



# RICH 2016

## Status and Perspectives of Gaseous Photon Detectors

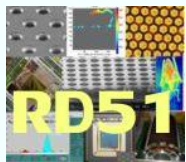
Fulvio Tessarotto  
INFN - Trieste





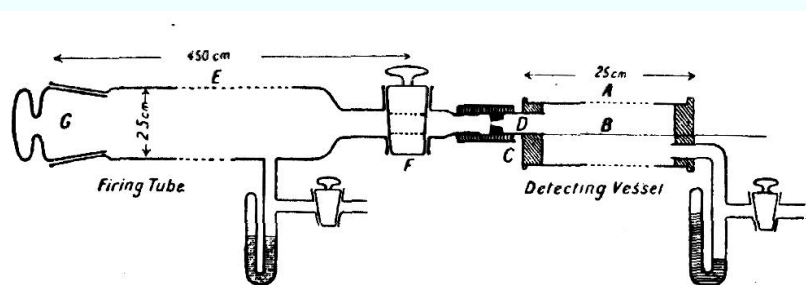
# Gaseous Photon Detectors

- Historical overview
- MWPCs with CsI Photocathodes
- GEM-based PDs
- THGEM-based PDs
- Other architectures
- Gaseous detectors for visible light
- Cryogenic gaseous photon detectors
- Large area coverage



# Glorious tradition: 100 years of gaseous detector developments

**1908: FIRST WIRE COUNTER  
USED BY RUTHERFORD IN THE STUDY OF NATURAL RADIOACTIVITY**

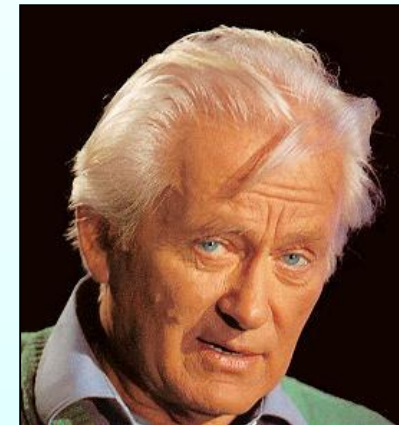
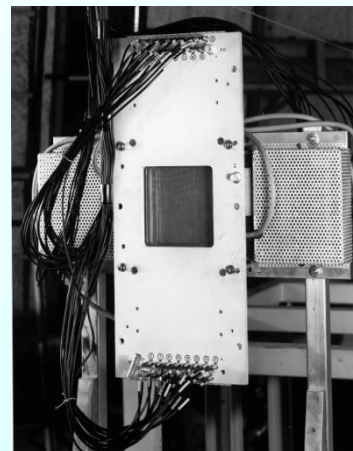


*E. Rutherford and H. Geiger,  
Proc. Royal Soc. A81 (1908) 141*



*Nobel Prize in Chemistry in 1908*

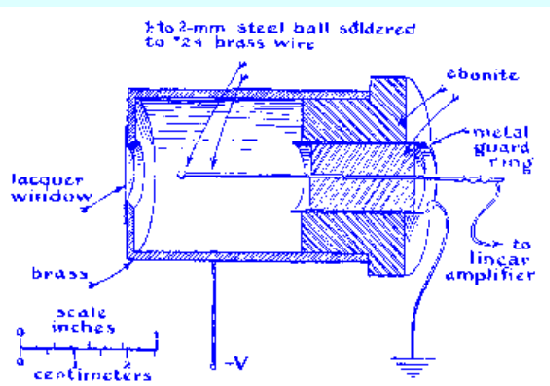
**1968: MULTIWIRE PROPORTIONAL CHAMBER**



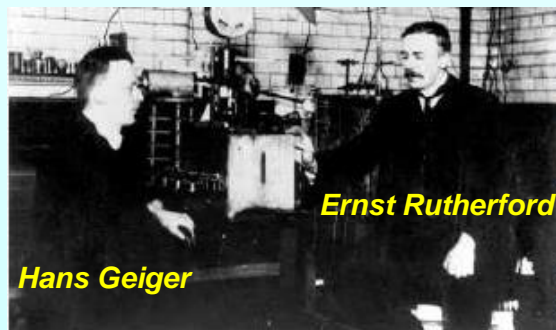
*Nobel Prize in 1992*

*G. Charpak, Proc. Int. Symp. Nuclear Electronics  
(Versailles 10-13 Sept 1968)*

