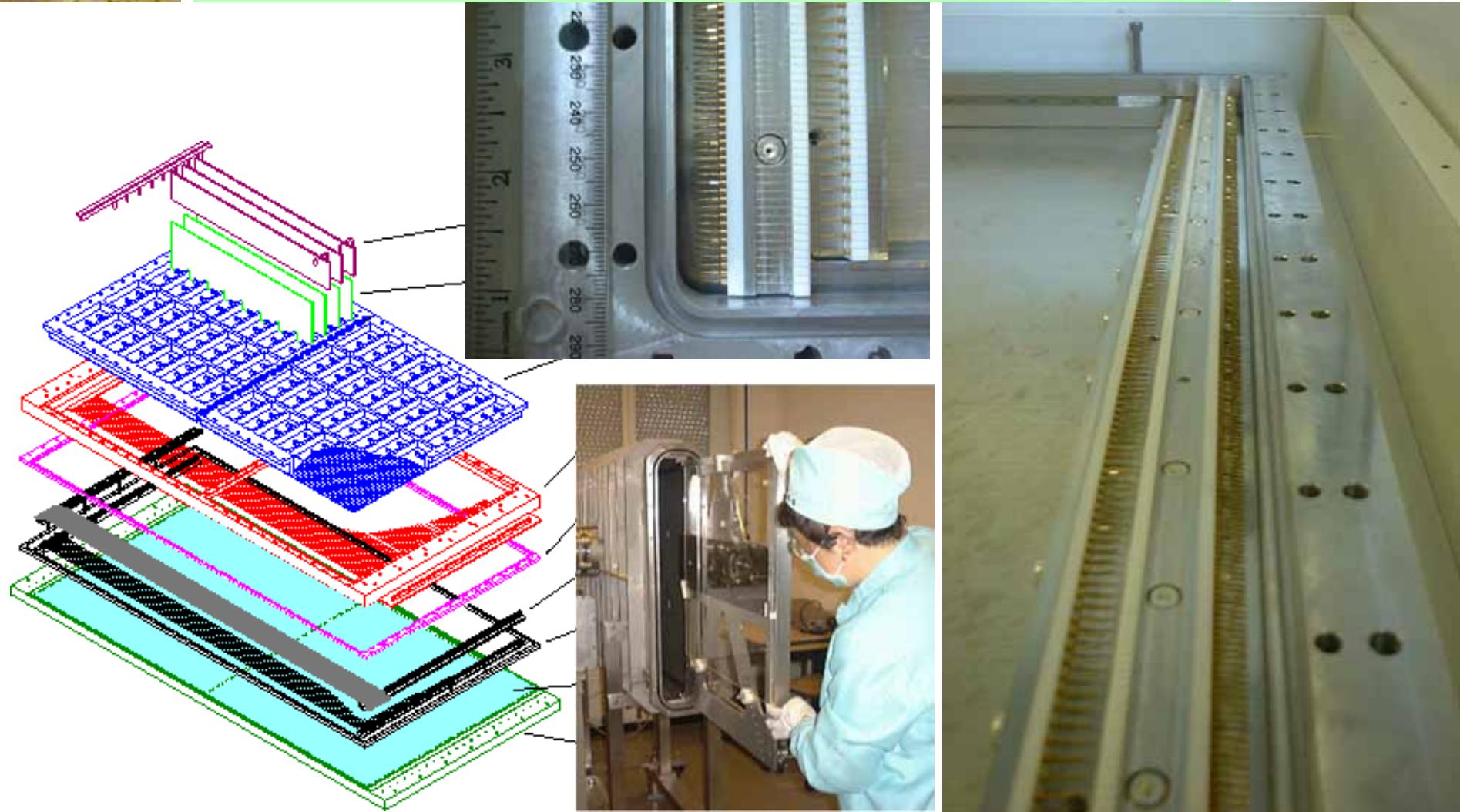
- Analogue pad readout



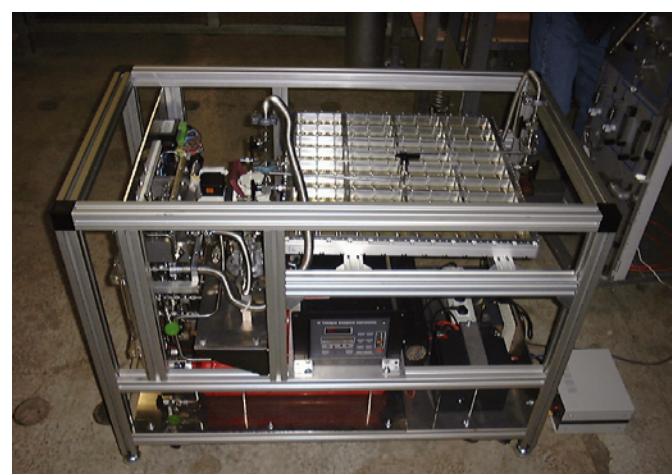
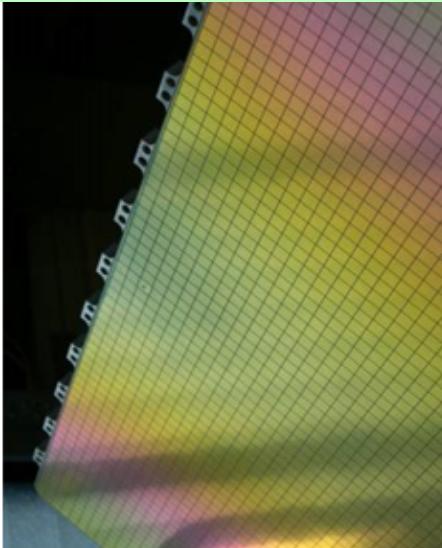
7 modules, each $\sim 1.5 \times 1.5 m^2$



COMPASS MWPC's with CsI

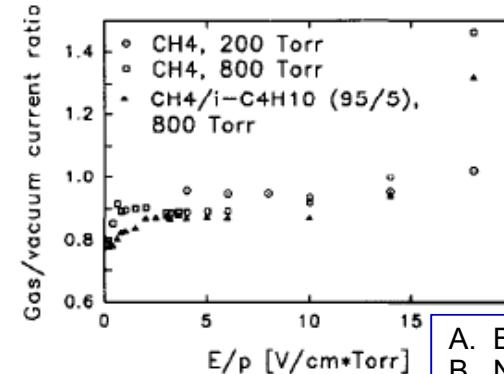
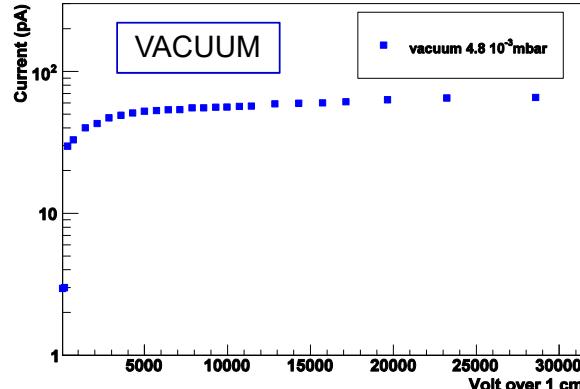


COMPASS photocathodes

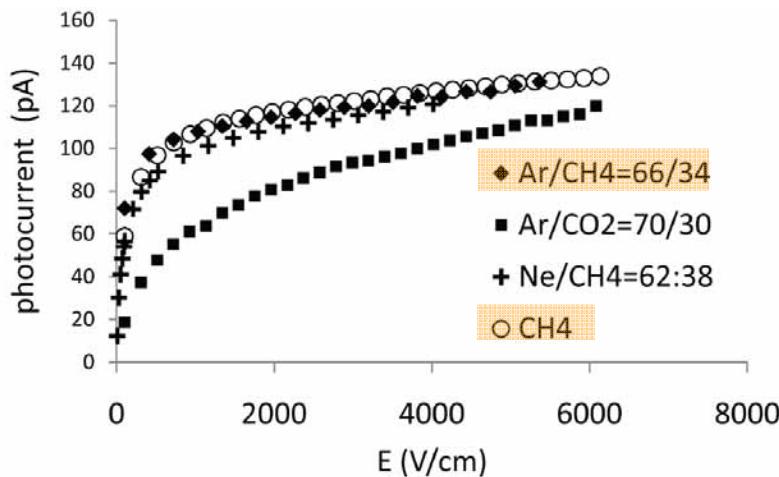


PHOTOELECTRON EXTRACTION

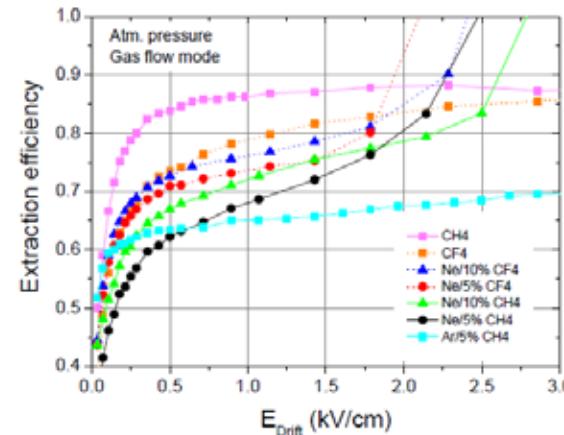
- Photoelectron extraction from a CsI film, the role of gas and E



A. Breskin et al.,
B. NIM A 367 (1995) 342



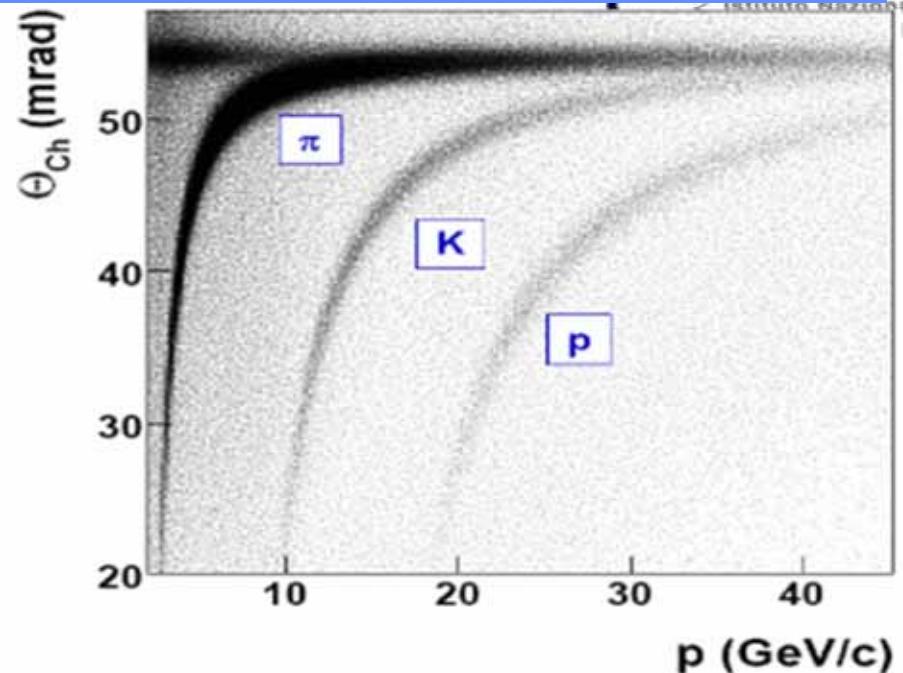
M. Alexeev et al., NIM A (2010) in press



C. D. R. Azevedo et al., 2010 JINST 5 P01002

MWPCs with CsI have good performance

- photons / ring ($\beta \approx 1$, complete ring in acceptance) : **14**
- $\sigma_{\theta\text{-ph}}$ ($\beta \approx 1$) : **1.2 mrad**
- σ_{ring} ($\beta \approx 1$) : **0.6 mrad**
- 2σ π - K separation @ **43 GeV/c**
- PID efficiency \sim **95%** for $\theta_{\text{ch}} > 30$ mrad
except for the very forward region



After a long fight for increasing electrical stability at high m.i.p. rates and systematic studies at the CERN G1 F we came to the same conclusion as Ypsilantis and Seguinot:

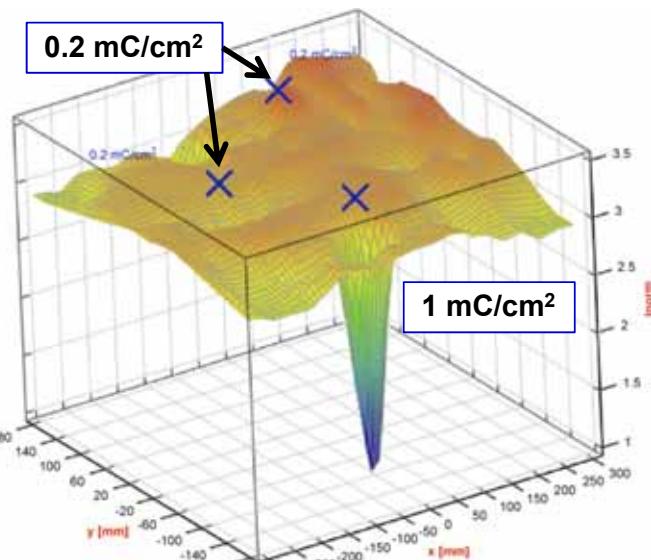
J. Seguinot et al., NIM A 371 (1996), 64:

CsI-MWPC with 0.5 mm gap to minimize ion collection time, fast front-end electronics (20 ns int. time):

stable operation is not possible at 10^5 gain because of photon feedback, space charge and sparks

MWPCs with CsI: the limits

- Severe recovery time (~ 1 d) after detector trips
 - Ion accumulation at the photocathode
 - Feedback pulses
 - Ion and photons feedback from the multiplication process
 - Aging after integrating a few mC/cm^2
 - Ion bombardment of the photocathode (IBF > 50%)
- moderate gain: $< 10^5$
(effective gain: $< 1/2$)
not fast



H. Hoedlmoser et al., NIM A 574 (2007) 28.

MWPCs \rightarrow slow signal formation

- + low gain \rightarrow "slow" electronics (signal integration, low noise level)

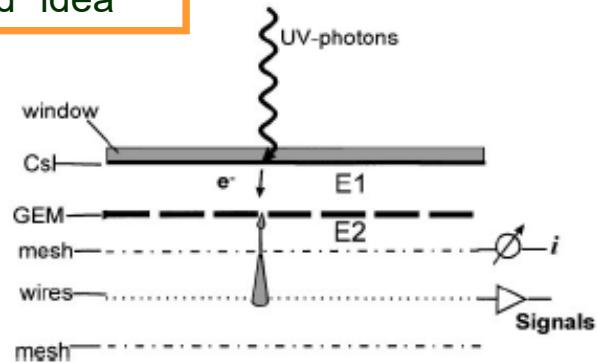
- *Gassiplex FE : integration time $\sim 0.5 \mu\text{s}$, time res $> 1 \mu\text{s}$*
- *APV (COMPASS RICH-1 upgrade) : resolution $\sim 400 \text{ ns}$*
- *Detector memory, i.e. not adequate for high rates*

MPGD technologies are needed to overcome these limits

GEM-based PDs

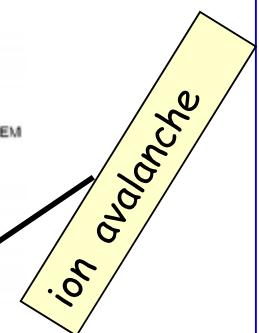
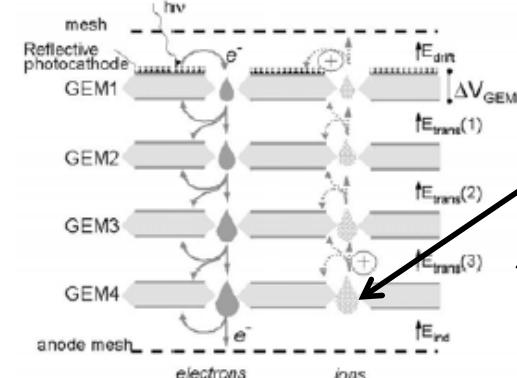
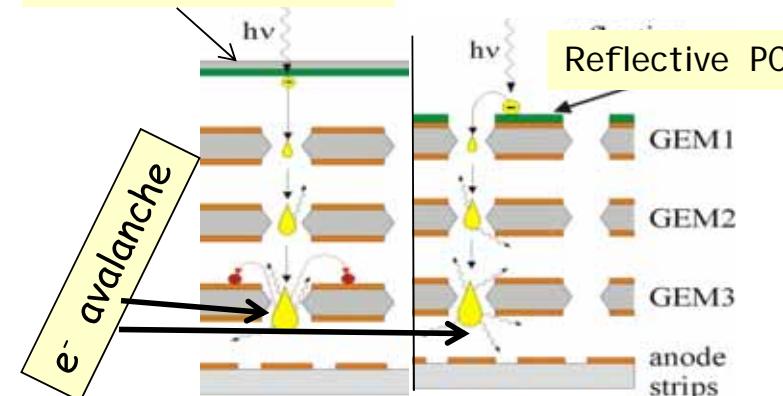
**NO photon feedback
Reduced ion feedback**