**1928: GEIGER COUNTER  
SINGLE ELECTRON SENSITIVITY**



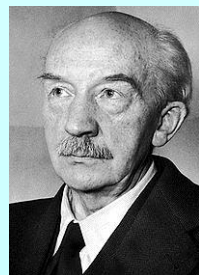
*H. Geiger and W. Müller,  
Phys. Zeits. 29 (1928) 839*



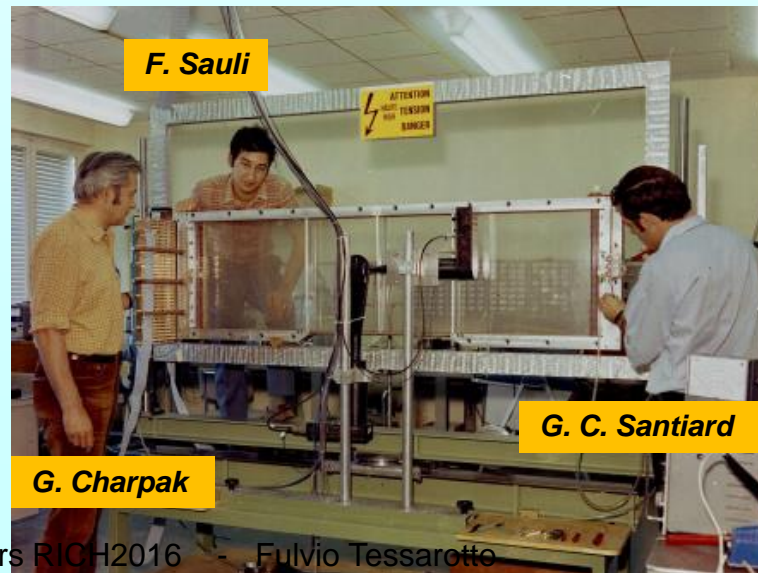
*Hans Geiger*

*Ernst Rutherford*

UK Science Museum



*Walther Bothe  
Nobel Prize in  
1954 for the  
"coincidence  
method"*



*F. Sauli*

*G. Charpak*

*G. C. Santiard*

# photon conversion and Cherenkov light



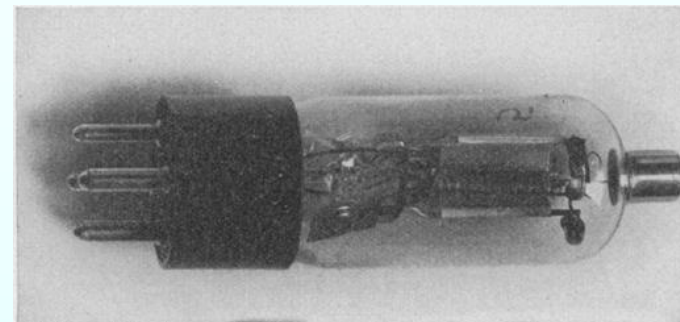
**John Sealy Townsend**



**Heinrich Rudolf Hertz**  
photoelectric effect, 1887



**A. Einstein, Nobel Prize in 1921**



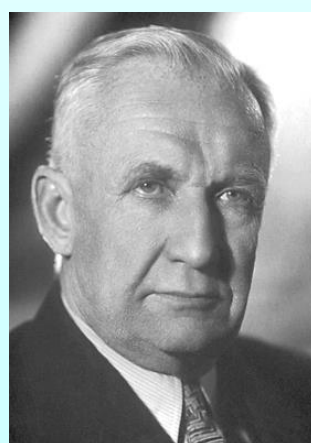
lams, H. E. and B. Salzberg, "The secondary emission phototube," Proc. IRE **23**, 55 (1935).



**Pavel Cherenkov 1904-1990**



**Ilya Frank**



**and Igor Tamm**

**Nobel Prize in 1958**



**Arthur Roberts 1912-2004**



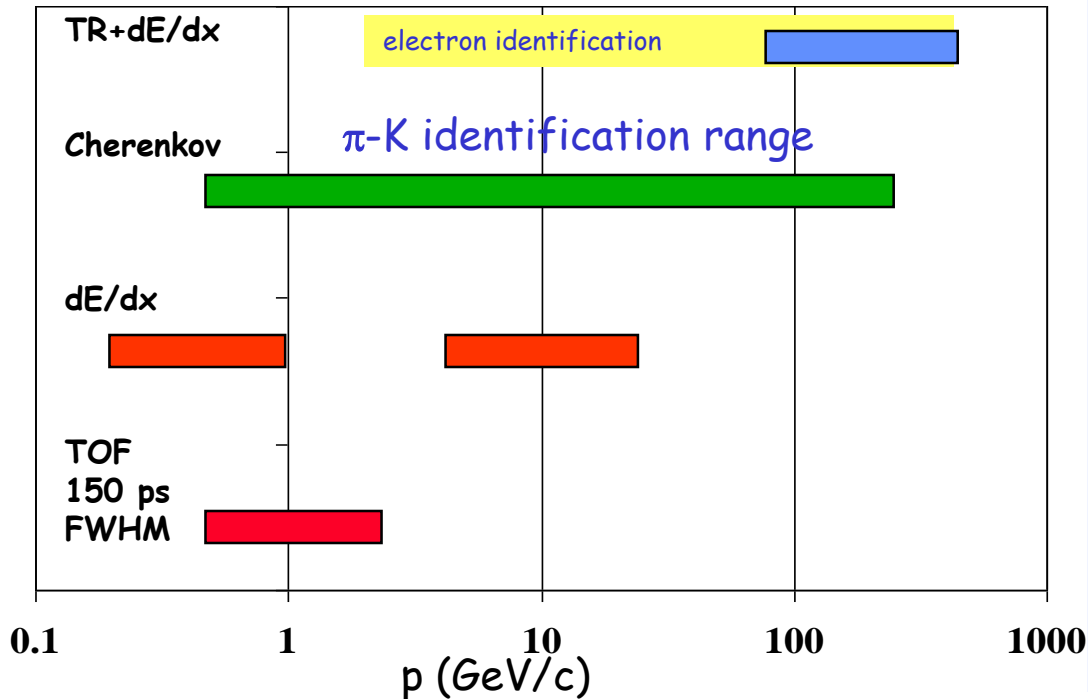
**Tom Ypsilantis 1928-2000**



# Motivation

- need for  $\pi$ -K identification from HEP Experiments
- Large momentum acceptance  $\rightarrow$  Cherenkov angle measurement technique
- Large angular acceptance  $\rightarrow$  large area of efficient single photon detection