An “old” idea



R. Chechik et al., NIM A 419 (1998) 423

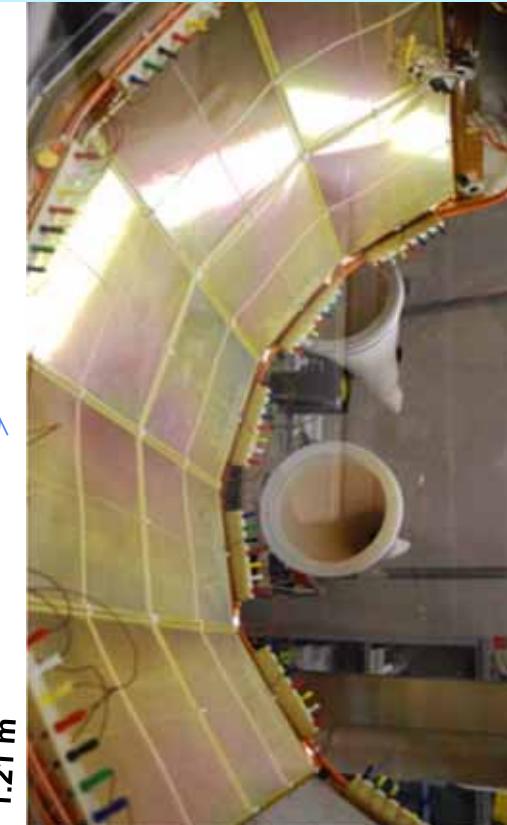
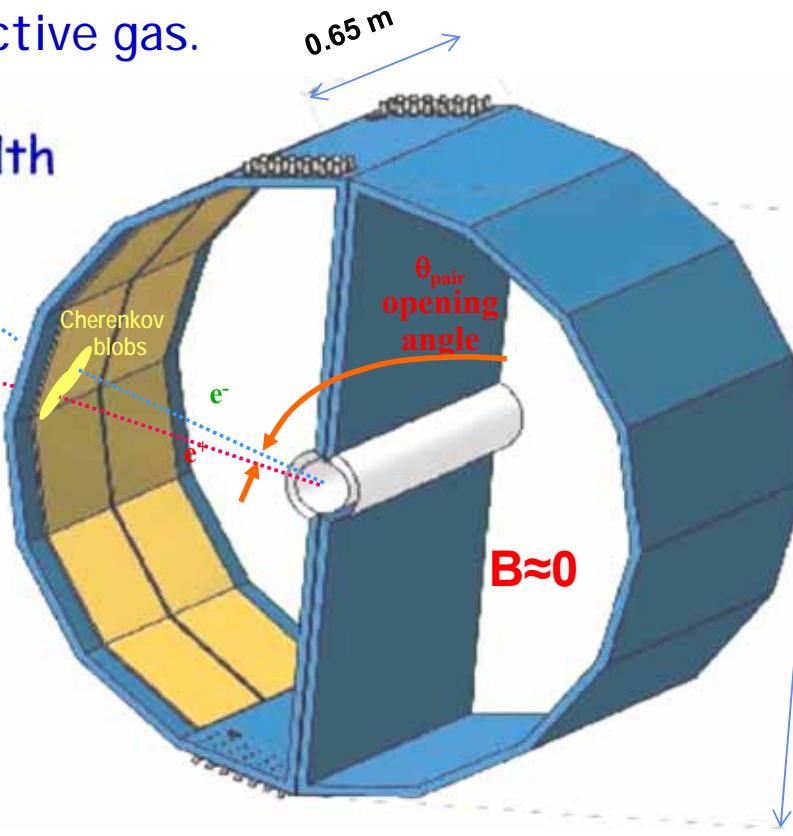
Semi-transparent PC



A. Breskin and R. Chechik, NIM A 595 (2008) 116

HBD- Cherenkov detector with GEMs +CsI

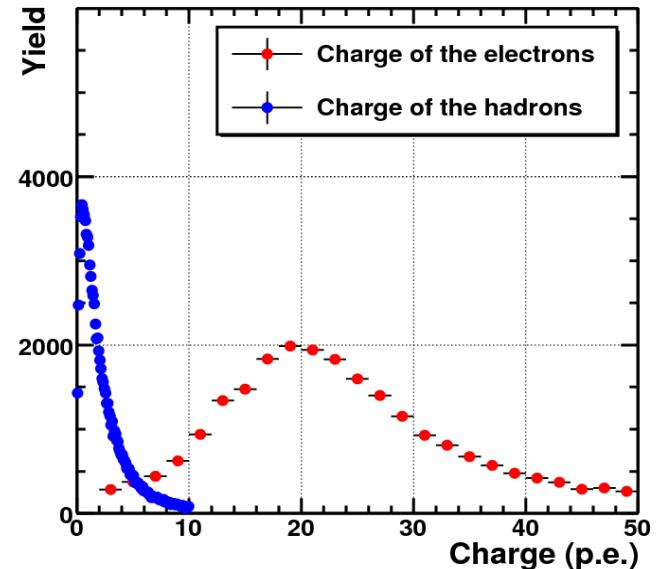
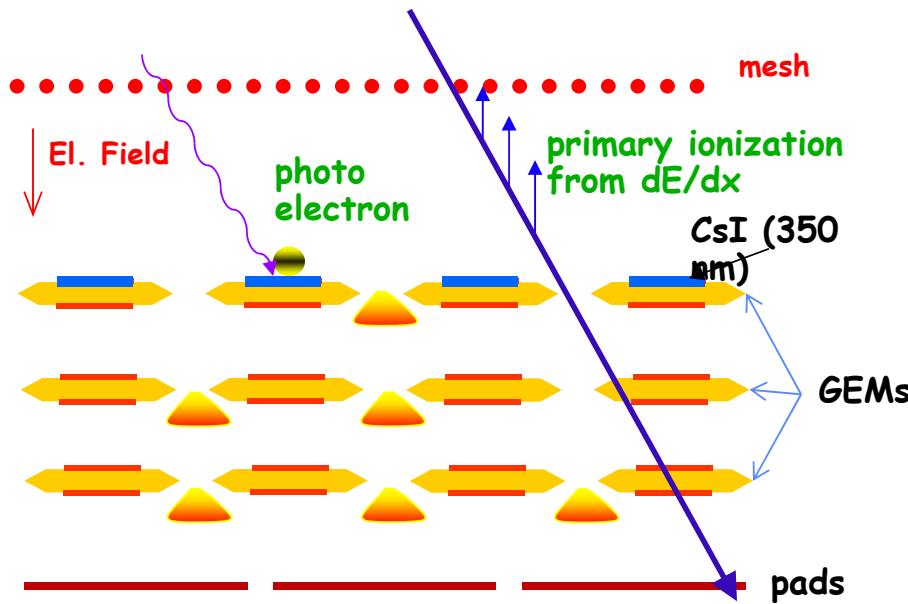
- ✓ Proximity focus configuration, no window, no mirror
- ✓ CF_4 radiator and active gas. $L_{\text{rad}}=50 \text{ cm}$
- ✓ Very large bandwidth 108 – 200 nm (6.2 - 11.5 eV)
- ✓ triple GEMs for signal multiplication
- ✓ CsI photocathode



W. Anderson et al.,
NIM A 646 (2011) 35

HBD - hadron blindness

- ❖ Electron signals are relatively rare (compared to hadrons)

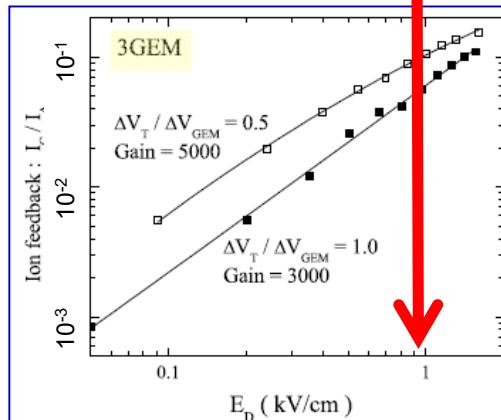
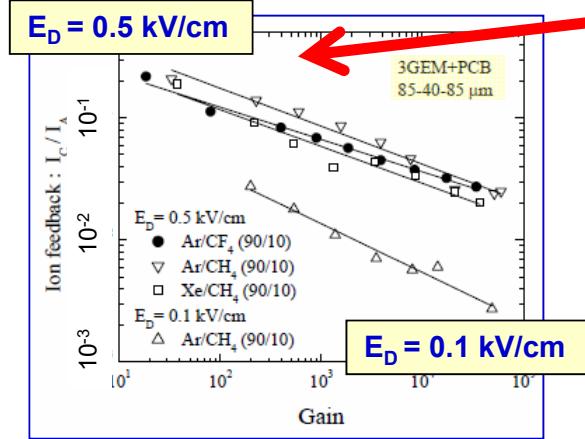


- Detector operated in reverse bias mode to repel the ionization charge from dE/dx
- Cherenkov light is formed only by e^+ or e^-
- Successfull operation at PHENIX since seversl years
- It is not a detector of single photons

GEM-based PDs and IBF

rich literature about IFB in GEM-based detectors
here examples with semi-transparent PC

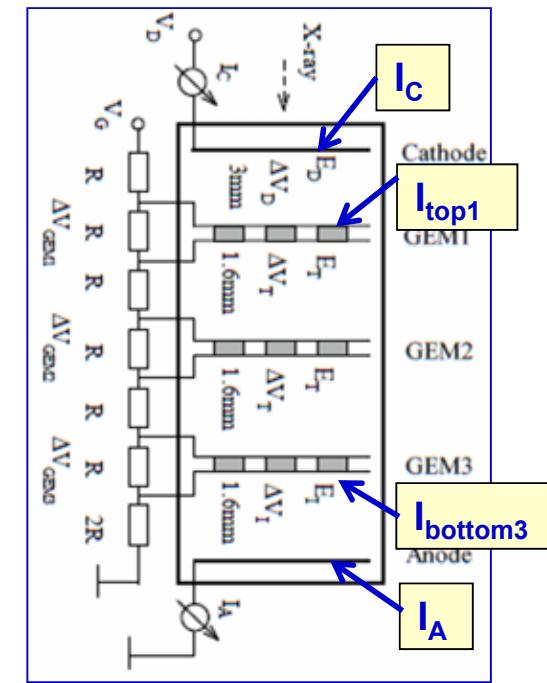
- strong dependence from gain and $E_{D\text{RIFT}}$
- poor dependence from pressure and gas type



A. Bondar et al., NIMA 496 (2003) 325

A. Breskin et al., NIMA 478 (2002) 225d

$E \sim 1 \text{ kV/cm}$ needed
for good photoelectron
extraction



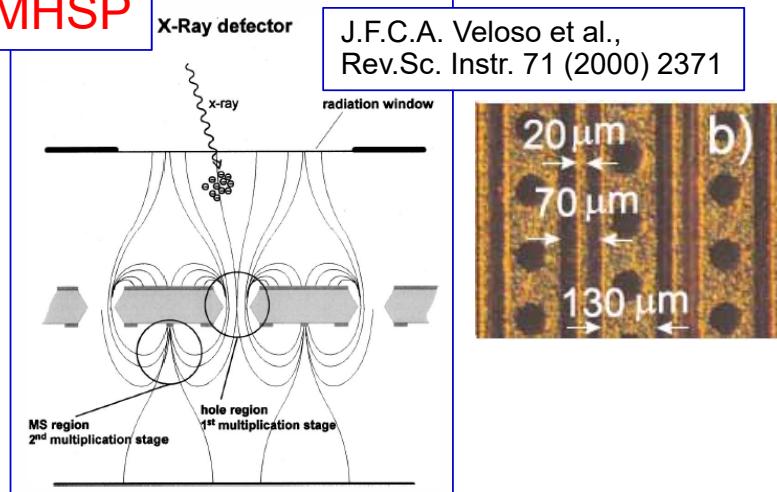
The same for reflective PCs :
small and reversed E_D is needed

IBF: a few % level in
effective GEM-based
photon detectors

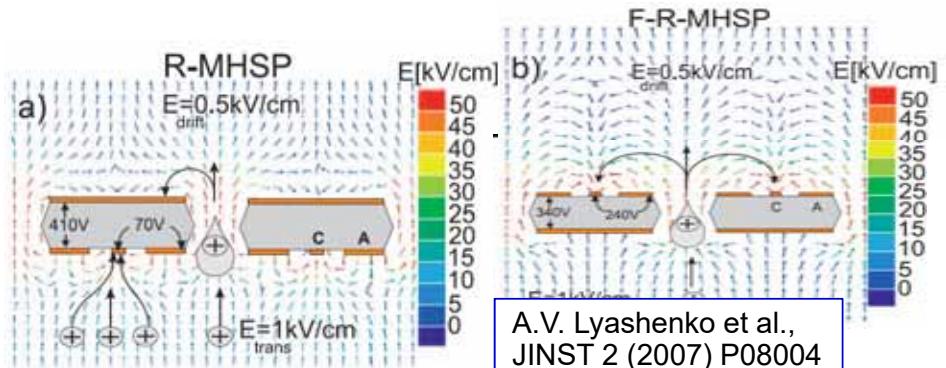
OVERCOMING IBF

More complex geometries needed with extra electrodes to trap the ions:

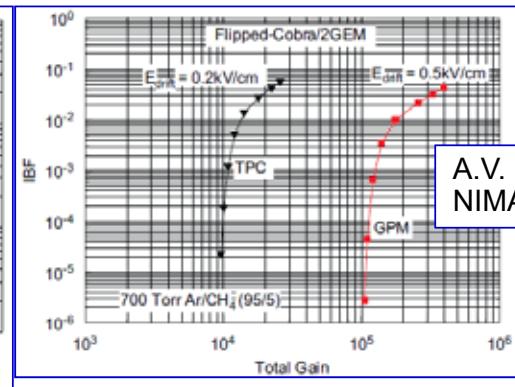
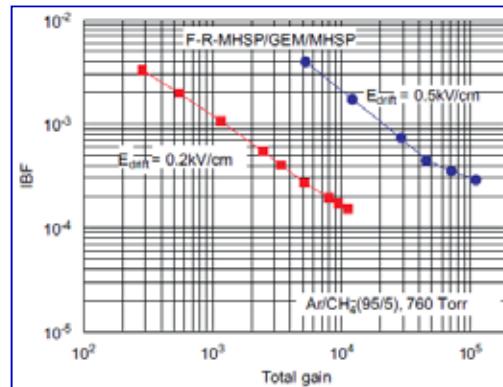
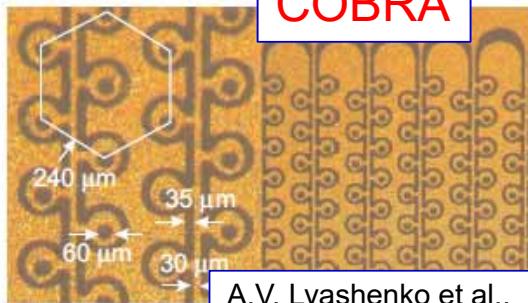
MHSP



Micro-Hole & Strip Plate (MHSP), COBRA



COBRA



A.V. Lyashenko et al., NIMA 598 (2009) 116

GEM-based PDs and GAIN

LARGE GAIN RELEVANT FOR SINGLE PHOTON DETECTION

- **GEM-based PDs in laboratory studies**
 - for single photoelectron detection, they have been operated at gains $> 10^5$ (see, for instance, the plots of the previous slides)
 - **GEM-based detectors in experiments**
 - Always a MIP flux and small rates of heavily ionizing fragments crossing the detectors (even when the detectors are used as photon detectors)
 - At COMPASS: $G \sim 8000$ (B. Ketzer, private comm.)
 - At LHCb: $G \sim 4000$ (M. Alfonsi NIMA 581 (2007) 283)
 - At TOTEM: $G \sim 8000$ (G. Catanesi, private comm.)
 - Phenix HBD: $G \sim 4000$ (W. Anderson et al., NIMA 646 (2011) 35)
- In experiments, small chances
to operate GEM-based PDs at gains $> 10^4$

THGEMs



GAS ELECTRON MULTIPLIER FORMED BY A **RIGID DIELECTRIC FOIL** BETWEEN ELECTRODES, PROVIDED WITH A PATTERN OF HOLES.

In a proper gas and with electric bias it can provide large electron multiplication

Material:

- FR4, permaglass, ...
- PTFE, PET, ARLON, ...
- glass, PEG3 (etchable glass), ...
- ceramic

Holes:

- mechanical drilling (1 € per 1000 holes)
- water jet
- laser
- chemical etching
- preformed (capillary plates)

THGEM-based PDs, why ?

PCB technology, thus:

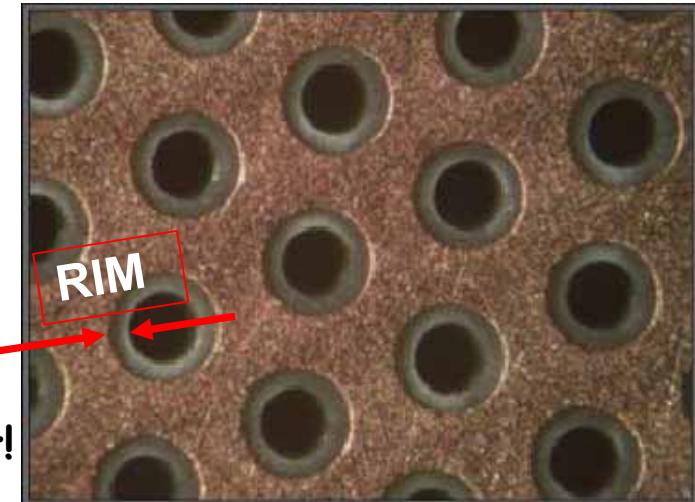
- robust
- mechanically self supporting
- industrial production of large size boards
- large gains have been immediately reported (rim !)

Comparing to GEMs

- Geometrical dimensions $\times \sim 10$
 - But e^- motion/multiplic. properties do not!
 - Larger holes:
 - dipole fields and external fields are strongly coupled
 - e^- dispersion plays a minor role

About PCB geometrical dimensions:

Hole diameter :	0.2 - 1 mm
Pitch :	0.5 - 5 mm
Thickness :	0.2 - 3 mm

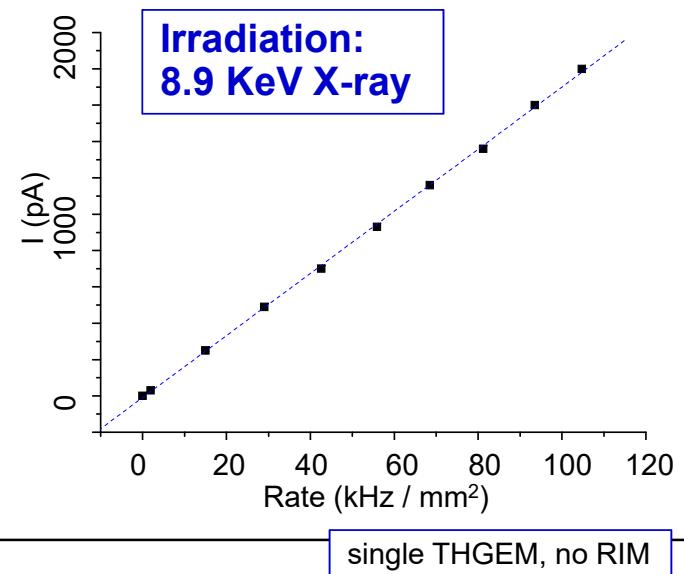


introduced in // by different groups:

- L. Periale et al., NIM A478 (2002) 377.
- P. Jeanneret, PhD thesis, Neuchatel U., 2001.
- P.S. Barbeau et al, IEEE NS50 (2003) 1285
- R. Chechik et al, NIMA 535 (2004) 303

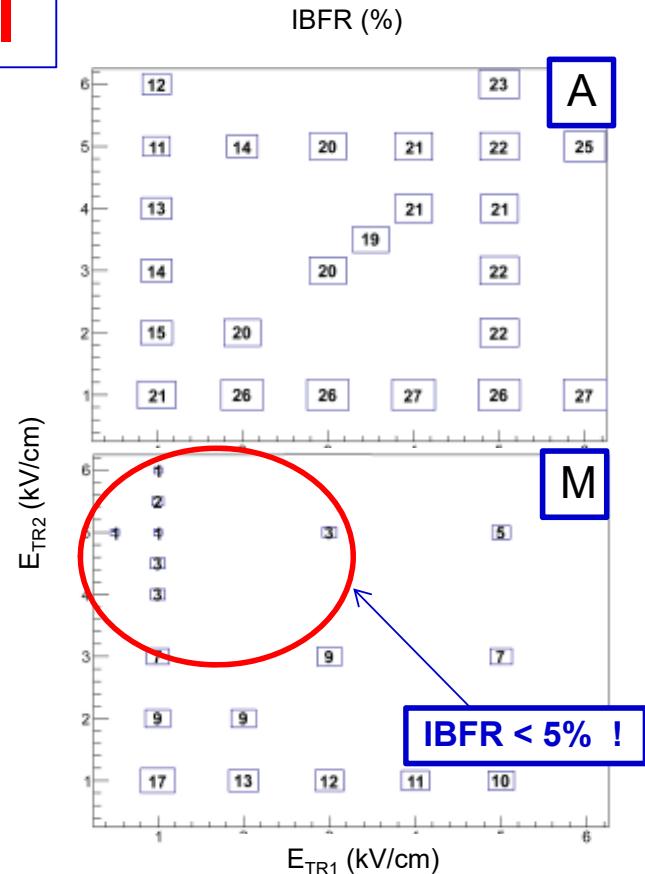
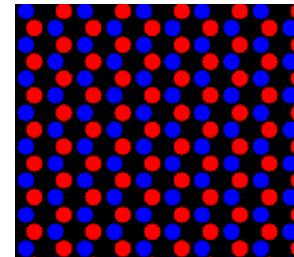
THGEM rate capability and IBF

High rate device



IBF control

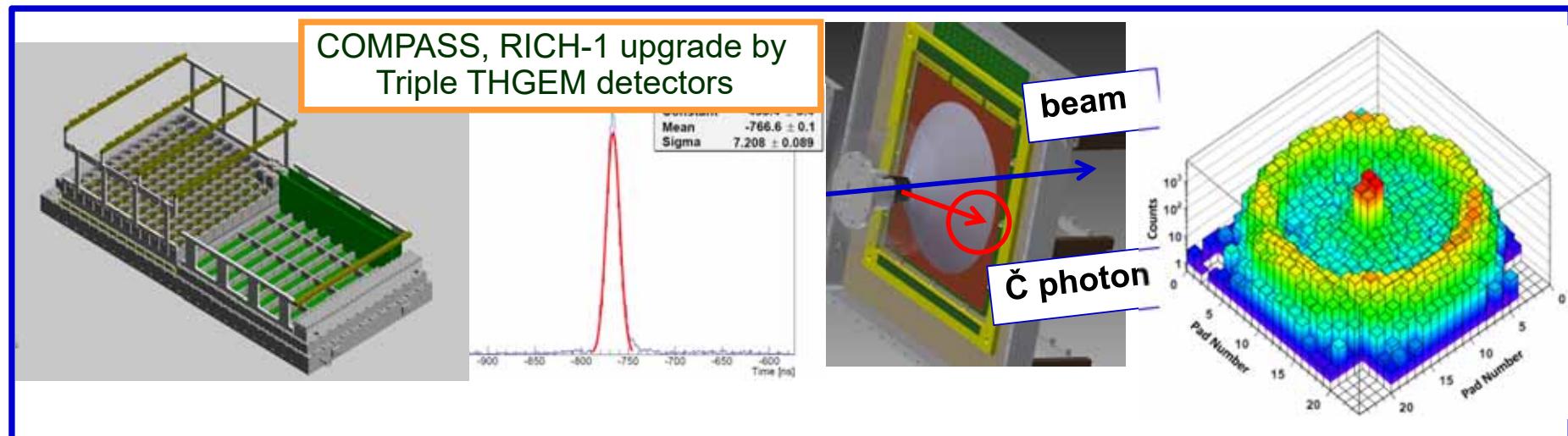
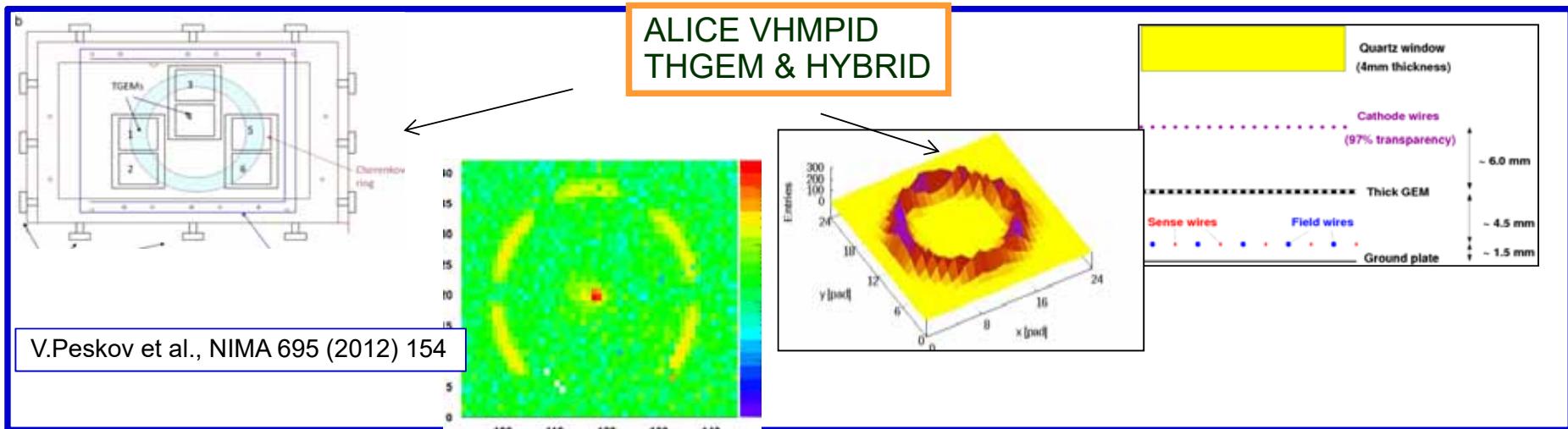
Triple THGEM:
Ion Back Flow
reduction by
staggering plates



M. Alexeev et al. JINST 10 (2015) P03026
The gain in Thick GEM multipliers and its time-evolution

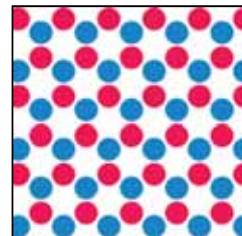
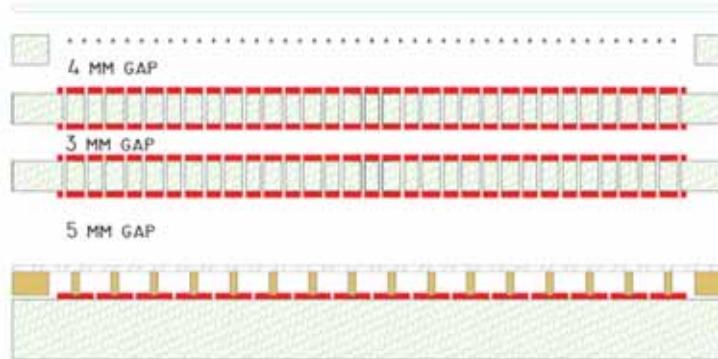
M. Alexeev et al., JINST 7 (2012) C002014

THGEM R&D for RICHes

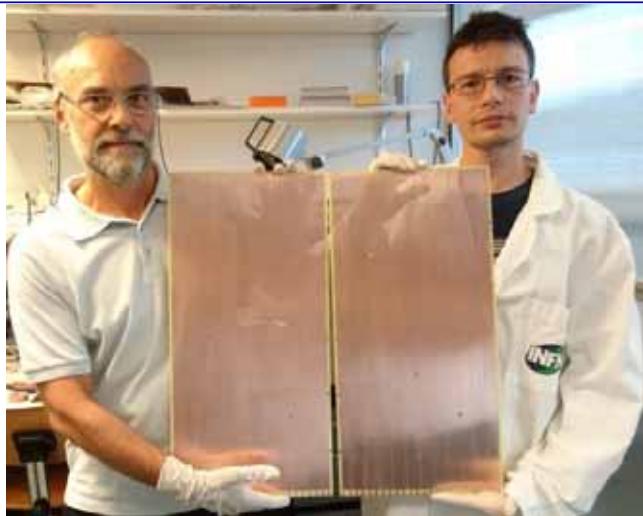


HYBRID MPGD PDs (THGEM + MM)

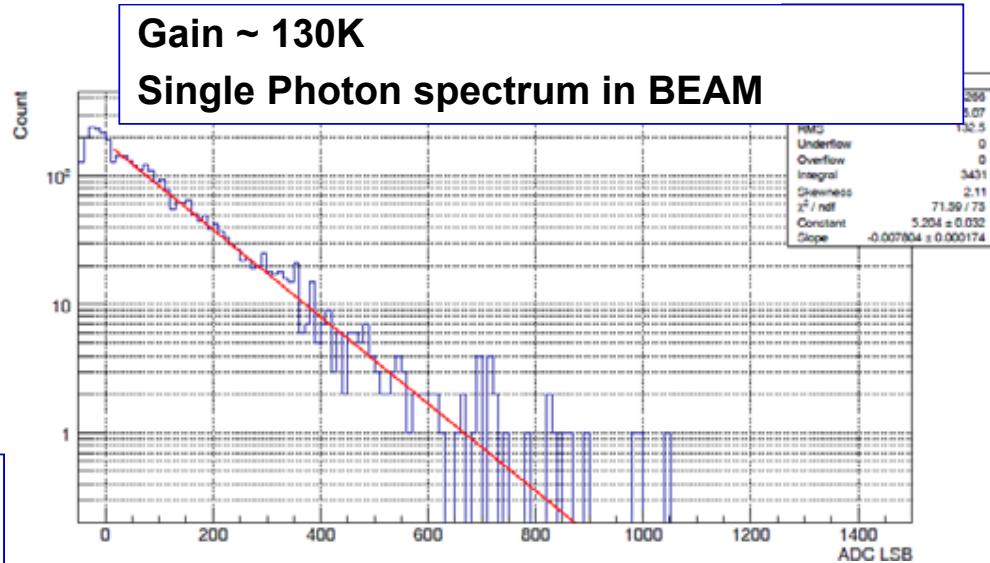
- The 1st THGEM forms the PC
- The 2nd THGEM (staggered) forces the electron diffusion
- The MM provides large gain, made larger by the diffusing the impinging electron cloud



THGEM
staggered !



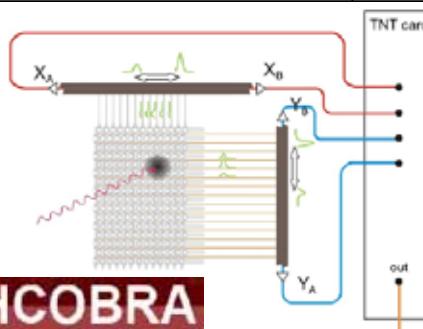
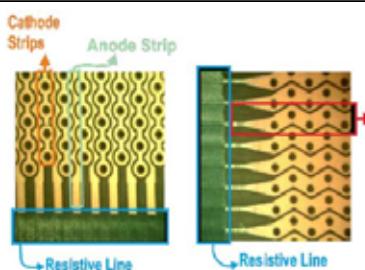
The same architecture is independently studied as GPM for DM searches





HYBRID MPGD PDs (THGEM + THCOBRA)

- 2 THGEMs
- a THCOBRA with 2 d R-O structure



2D-THCOBRA



2D THCOBRA

THGEMs

CsI Photocathode

Parameters			
Structure	Hole Diameter (μm)	Pitch (μm)	RIM (μm)
THGEM 1	400	800	5
THGEM 2	700	1300	100
2D THCOBRA	400	1000	80

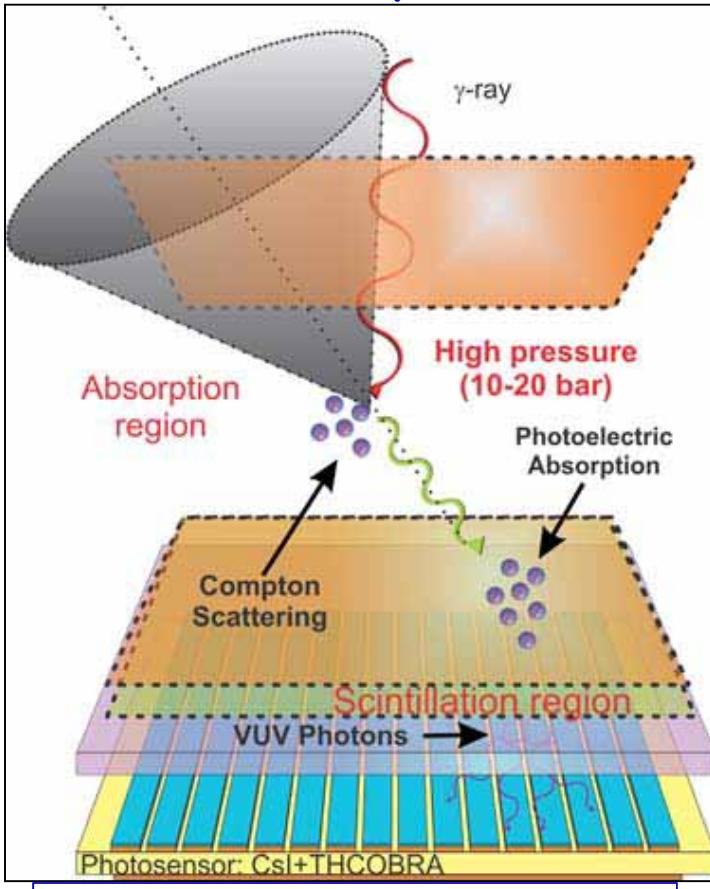
• Gas Photomultiplier (GPM) : 2D-THCOBRA

- Good Performance
 - Gain of 10^6
 - IBF values of about: 20%
- 2D-THCOBRA adequate to obtain image
- Position Resolution: FWHM = 300 μm, $\sigma = 128\mu\text{m}$
- Count rate of 100kHz

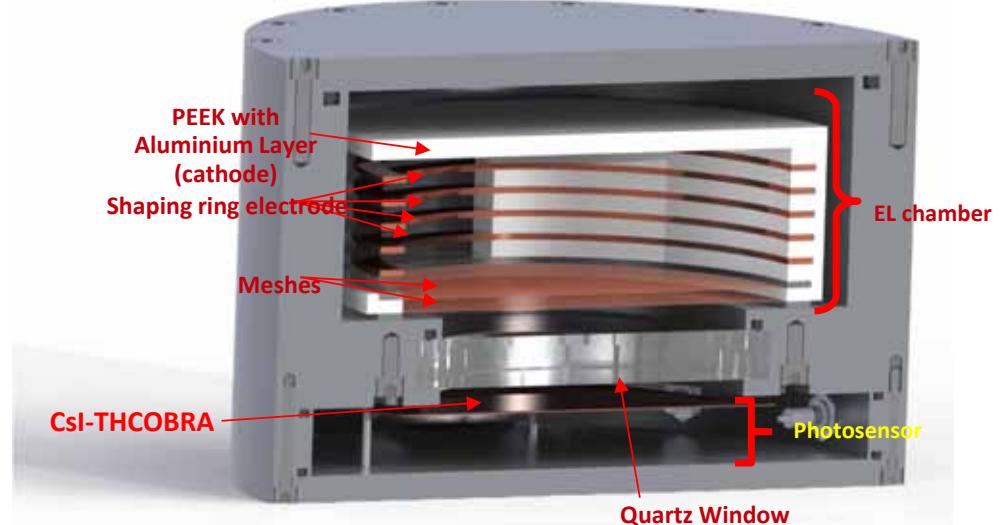
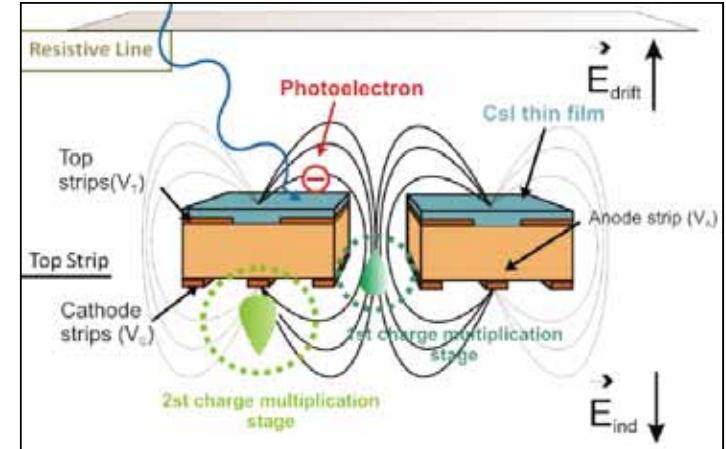
T. Lopes 2013 JINST 8 P09002

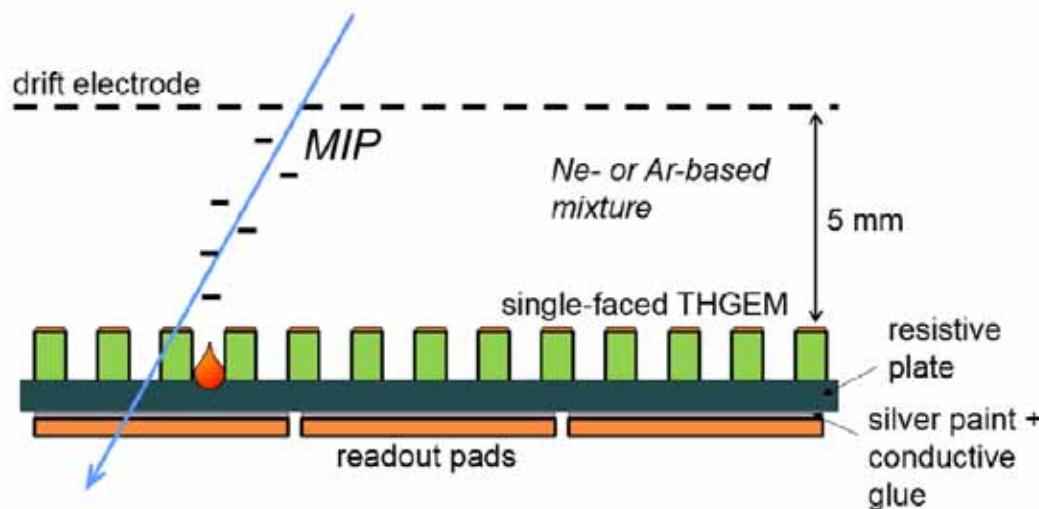
Gaseous Compton camera

Electroluminescence light is detected by THCOBRA with 2D R-O
Drift time provides the third coordinate