## Particle Identification Techniques:

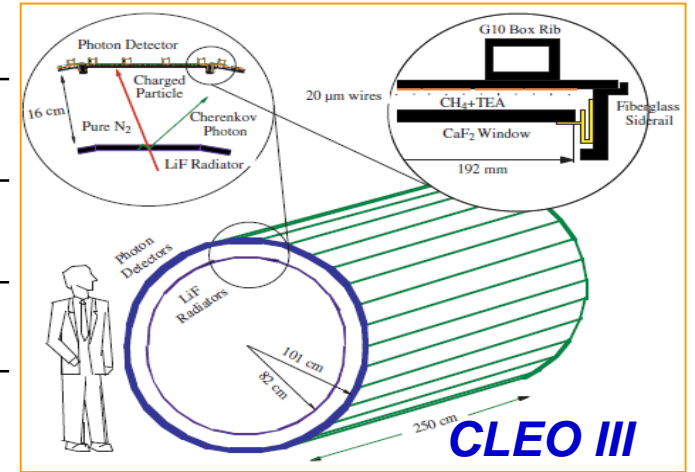
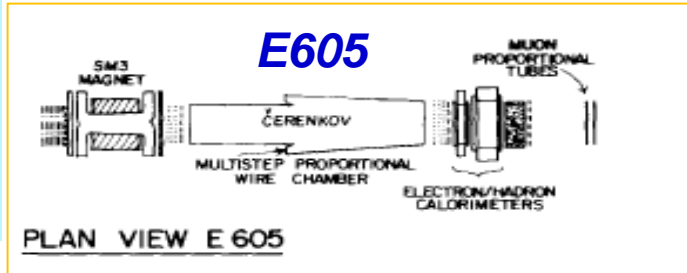
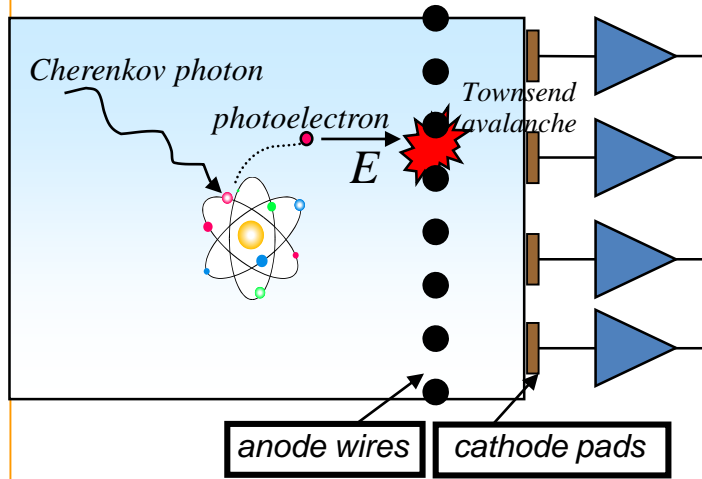
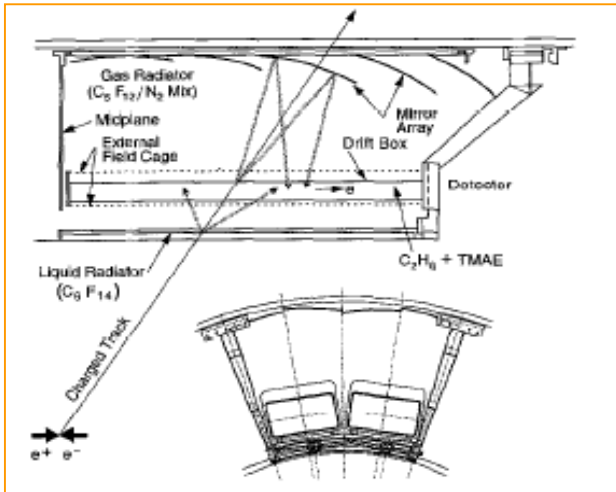


- 1970s: large area position sensitive gaseous detectors available
- Suitable photo-ionizing agent:
  - benzene: Seguinot-Ypsilantis NIM 142 (1977) 377,
  - TEA** (7.6 eV) NIM 173 (1980) 283,
  - TMAE** (5.3 eV) NIM 178 (1980) 125.
- a gas gain high enough to detect single photoelectrons
- $\rightarrow$  conflicting requirements because of the copious UV emission by the multiplication avalanche.
- solution: multistep avalanche chamber (Charpak-Sauli Phys. Lett. B 142 (1977) 377) or TPC