C. D. R. Azevedo et al., NIM A 732 (2013) 551

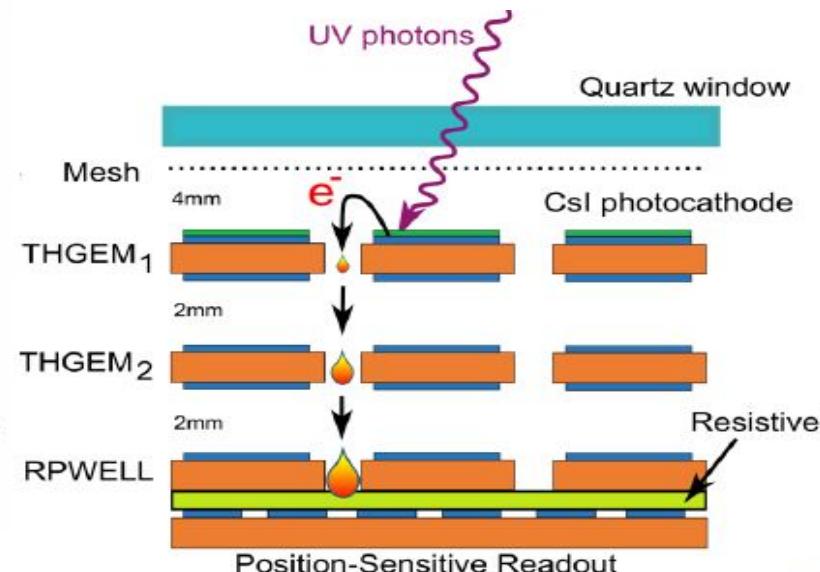




$10^9 \Omega \text{ cm}$ resistive plate \rightarrow discharge free operations. 99% eff. up to $\sim 10^5 \text{ Hz/cm}^2$

proposed for digital hadron calorimetry

L. Moleri et al., NIMA 845 (2017) 262



proposed for UV photon detection

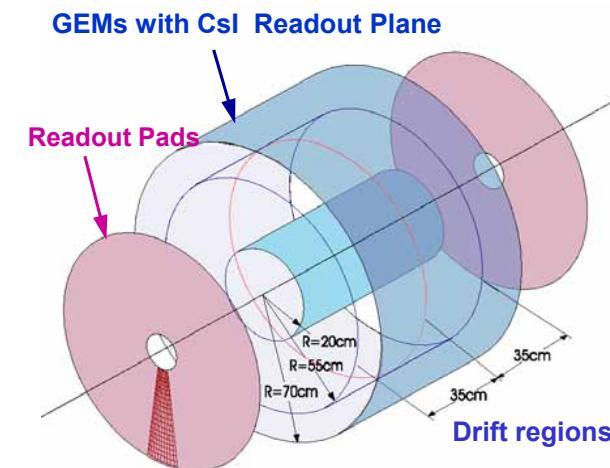
S. Bressler et al. JINST July 2013

arXiv:1305.4657

L. Arazi et al. 2012 JINST 7 C05011 arXiv:1112.1915

MPGD-based PDs for a detector at EIC

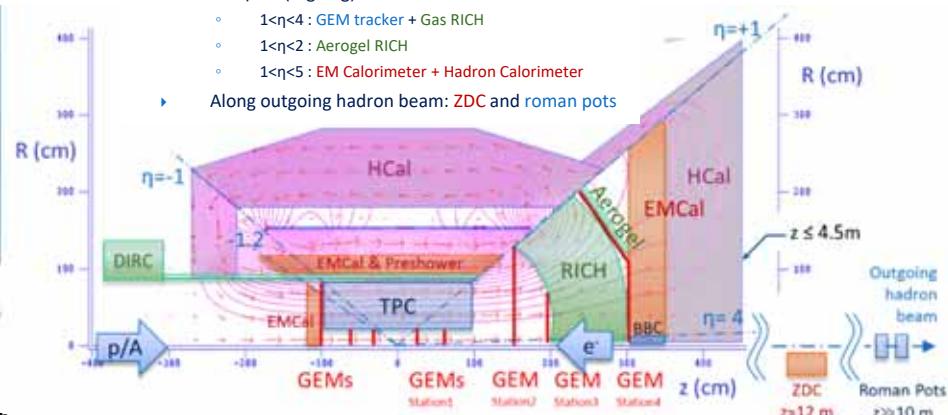
Current PHENIX	sPHENIX (+fsPHENIX)	An EIC detector
<ul style="list-style-type: none"> 15 years of operation Produced broad spectrum of studies of QGP and Hadron Physics 140+ published papers to date Last run in 2016 <p>~2000 2017→2020 ~2025 Time</p> <p>RHIC: A+A, spin-polarized p+p, spin-polarized p+A</p>	<ul style="list-style-type: none"> Comprehensive upgrade based on the BaBar magnet Goal is to do a systematic study the QGP near T_c by measuring jets and heavy quarkonia Possible addition of a forward spectrometer (fsPHENIX) to continue study of Spin and CNM 	<ul style="list-style-type: none"> Further upgrades of sPHENIX leads to a capable Day 1 detector for eRHIC Study polarized ep and eA physics Large coverage of tracking, calorimetry and PID



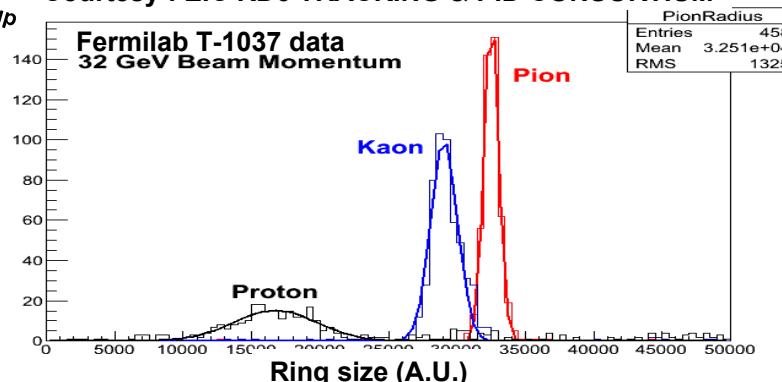
M.Blatnik et al. arXiv:1501.03530

sPHENIX → eRHIC Detector LOI: arXiv:1402.1209

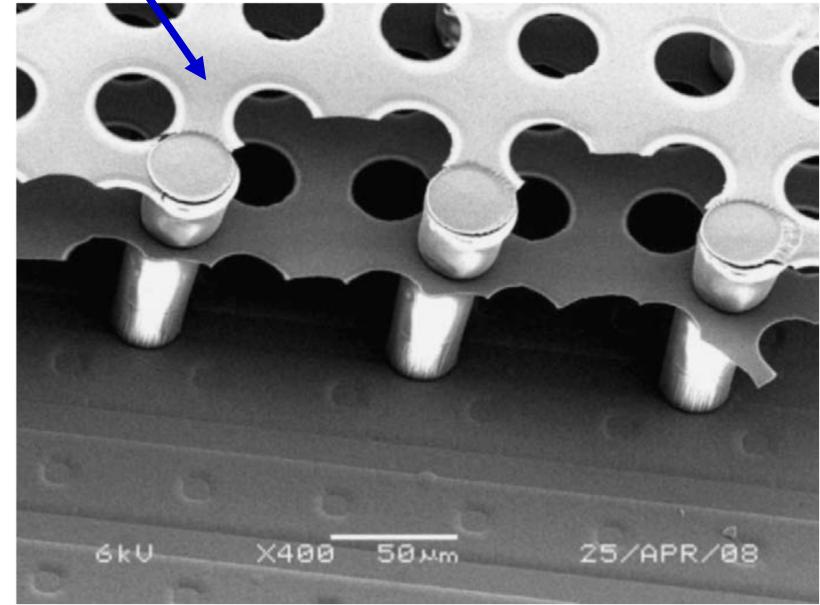
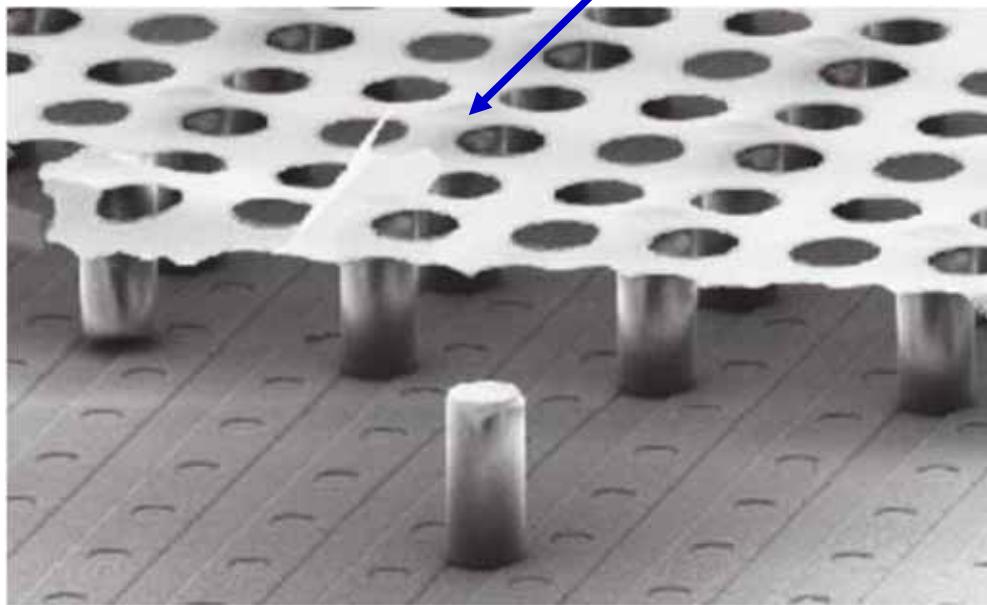
- 1 < η < +1 (barrel) : sPHENIX + Compact-TPC + DIRC
- 4 < η < -1 (e-going) : High resolution EM calorimeter + GEM trackers
- +1 < η < +4 (h-going) :
 - 1 < η < 4 : GEM tracker + Gas RICH
 - 1 < η < 2 : Aerogel RICH
 - 1 < η < 5 : EM Calorimeter + Hadron Calorimeter
- Along outgoing hadron beam: ZDC and roman pots



Courtesy : EIC RD6 TRACKING & PID CONSORTIUM



InGrid and TwinGrid



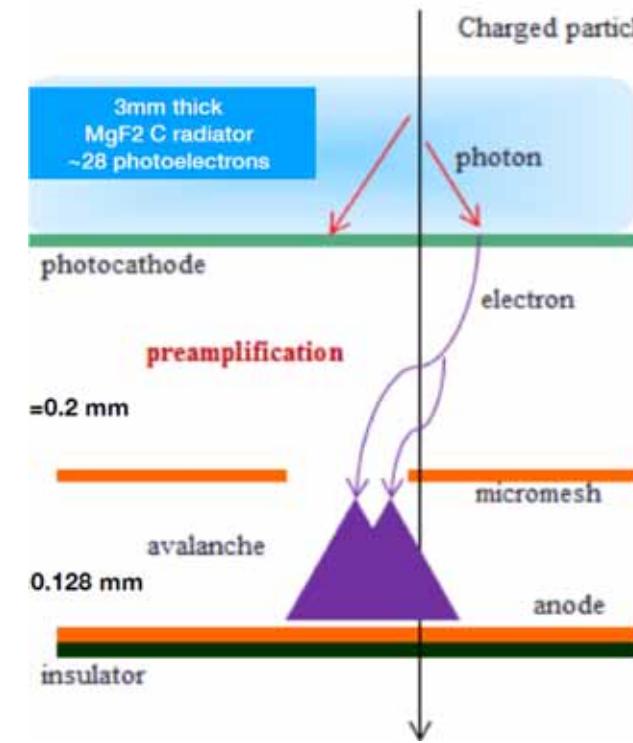
M.Chefdeville et al., Nucl. Instrum. Meth. A 556 (2006) 490.

Y.Bilevych et al., Nucl. Instrum. Meth. A 610 (2009) 644.

Excellent space resolution for single UV photons provided by InGrid with CsI coating of a micro-grid directly integrated by wafer post-processing production onto a CMOS pixel detector with the complete readout system.

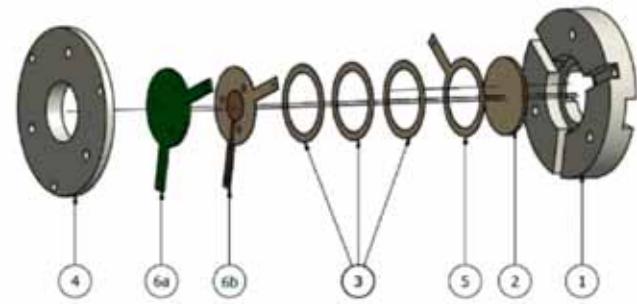
The array of microgrid round holes corresponds to the array of CMOS pixel centers

“picosecond” timing with MPGDs?

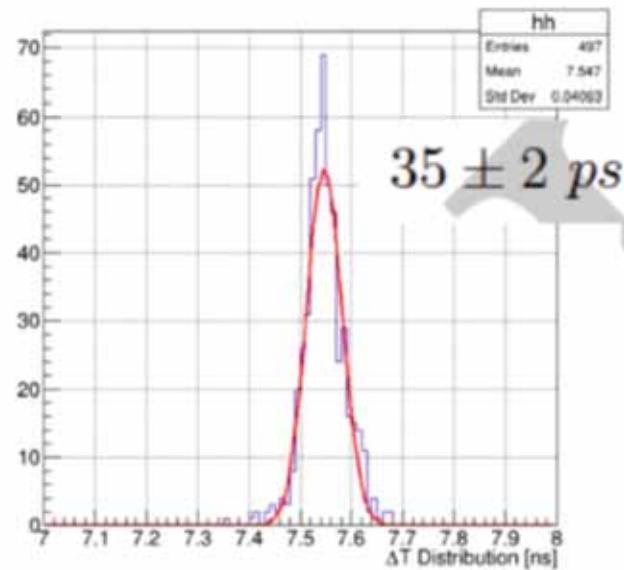


T.Papaevangelou et al, Fast Timing for High-Rate Environments with Micromegas, arXiv:1601.00123 (Jan. 2016).

Talk of Sebastian White



$t_{\text{PICOSEC}} - t_{\text{MCP}}$



Many configuration tested

Successfully achieved ~35 ps resolution in test beam.

Puzzling shift in spite of CF.

A NEW PHOTOCONVERTER: Rich-graphite Nano-Diamond film

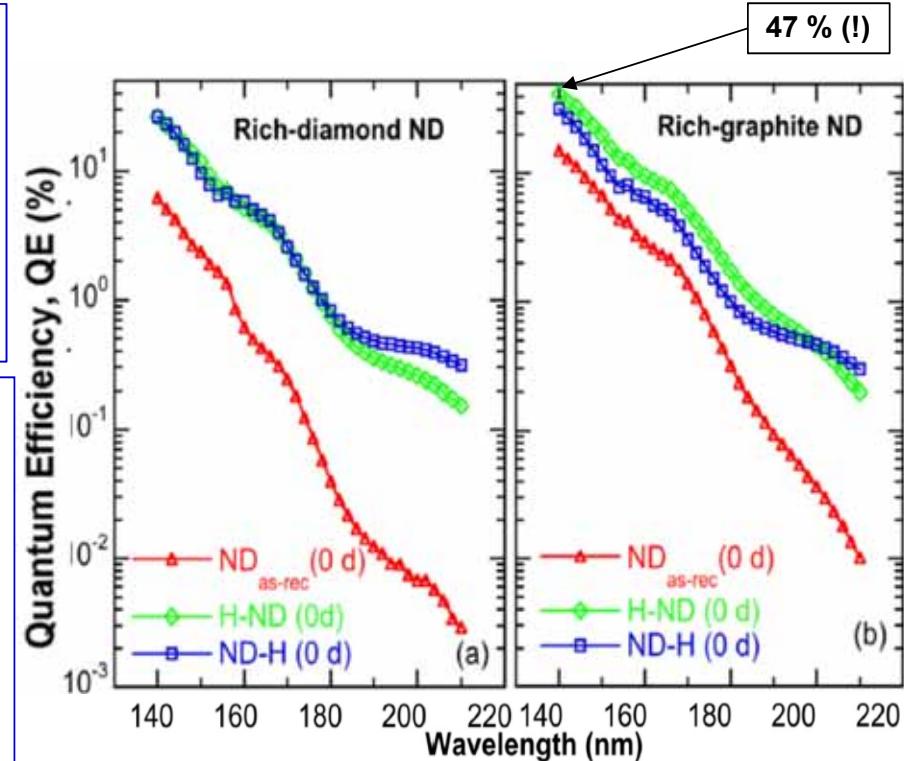
CsI bandgap: 6.2 eV; electron affinity: 0.1 eV;
hygroscopic; ages by ion bombardment ($\sim \text{mC/cm}^2$)

Diamond bandgap: 5.5 eV; chemically inert and robust; if hydrogenated: electron affinity -1.27 eV

Hydrogenated chemical vapor deposited diamond films (4-6 μm) known to have QE $\sim 15\%$ @ 140 nm.

Heterogranular-structured diamond-gold nanohybrids proposed as stable field emission cathode material

Nano-Diamond grains (size: ~ 250 nm), with variable sp^2 (graphite phase) and sp^3 (diamond phase) hybridized carbon contents treated in H₂ microwave plasma show large QE: $\sim 50\% @ 140$ nm



L.Velardi, A.Valentini, G.Cicala, Diamond & Related Materials 76 (2017) 1

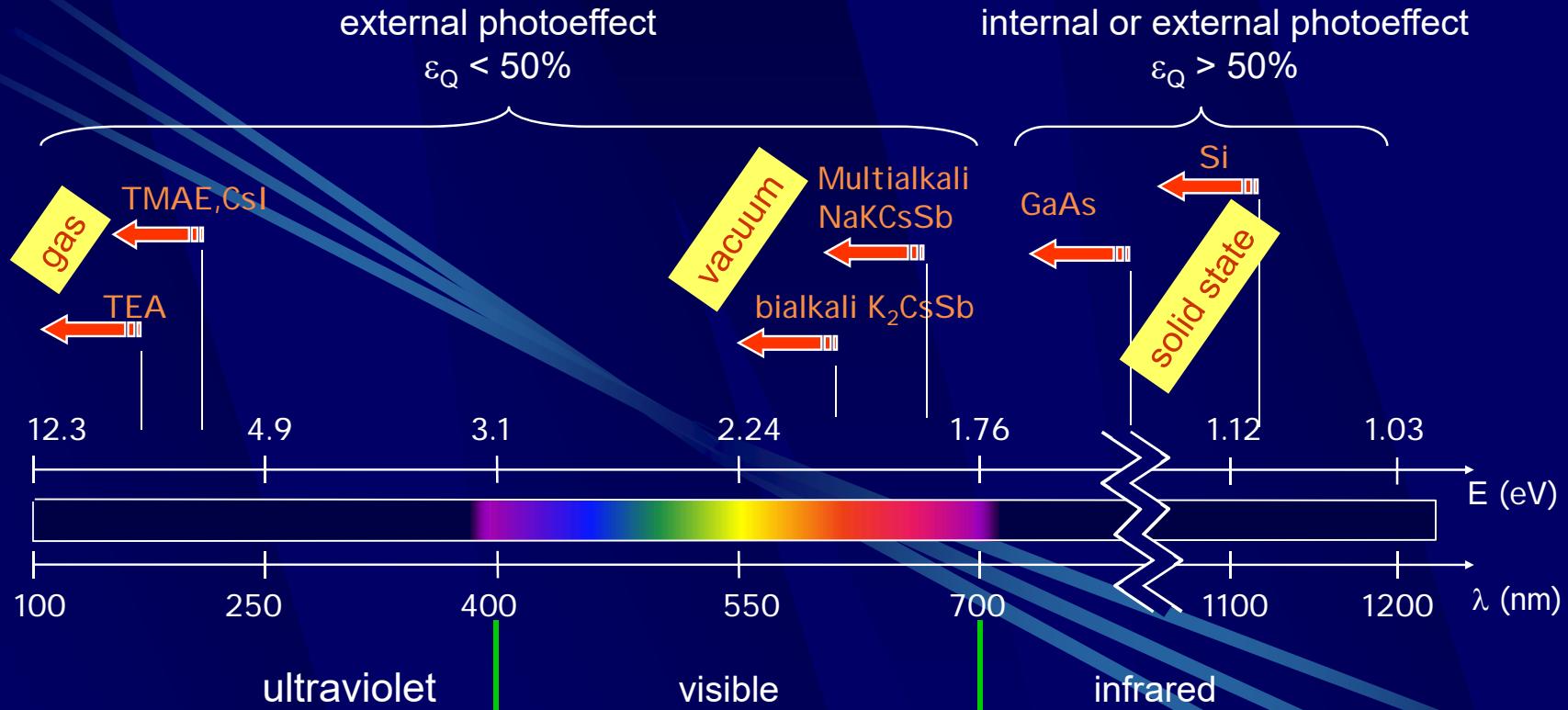
NEW !!!