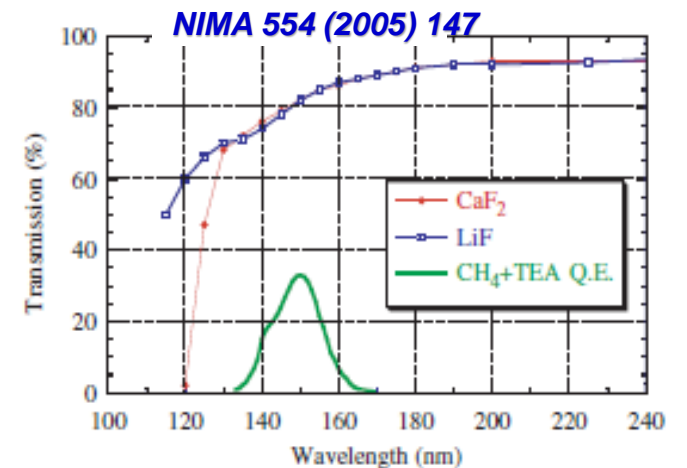
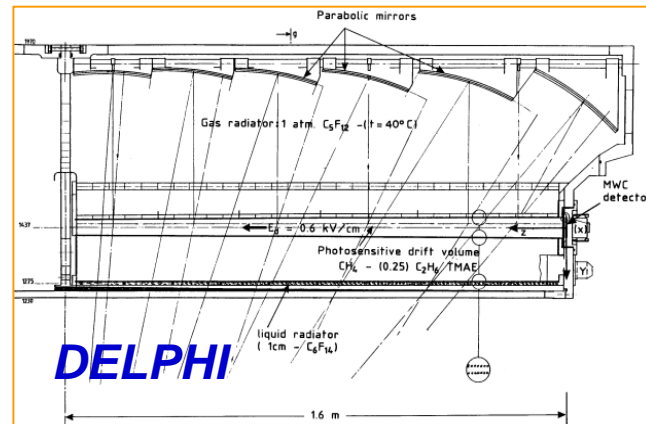
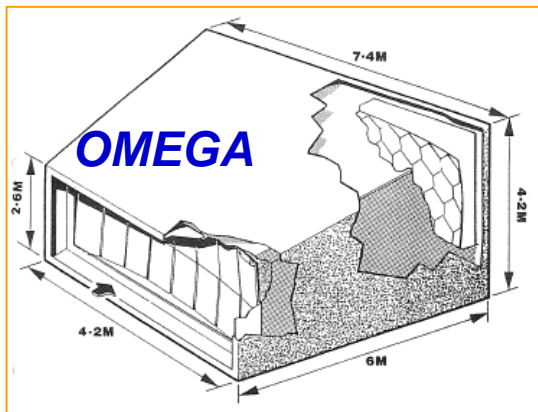
Gaseous detectors: 1) cheap, 2) magnetic insensitive, 3) low material budget

# RICH with large area gaseous PD's 1<sup>st</sup> generation: photoconverting vapours

## SLD - CRID



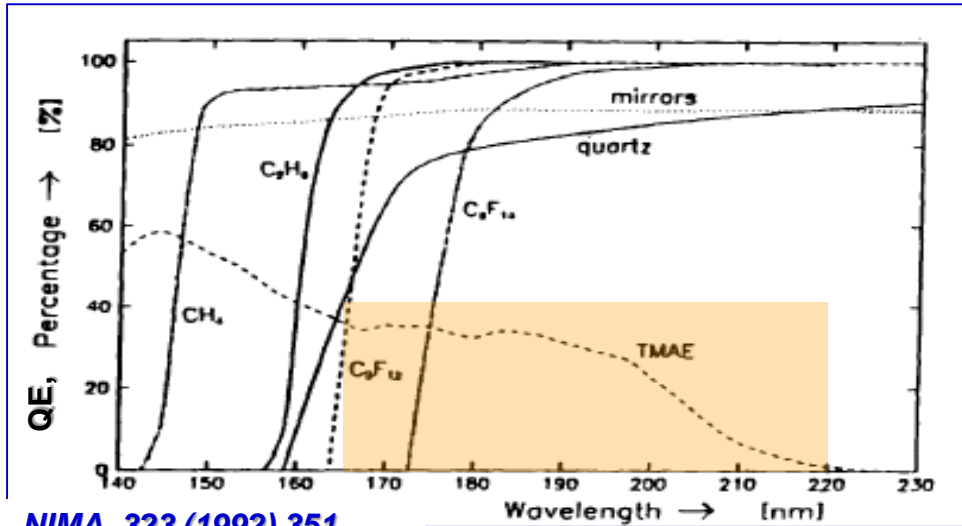
## TEA (Tri-Ethyl-Amine)