Photocatodes: diamond film obtained with Spray Technique

Spray technique: $T \sim 120^\circ$ (instead of $\sim 800^\circ$ as in standard techniques)

Different photocathodes and their thresholds

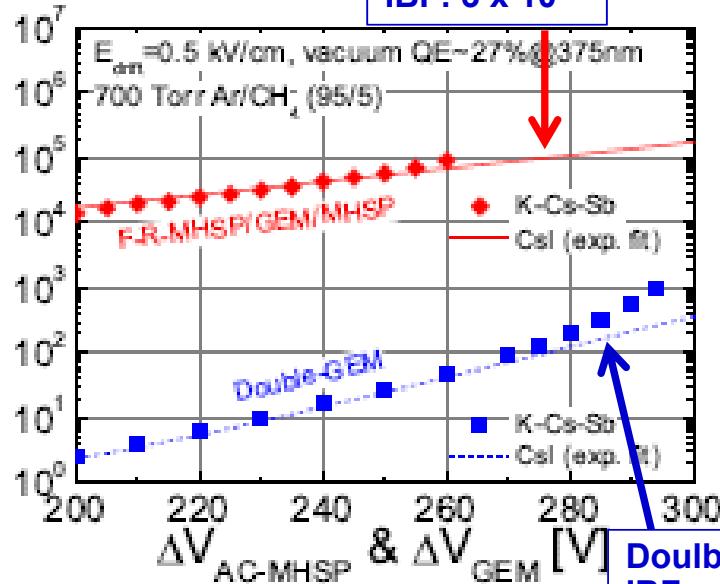
$$\varepsilon_Q = \text{Q.E.} = N_e / N_\gamma$$



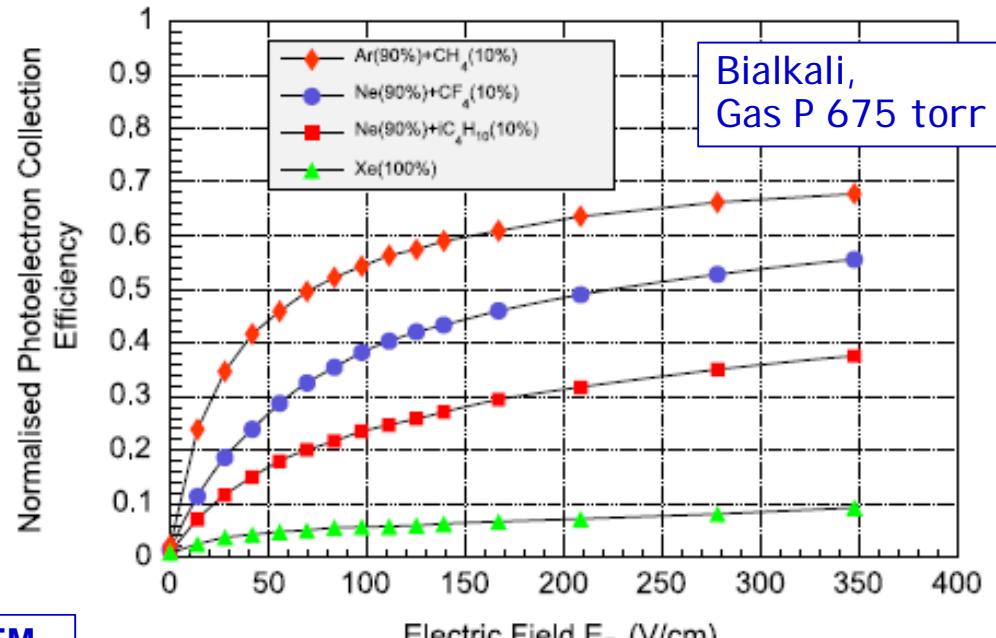
- Photon detection involves often materials like K, Na, Rb, Cs (alkali metals). They have the smallest electronegativity → highest tendency to release electrons.
- Most photocathodes are VERY reactive; Exceptions: Si and CsI.

- Chemical reactivity (gas purity better than ppm level needed → UHV materials and sealed detectors)
 - PC stability under ion bombardment - work function lower than CsI one
 - AGEING CsI: -16% QE at $25\mu\text{C}/\text{mm}^2$
 Bilkali: -20% QE at $0.4\mu\text{C}/\text{mm}^2$
- F.Tokanai et al., NIMA 628 (2011) 190
 T.Moriya et al., NIMA 732 (2013) 263

K-Cs-Sb vs CsI



A.V.Lyashenko et al., 2009 JINST 4 P07005

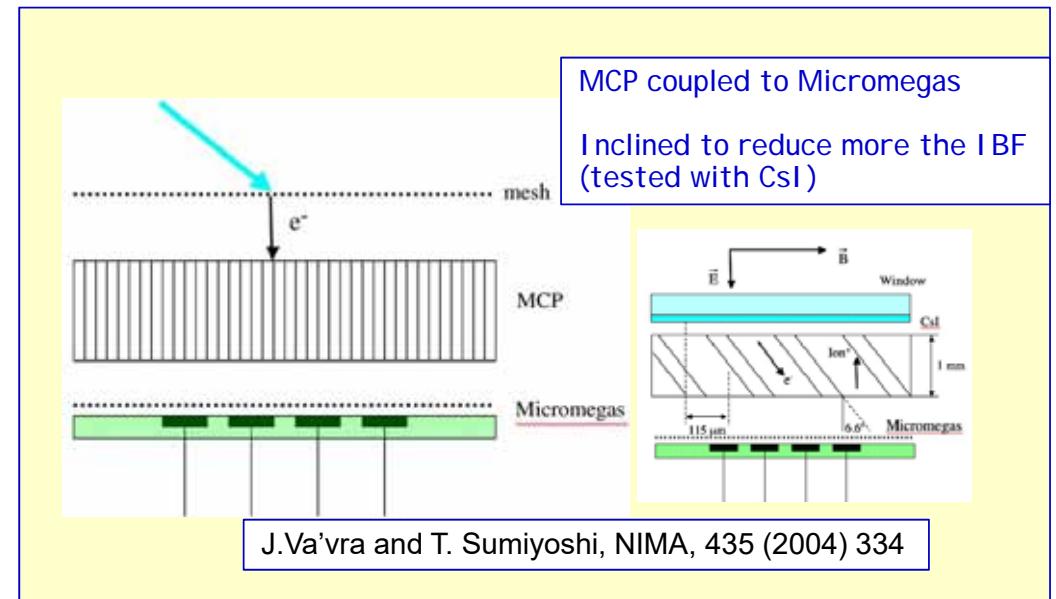
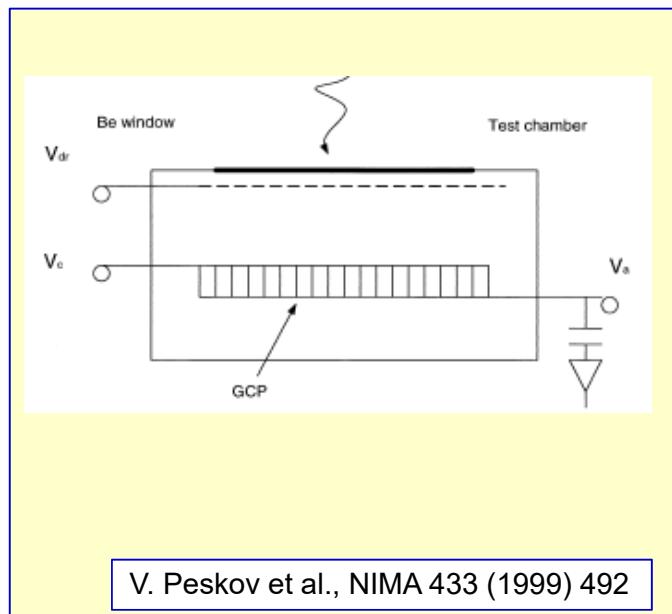


F. Tokanai et al., NIMA 610 (2009) 164



FIRST GASEOUS DETECTORS FOR VISIBLE LIGHT

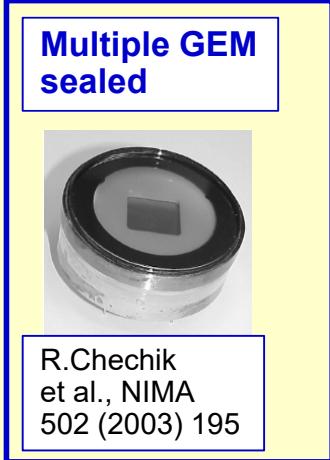
The Capillary Plate (CP) approach



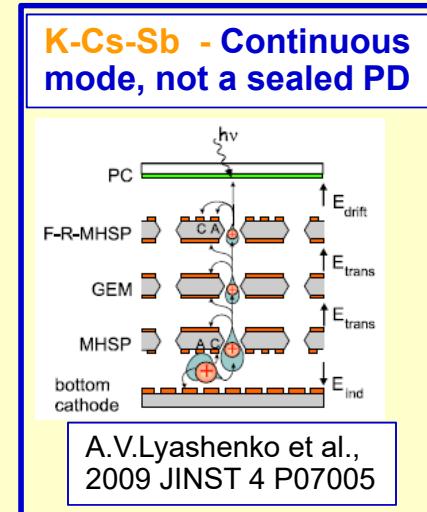
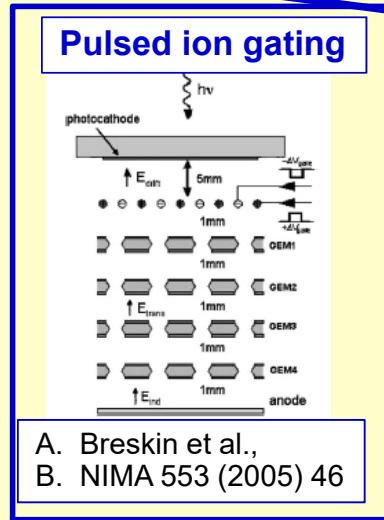
GASEOUS DETECTORS FOR VISIBLE LIGHT



the GEM approach



R.Chechik
et al., NIMA
502 (2003) 195



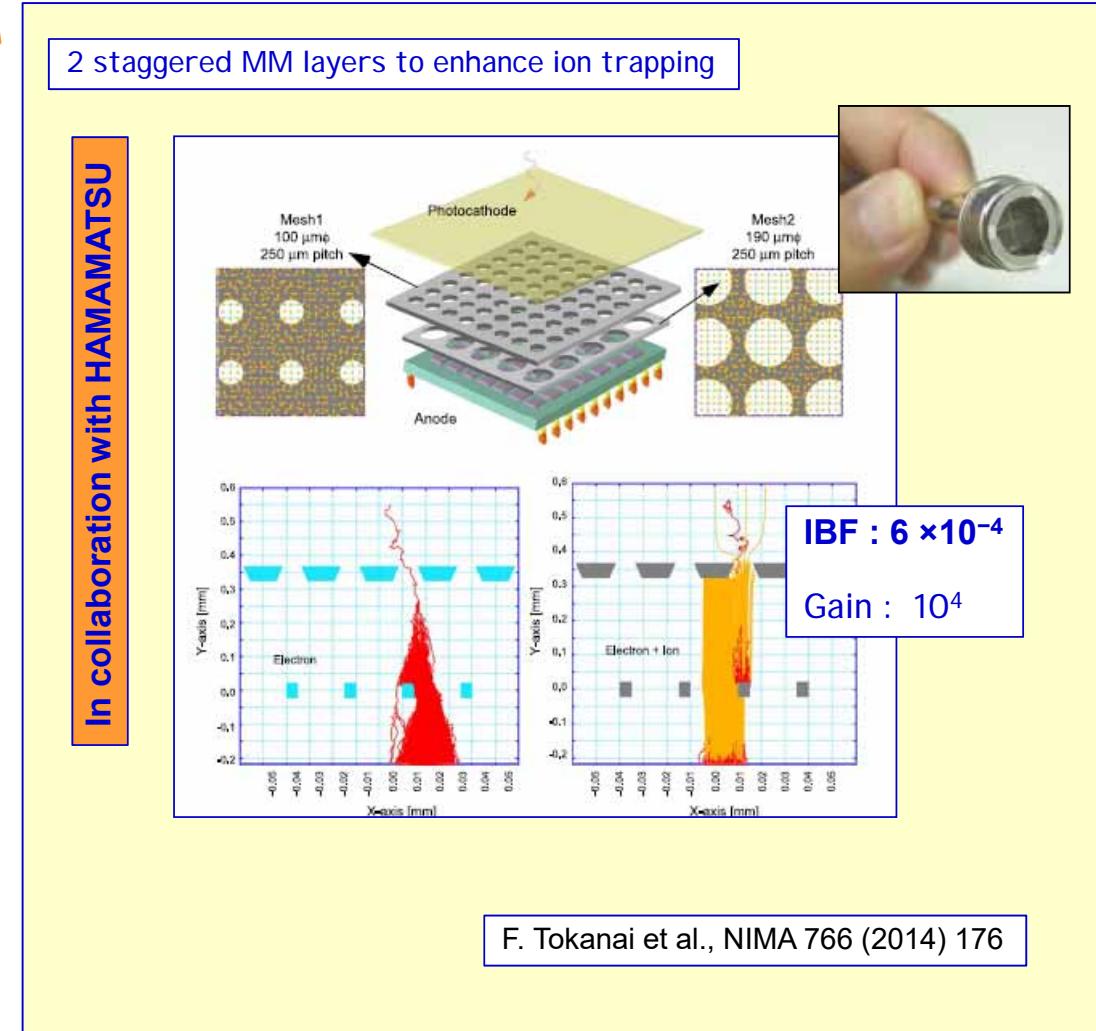
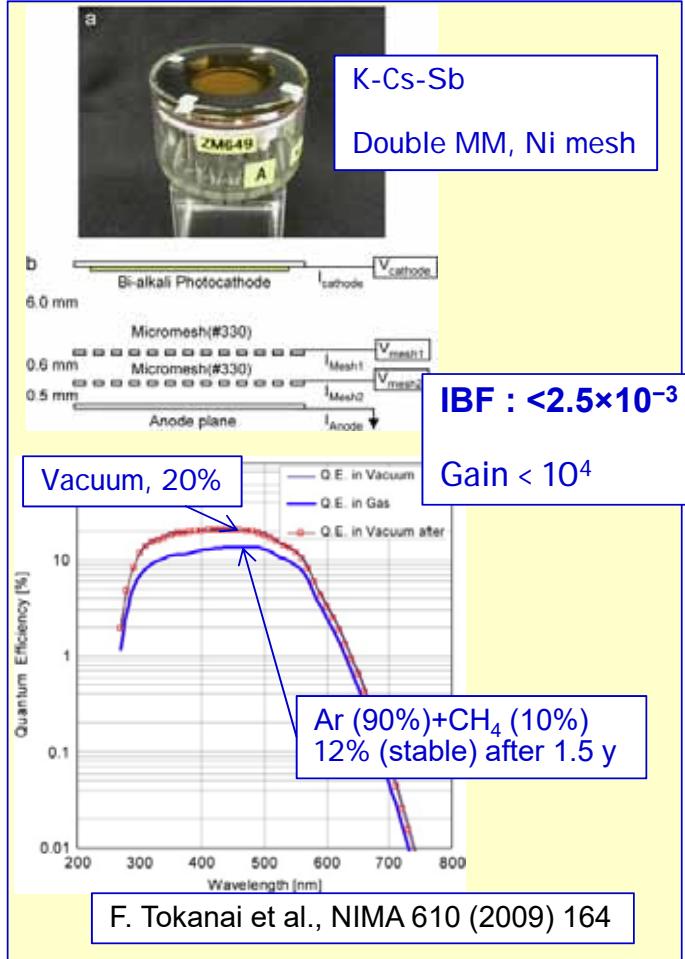
Poor compatibility of bialkali and GEM material ?