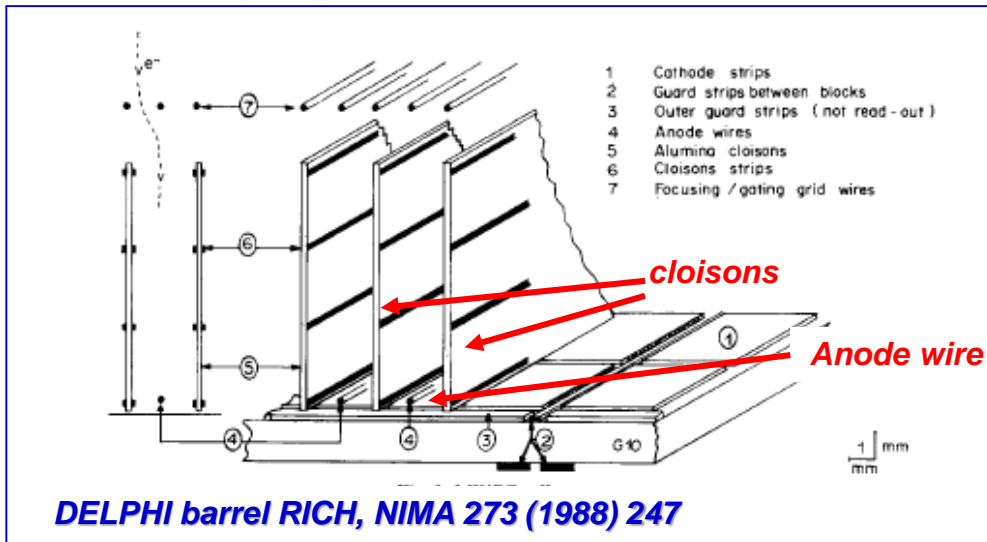
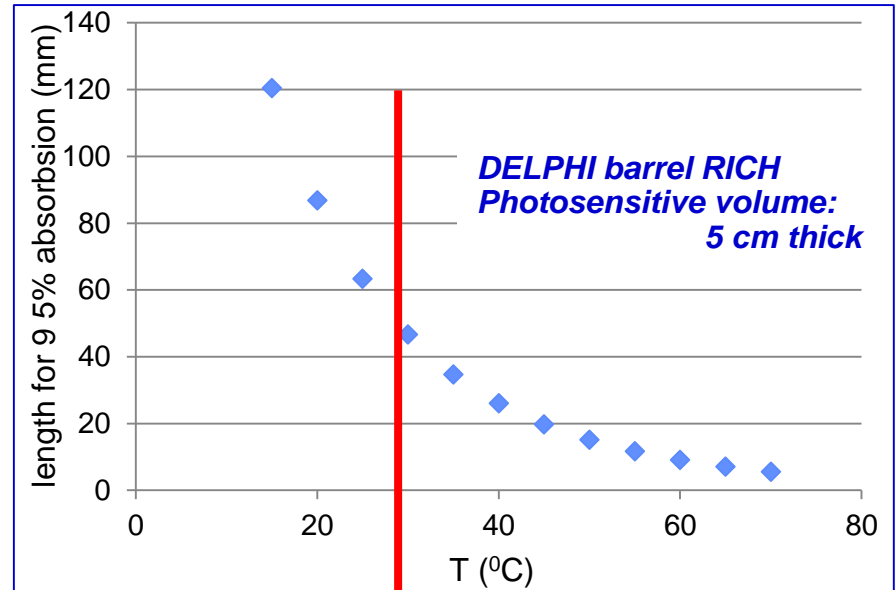
## TMAE (Tetrakis-Dimethylamine-Ethylene)



# TMAE



NIMA 323 (1992) 351



DELPHI barrel RICH, NIMA 273 (1988) 247

- thick photosensitive volume (slow photon detectors, parallax error)
- heating and temperature control ( $T_{\text{bubbling}} < T_{\text{operation}}$ )
- photon feed-back from amplification region (protections)
- chemically extremely reactive



# Thin CsI film

1956:  
CsI layer has large QE for photons with  $h\nu > 6$  eV  
(Philipp and Taft)

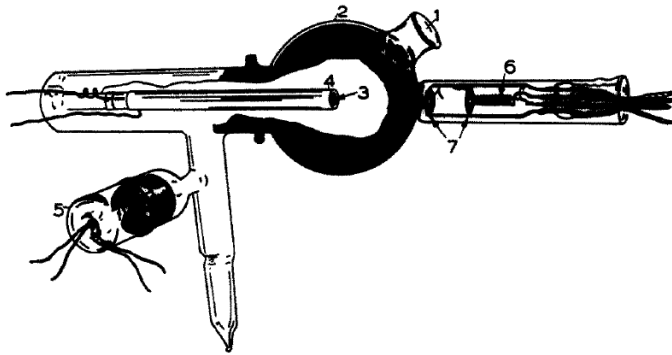


FIG. 1. Cutaway sketch of phototube; (1) 9741 glass bubble window, (2) graphite coated collector sphere 4 inches in diameter, (3)  $\frac{1}{8}$  inch glass tube, platinum painted, (4) nickel sleeve insulated from tube by glass beads, (5) ion gauge, (6) evaporating cylinder and helix platinum heater, (7) collimating shields.

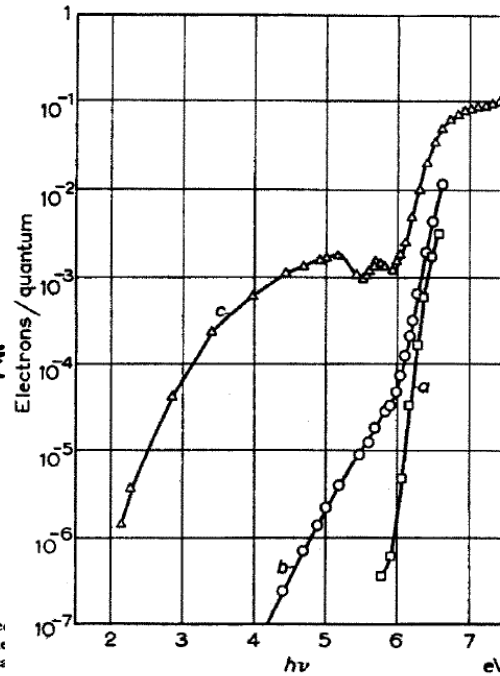
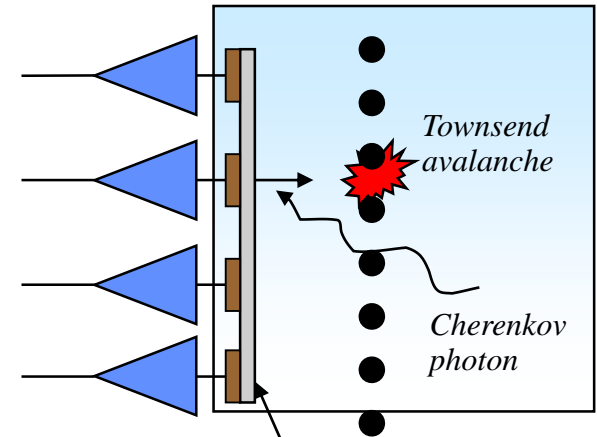


FIG. 2. Spectral distribution of the photoelectric yield for CsI surfaces: (a) thick film, (b) single crystal, (c) thin film evaporated in presence of excess Cs.

CsI is highly reactive with water: it took many years to develop appropriate substrate preparation, deposition method, handling technology for high QE gaseous PDs

Semitransparent photocathode:  
thin metallic film and precise thickness.

Reflective photocathode:  
no metallic film and non critical thickness  
(important for large area)



thin layer (300 - 500 nm) of CsI on a cathode pad plane

*J. Phys. Chem. Solids.* Pergamon Press 1956. Vol. 1. pp. 159-163.