Extremely poor QE of the bialkali PC:
the material of the GEM chemically reacts with the bialkali metals

F. Tokanai et al., NIMA 610 (2009) 164

GASEOUS DETECTORS FOR VISIBLE LIGHT

the MicroMegas approach

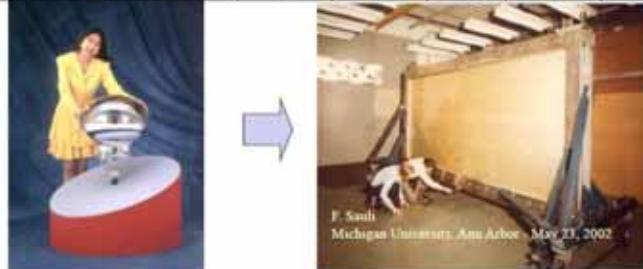


GASEOUS DETECTORS FOR VISIBLE LIGHT

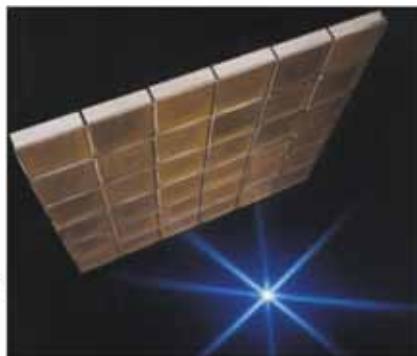
Gaseous PMT

Yamagata U. TMU, HAMAMATSU

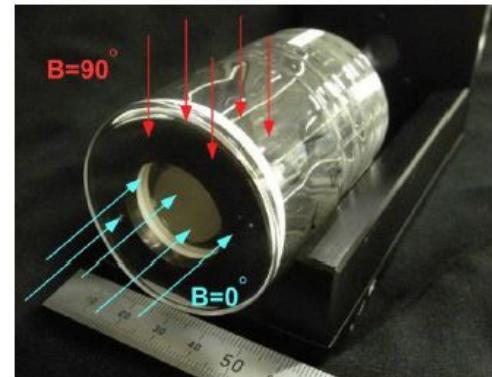
Sensor type	Sensitivity	Position Resolution	Timing Resolution	Uniformity	Price	Magnetic Field	Effective Area
Vacuum PMT	⊗	△	⊗	△	○	△	○
CCD / CMOS	△	⊗	X	⊗	△	⊗	X
Gaseous PMT	○	○	○	○	●	●	●



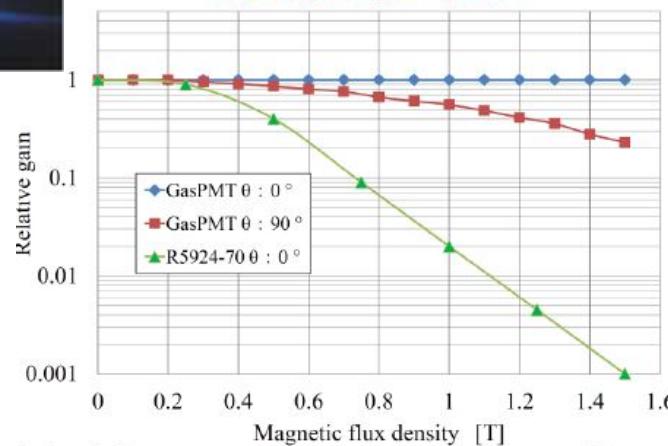
- The advantage of the **gaseous PMT**:
 - It can achieve a **very large effective area** with moderate **position** and **timing** resolutions.
 - It can be easily operated under a **very high magnetic field**.



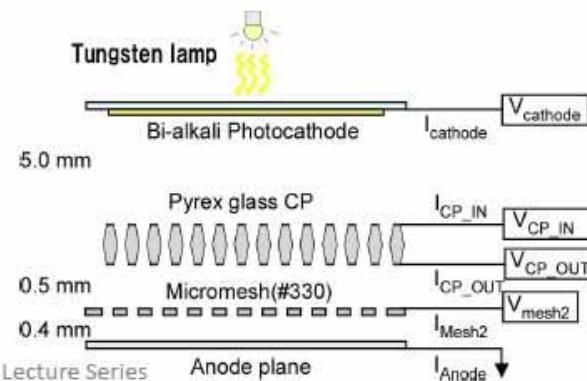
Operation in magnetic field environment



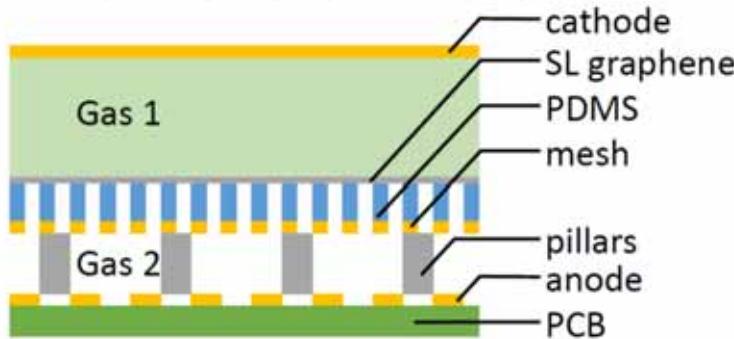
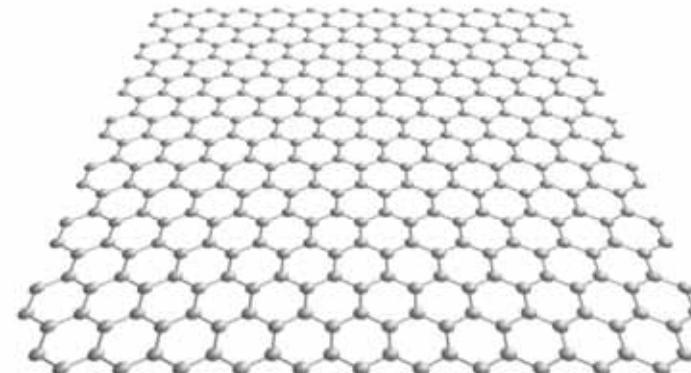
Ar(90%)+CH₄(10%) 1気圧



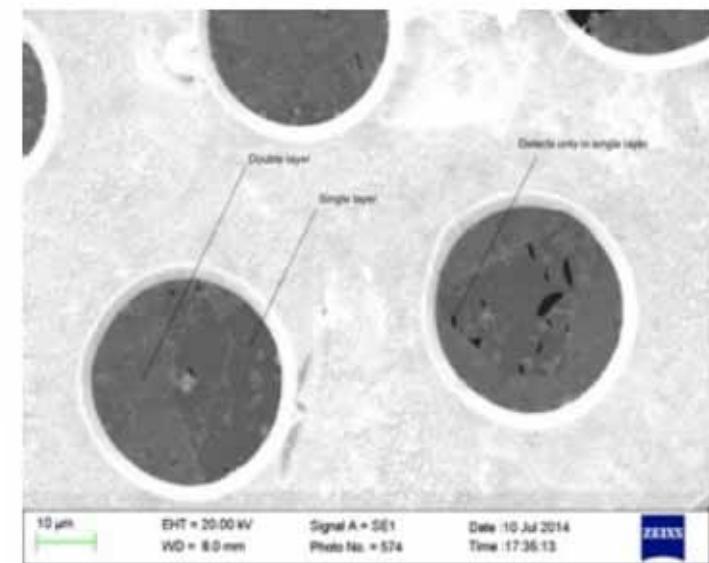
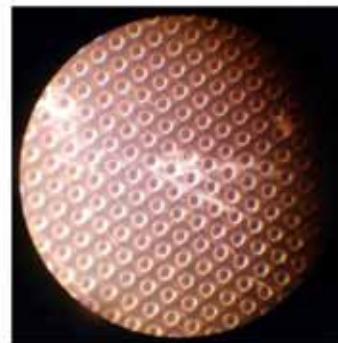
H. Sugiyama et al., NIMA (2016) in press



GRAPHENE



- Single layer of C atoms in hexagonal lattice
- Thinnest possible conductive mesh with 0.6 Å pores
- Very good ion blocking
- Charge transfer properties through graphene layers have been measured



P. Thuiner^{1,2}, R. Hall-Wilton³, R. B. Jackman⁴, H. Müller¹,
 T. T. Nguyen⁴, R. de Oliveira¹, E. Oliveri¹, D. Pfeiffer^{1,3}, F. Resnati¹,
 L. Ropelowski¹, J. A. Smith⁴, M. van Stenis¹, R. Veenhof⁵

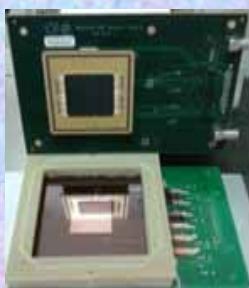
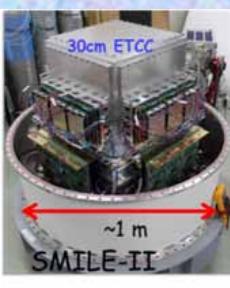
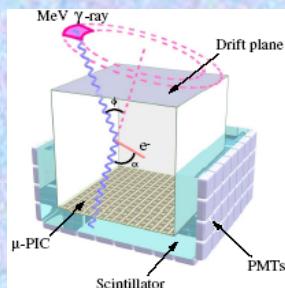
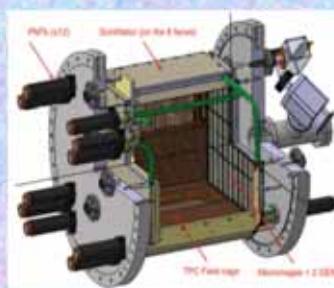
¹CERN, ²Technische Universität Wien, ³ESS,
⁴University College London, ⁵Uludağ University

Promising new idea

MPGD Technologies for X-Ray Detection and γ -Ray Polarimetry

M. Titov, MPGD 2017

Experiment/ Timescale	Application Domain	MPGD Technology	Total detector size / Single module size	Operation characteristics / Performance	Special Requirements/ Remarks
KSTAR @ Korea Start: 2013	X-ray Plasma Monitor for Tokamak	GEM GEMPIX	Total area: 100 cm ²	Spat. res.: ~ 8x8 mm ² 2 ms frames; 500 frames/sec	
			Total area: 10-20 cm ²	Spat. res.: ~ 50x50 μm^2 1 ms frames; 5 frames/sec	
PRAxyS Future Satellite Mission (US-Japan): Start 2020 - for 2years	Astrophysics (X-ray polarimeter for relativistic astrophysical X-rays)	TPC w/ GEM	Total area: 400 cm ³ Single unit detect. (8 x 50cm ³) ~400cm ³	Max.rate: ~ 1 lcps Spatial res.: ~ 100 μm Time res.: ~ few ns Rad. Hard.: 1000 krad	Reliability for space mission under severe thermal and vibration conditions
HARPO Balloon start >2017?	Astroparticle physics Gamma-ray polarimetry (Tracking/Triggering)	Micromegas + GEM	Total area: 30x30cm ² (1 cubic TPC module) Future: 4x4x4 = 64 HARPO size mod.	Max.rate: ~ 20 kHz Spatial res.: < 500 μm Time res.: ~ 30 ns samp.	AGET development for balloon & self triggered
SMILE-II: Run: 2013-now	Astro Physics (Gamma-ray imaging)	GEM+ μ PIC (TPC+ Scintillators)	Total area: 30 x 30 x 30 cm ³	Point Spread Function for gamma-ray: 1°	
ETCC camera Run: 2012-2014	Environmental gamma-ray monitoring (Gamma-ray imaging)	GEM+ μ PIC (TPC+ Scintillators)	Total area: 10x10x10 cm ³	Point Spread Function for gamma-ray: 1°	



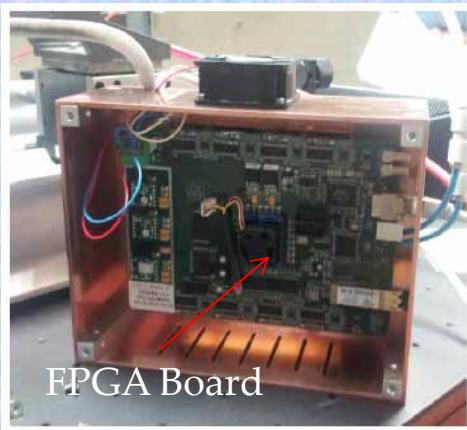
GEM (X-Ray) Detector for Tokamak Plasma Diagnostics

M. Titov, MPGD 2017

10x10 cm² GEM installed since 2013

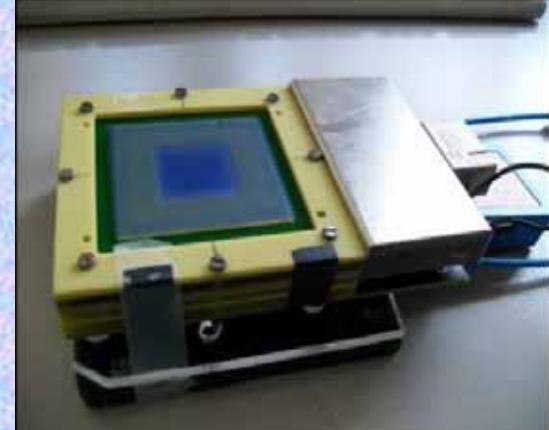


KSTAR
Tokamak



FPGA Board

GEMPIX (GEM + Timepix)



Plasma images (GEM) measured in 2015:
Movie of 200 images per sec

GEMPIX for Fusion:
2015 measurement campaign:

F. Murtas

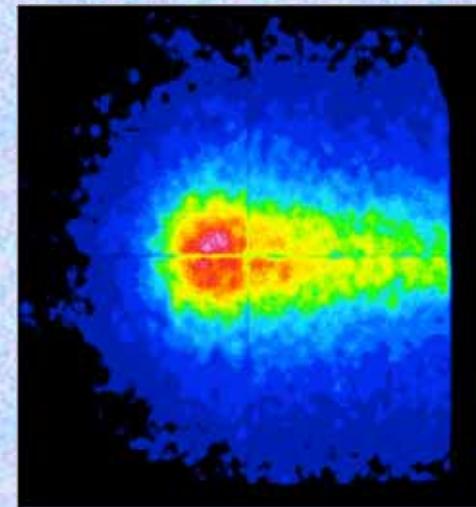
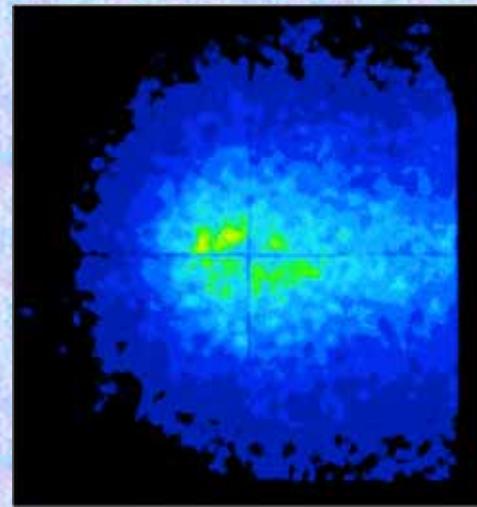
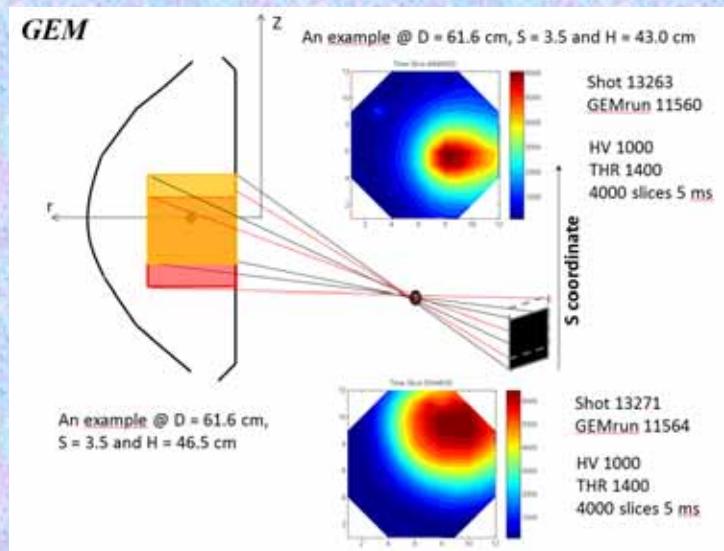


Image of KSTAR Plasma with spectroscopy measurements

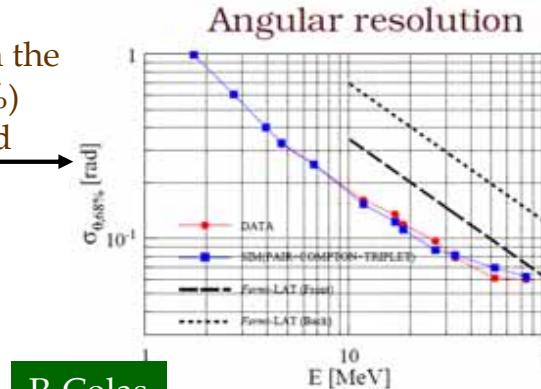
MPGDs Technologies for MeV-GeV Polarimeter and γ -Ray Telescope

M. Titov, MPGD 2017

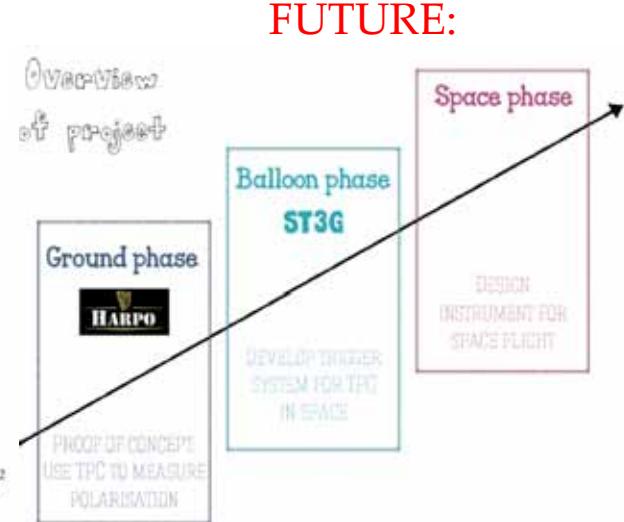
HARPO: TPC as a γ -ray Telescope and Polarimeter:

High-Pressure HARPO GEM+MM TPC (2014):

- Sensitivity to polarization in the 1-100 MeV range ($A \sim 10-15\%$)
- Angular resolution improved wrt FERMI-LAT

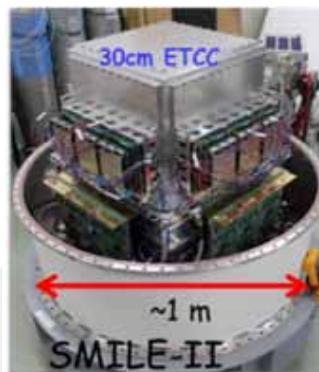
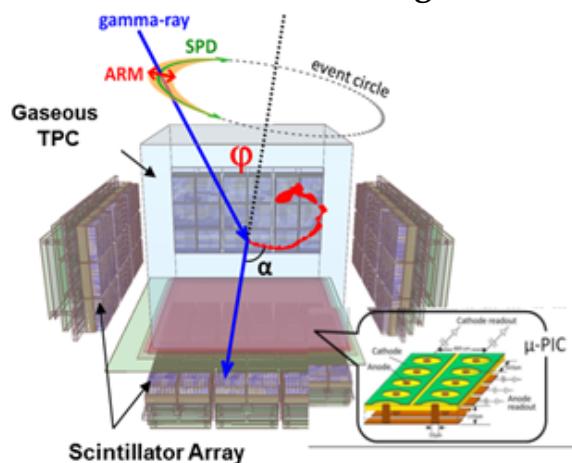


P. Colas



Electron Tracking Compton Camera: μ PIC+TPC:

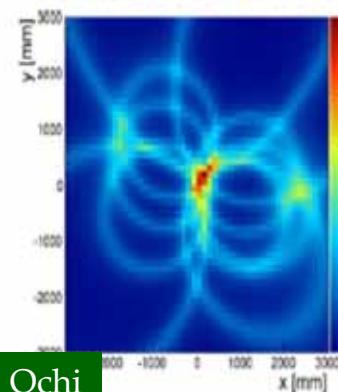
→ tracking of recoil electron



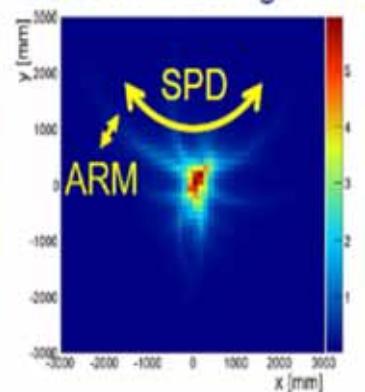
SMILE-II:

Both γ -ray imaging and spectroscopy are available → color image

Conventional method



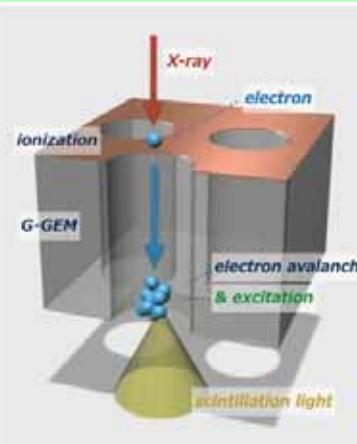
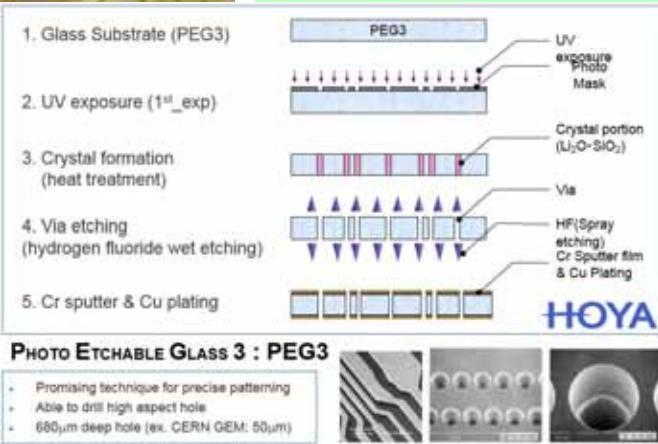
Electron Tracking method



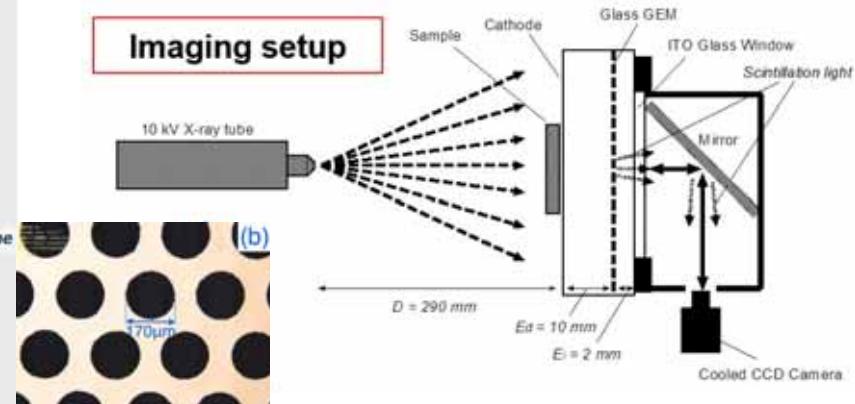
A. Ochi



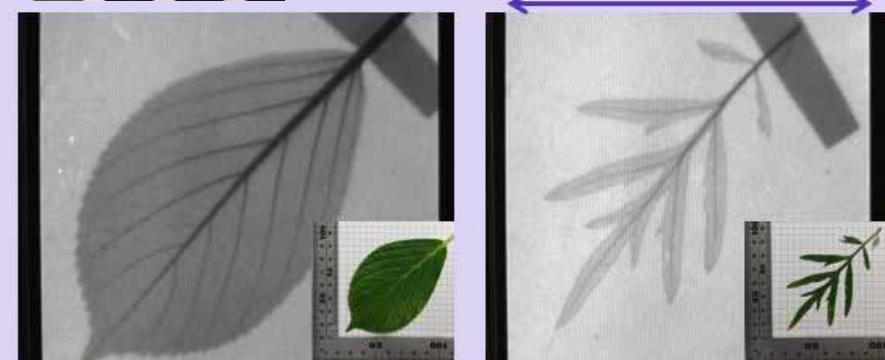
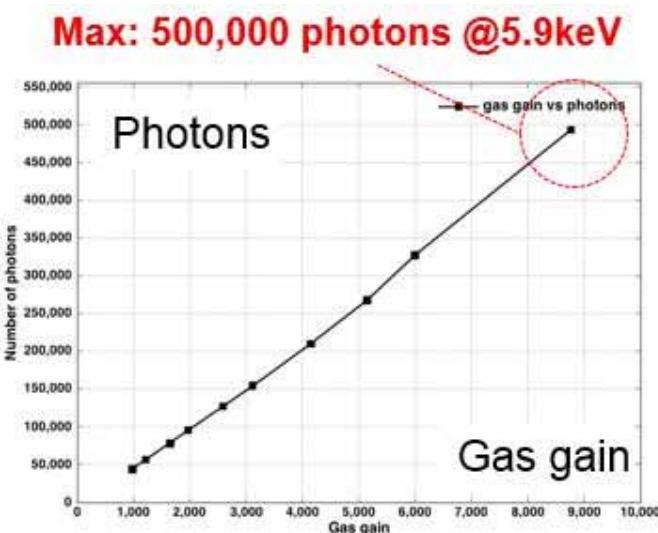
Scintillating Glass-GEM imager



Imaging setup



T. Fujiwara, et al., JINST 8 C12020, 2013



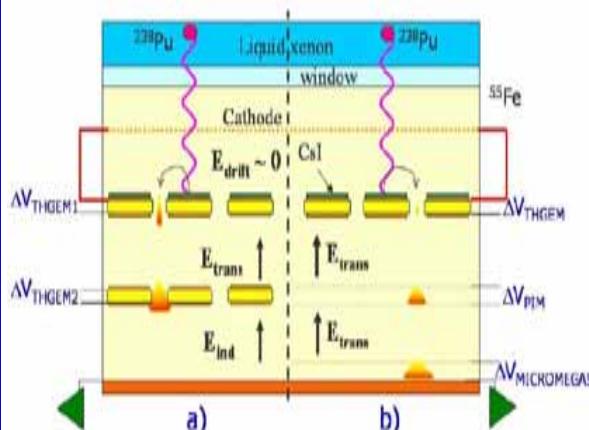
Obtained image of leaves (2 sec integration time)^[10]

Excellent spatial resolution ≈500μm
Quick imaging of low Z material with low energy X-rays (~7 keV)

Read-out elements of cryogenic noble liquid detectors

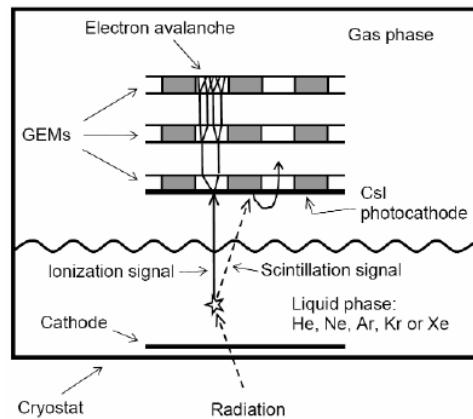
- Rear event detectors (ν , DM)
- Detecting the scintillation light produced in the noble liquids
- Options of scintillator light and ionization charge detection by a same detector !

with WINDOW



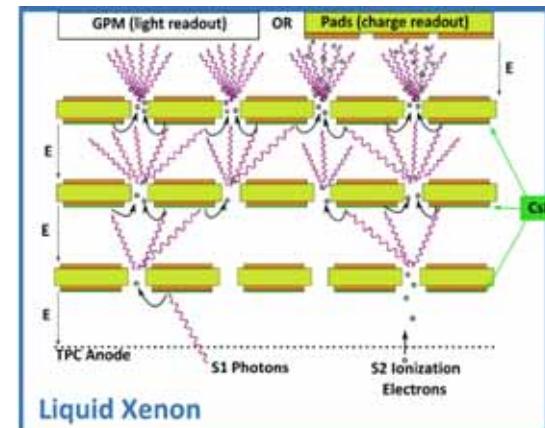
S.Duval et al., JINST 6 (2011) P04007

WINDOWLESS (2-PHASES)



A. Bondar et al., NIMA 556 (2006) 273

OPERATED IN THE CRYOGENIC LIQUID

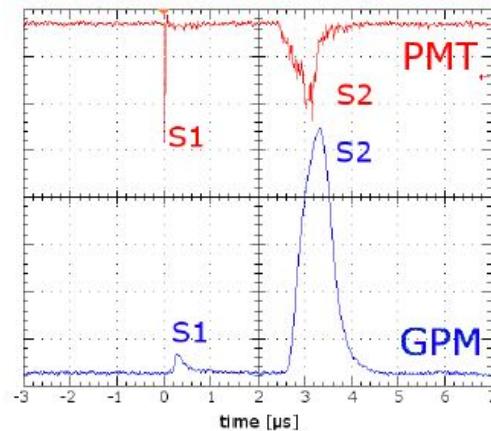


L.Arazi et al., JINST 8 (2013) C12004

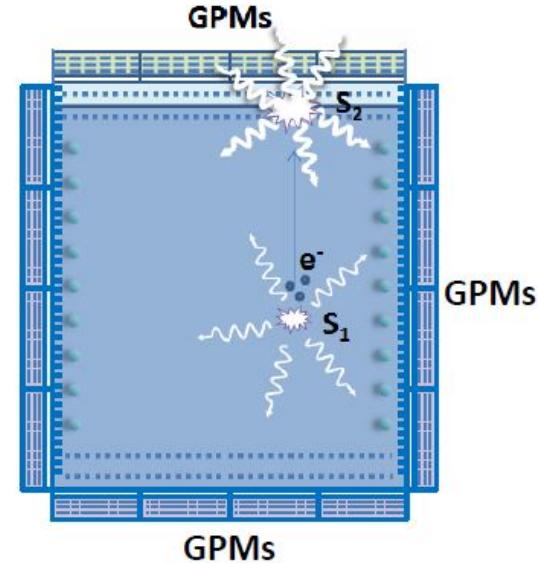
Triple THGEM Gaseous Photo-Multiplier for DM

- WIS R&D on **GPMs** for future multi-ton LXe TPCs for dark matter searches (within DARWIN)
- Aim for **4π coverage** – not practical with PMTs (cost, bulkiness) or SiPMs (dark count rate)
- Successful demonstration of 4" cryogenic **triple-THGEM GPM** with reflective CsI coupled to dual phase LXe TPC: ([arXiv:1509.02354](https://arxiv.org/abs/1509.02354))

- Stable gain $\sim 10^5$
- Large dynamic range: 1 – $O(10^3)$ photoelectrons
- 1 ns timing (~ 200 PEs)
- Expected PDE $\sim 15\%$ after optimization

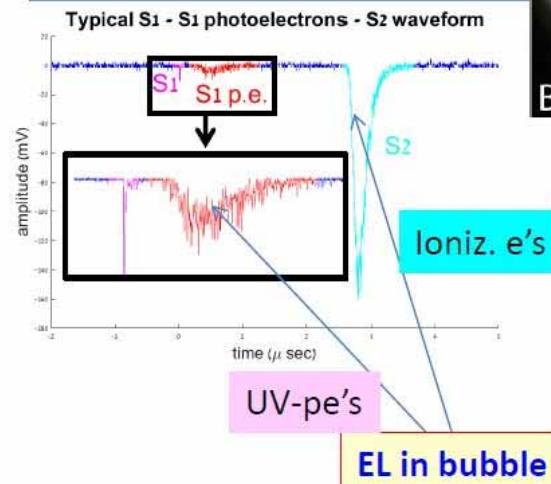
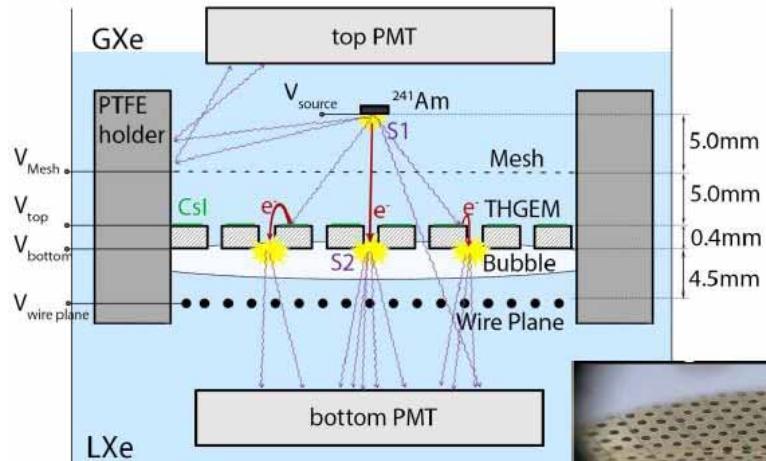


- Also: on-going R&D on **n/γ imaging** with pixilated readout ([arXiv:1501.00150](https://arxiv.org/abs/1501.00150))



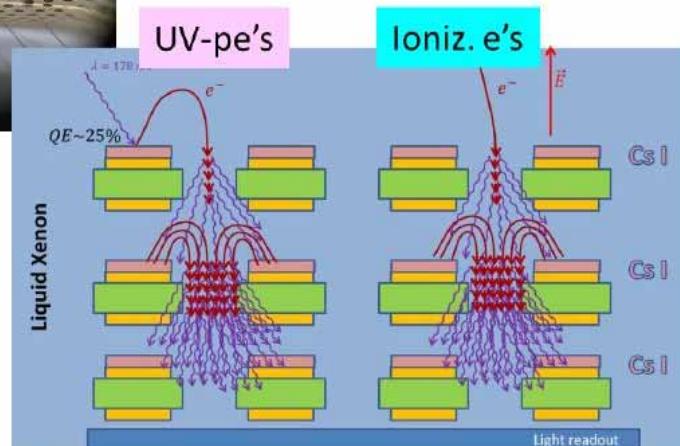
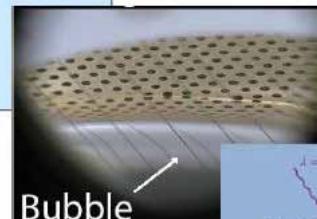
Bubble-assisted electroluminescence in LXe

A “local dual-phase” noble-liquid detector



TOWARDS LARGE-SCALE NOBLE-LIQ DETs

Energy resolution 5MeV alphas: $\sigma/E=7.5\%$
 Time resolution: $\sigma=10\text{ns}$
 Bubble (under THGEM, GEM) stable for days
 CsI on THGEM: high pe extraction

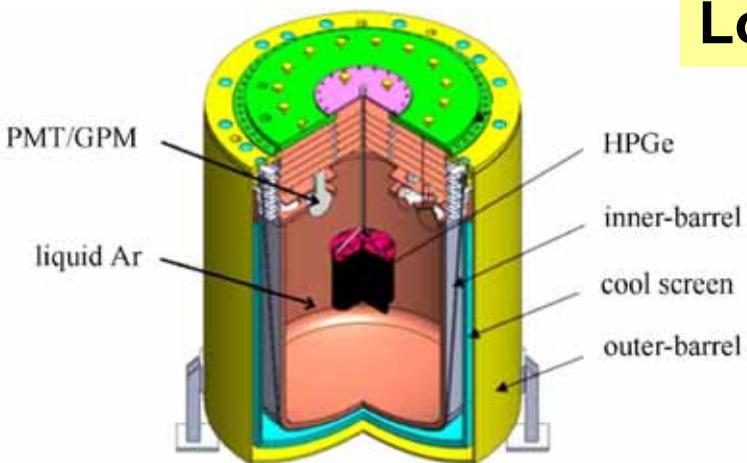


A Breskin MPGD 2015 Trieste

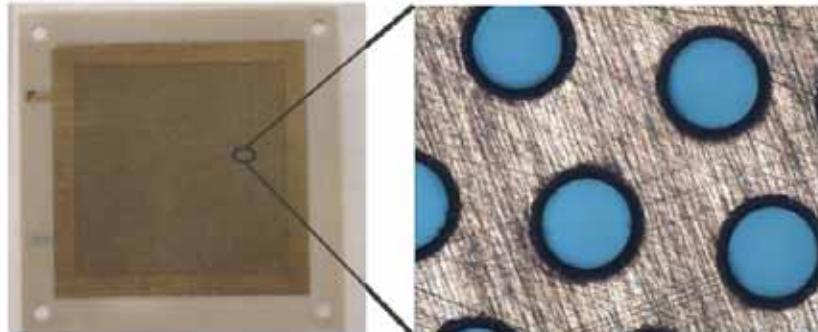
Breskin, J. Phys. Conf. Ser. 460(2013) 012020

33

PTFE THGEMs for CDEX



Low level radioactivity THGEMs



$t = 0.38 \text{ mm}$,
 $d = 0.3 \text{ mm}$,
 $p = 0.7 \text{ mm}$,
 $r = 30 \mu\text{m}$

Fig. 2. A photograph of the PTFE-THGEM; the enlarged part (right) shows the holes and the rims surrounding them.

Wen-Qing XIE et al, Chinese Physics C, Vol. 37, N. 11 (2013) 116001

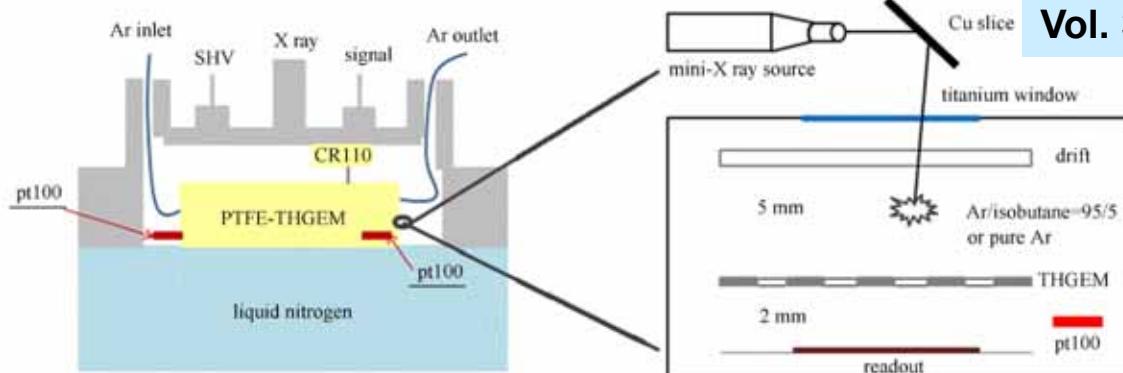
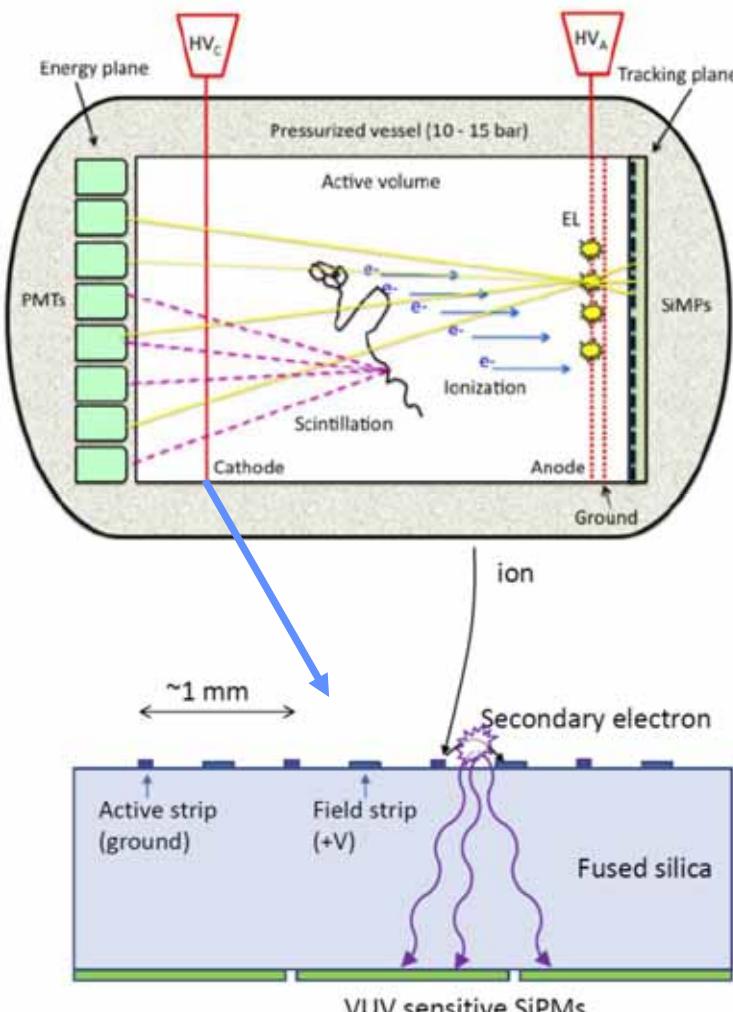


Fig. 3. Schematic diagram of the experimental setup. On the left is the test of the PTFE-THGEM at cryogenic temperature; and on the right is the internal structure of the PTFE-THGEM used.

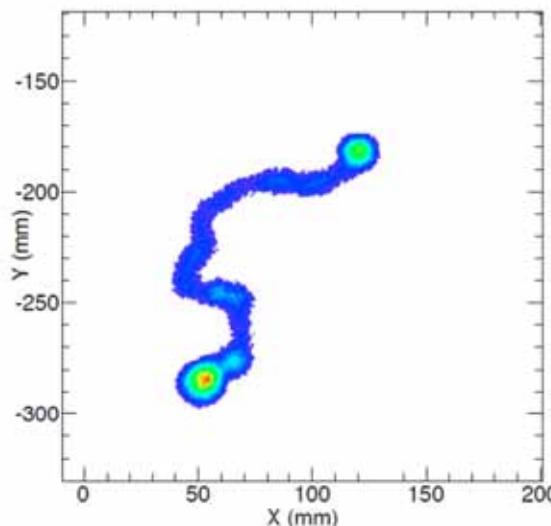
PMT disadvantages:
high cost, limited area,
unmatched spectral
response, radioactivity
 \rightarrow **PTFE THGEMs**
Almost no radioactivity,
demonstrated to operate
nicely at 117 K

Exotic proposal for NEXT $\beta\beta 0\nu$

LIOR ARAZI, WIZEMAN INSTITUTE OF SCIENCE



MC of $\beta\beta 0\nu$ event – 2 blobs at the ends



10 bar Xe, 300 V/cm
 10^5 electron ion pairs/ev.

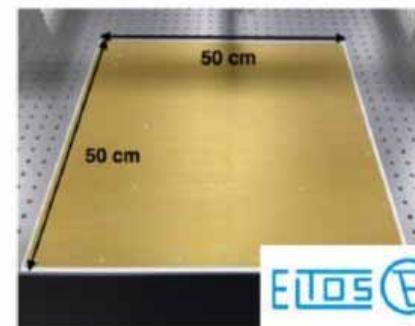
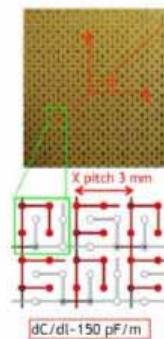
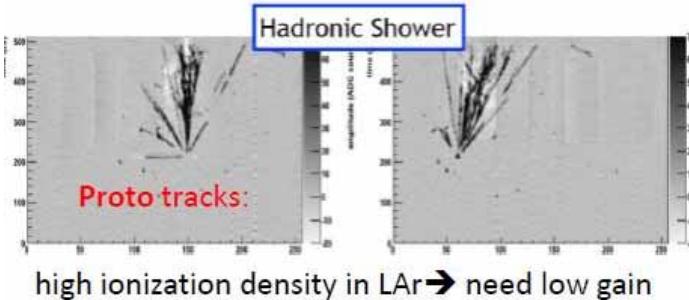
diffusion over $L = 2\text{ m}$:

$\sigma_{\text{electrons}} = 18 \text{ mm}$,
 $\sigma_{\text{ions}} = 1.8 \text{ mm.}$

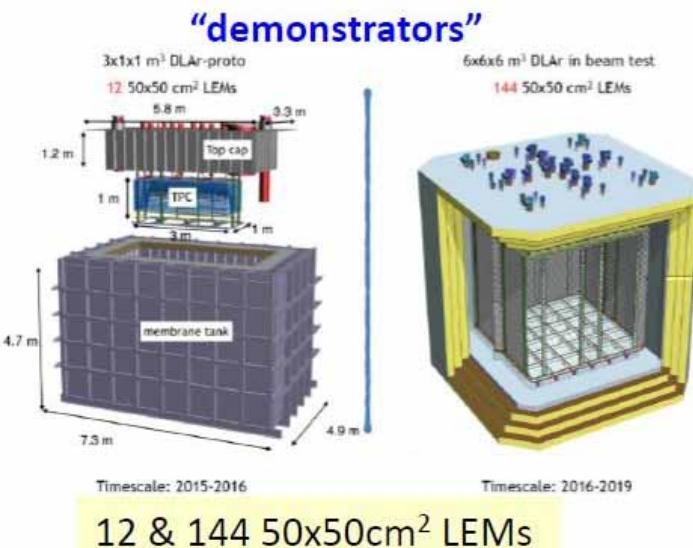
All Xe^+ atoms convert to Xe_2^+
 They reach the cathode in < 1 sec
 Xe_2^+ on active cathode strips emit e^- by "Auger Neutralization" if the surface has electron affinity < 0
 Secondary electrons go to field strips producing small avalanches with $O(10^3)$ photons per detected ion.

Dual-phase LAr LEM TPC

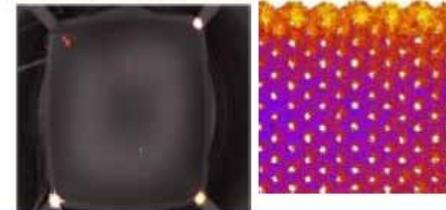
Goal: Neutrino oscillation experiments: WA105 (on ground) and future (underground) DUNE.



- Optimised values
- 40 μm rim
 - 1 mm FR4 thickness
 - 500 μm diameter hole
 - 800 μm hole pitch and hexagonal layout



DC: 5nA/LEM(50x50)
Stable gain ~20 (fine)



C.Cantini et al.,
JINST 10 P03017
(2015)

~3500V; spark on edges
(use COMPASS RICH solution?)
Charging up of rims: gain stabilizes. OK

DUNE: ~3000 LEMs (50x50)

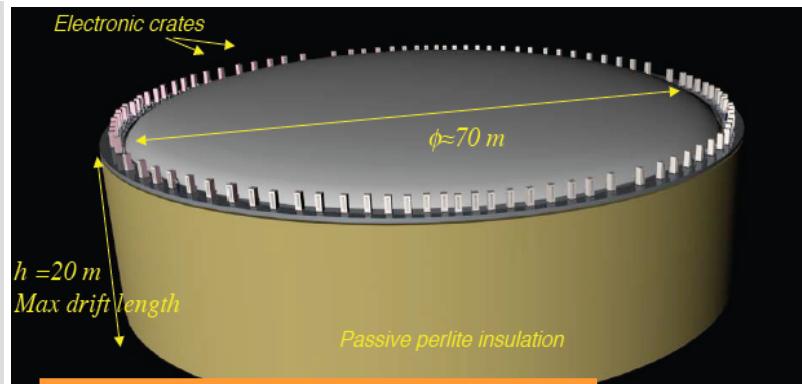
Ongoing R&D on RESISTIVE WELL concepts



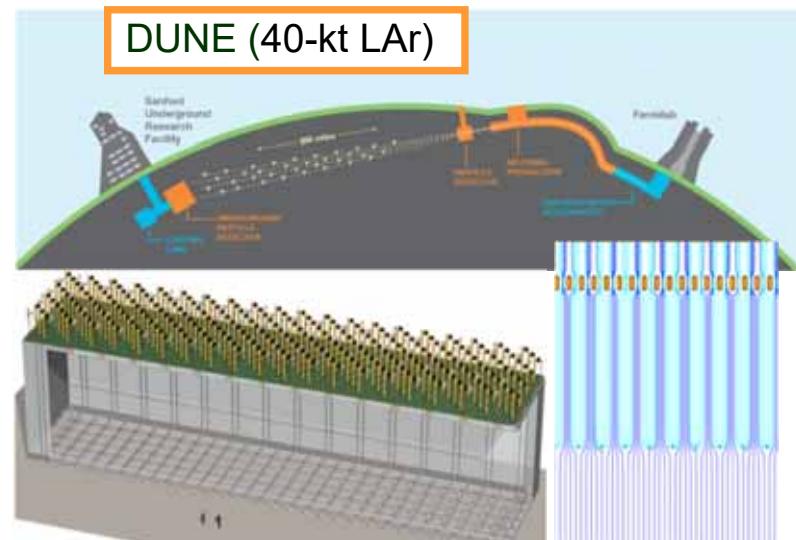
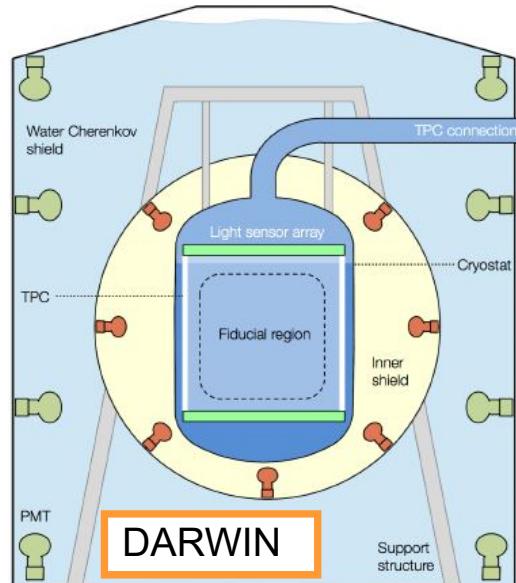
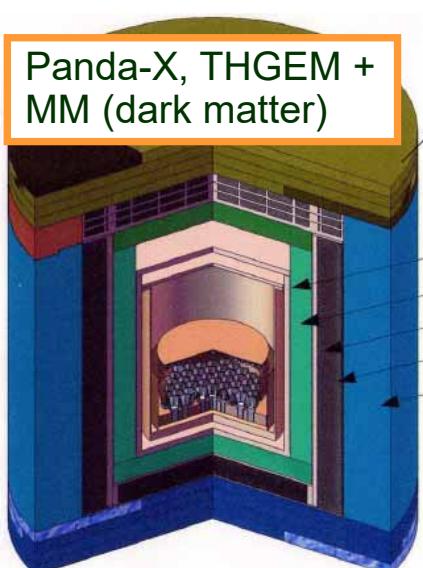
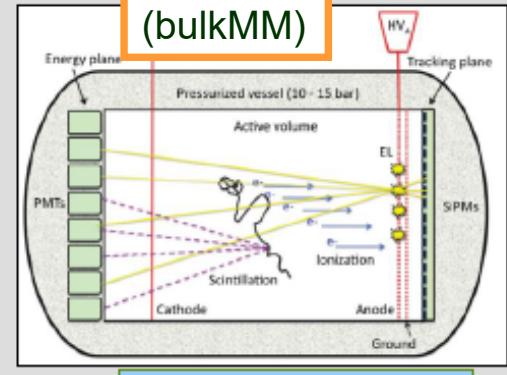
LARGE SIZE PROJECTS



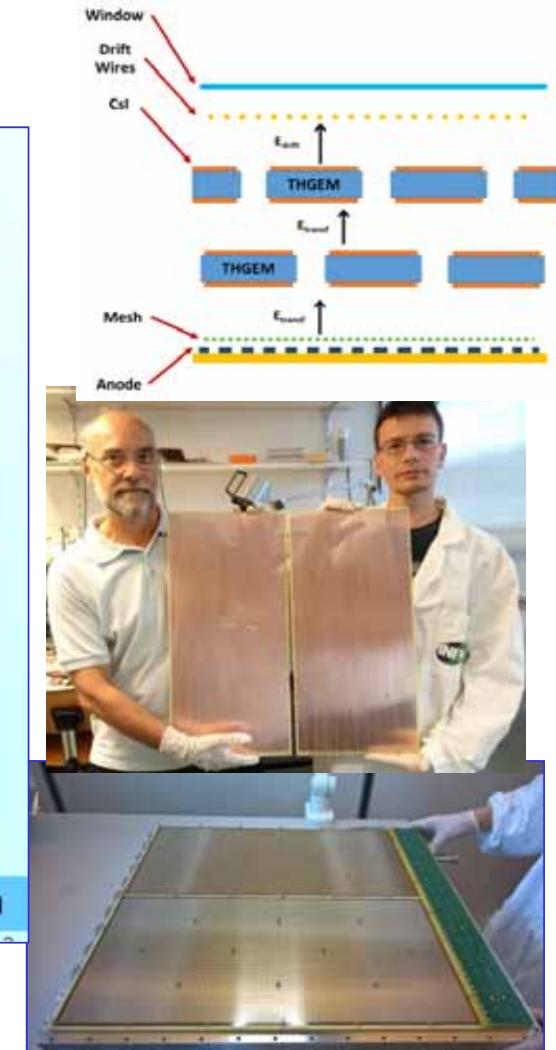
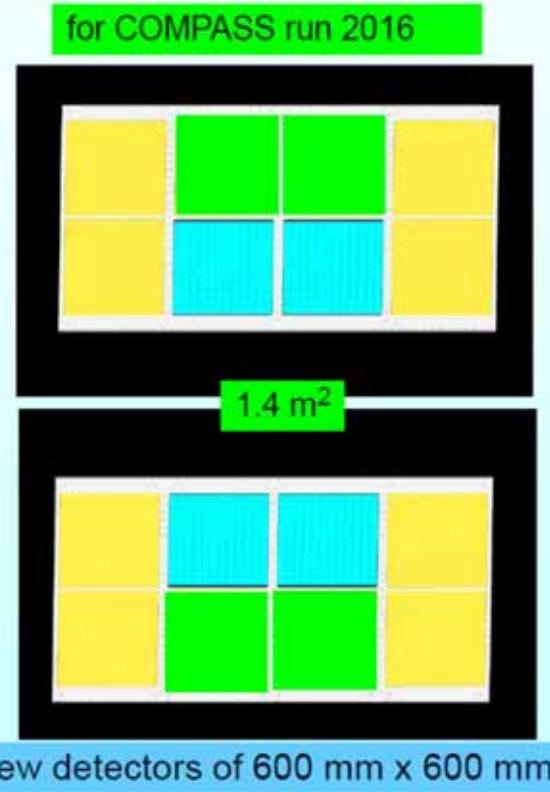
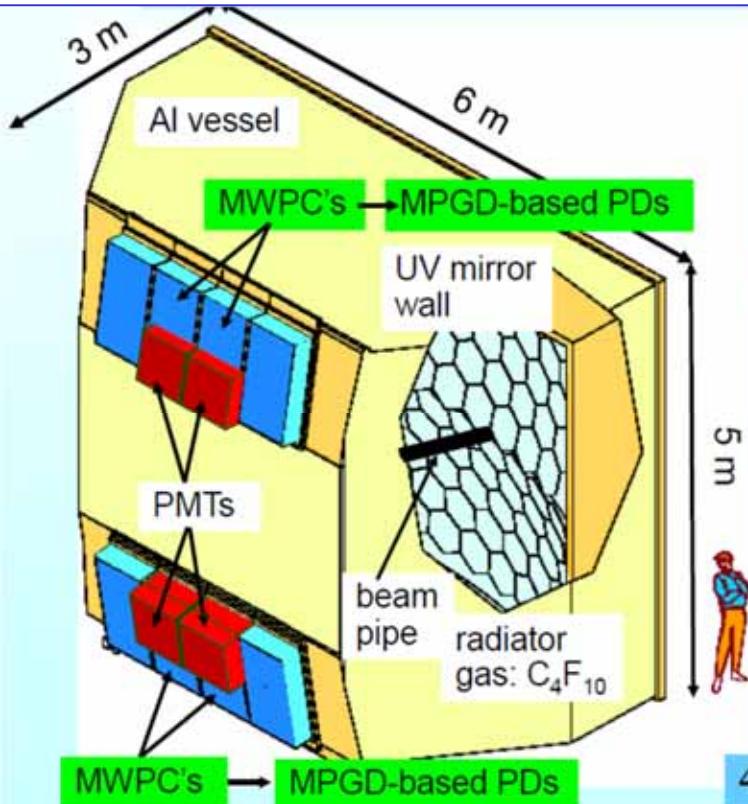
XENON (dark matter)



hep-ph/0402110
Venice, 2003



COMPASS RICH UPGRADE



talk by Shuddha Shankar Dasgupta



SUMMARY / CONCLUSIONS



- **GASEOUS PHOTON DETECTORS**
 - Most effective approach to instrument large surfaces at affordable costs
- **MPGD-BASED PHOTON DETECTORS**
 - Allow to overcome the limitations of open geometry gaseous PDs
 - A wide effort to refine and consolidate the technology
- **MANY APPLICATIONS OF MPGD-BASED PHOTON DETECTORS**
 - From PID to ν , DM, medical applications ...
 - First step toward large area: Hybrid THGEM+MM for COMPASS
- **BRIGHT FUTURE FOR:**
 - Inventions: new ideas, new techniques
 - Technology consolidation, new applications
 - Large scale projects