## PHOTOELECTRIC EMISSION FROM THE VALENCE BAND OF CESIUM IODIDE

H. R. PHILIPP AND E. A. TAFT

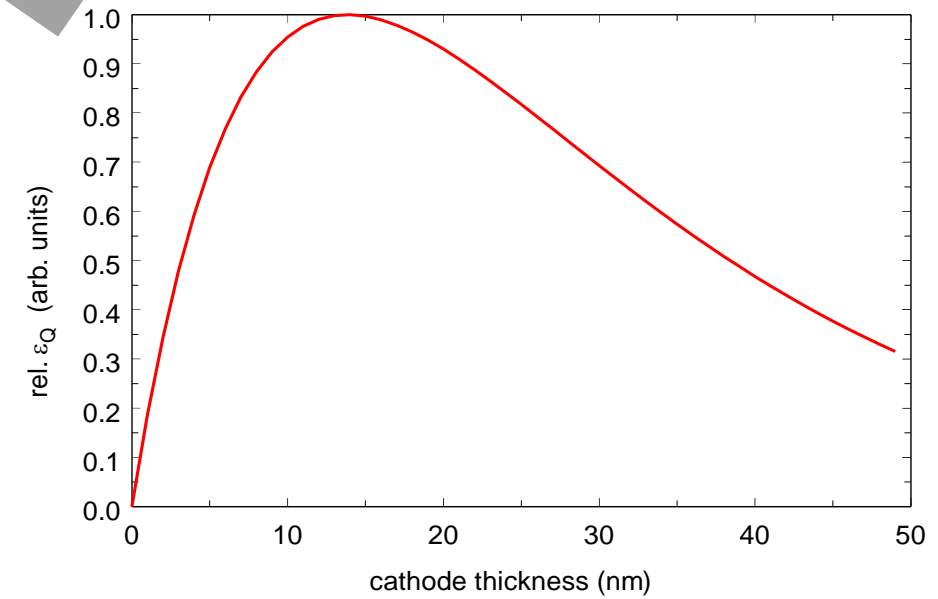
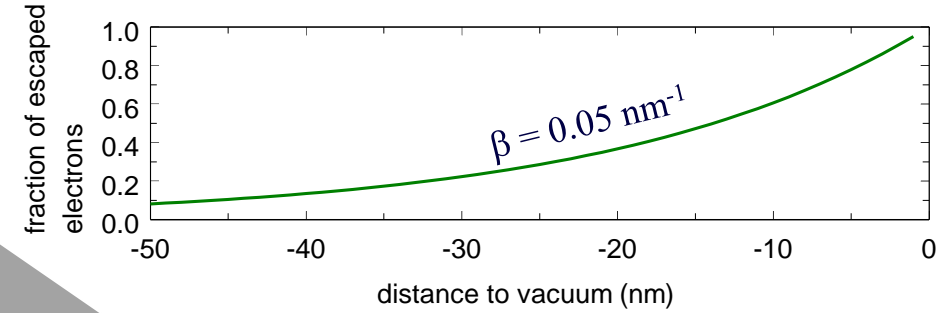
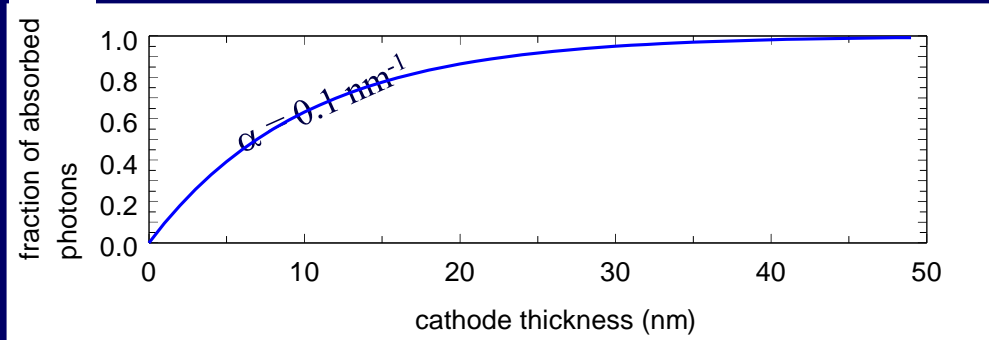
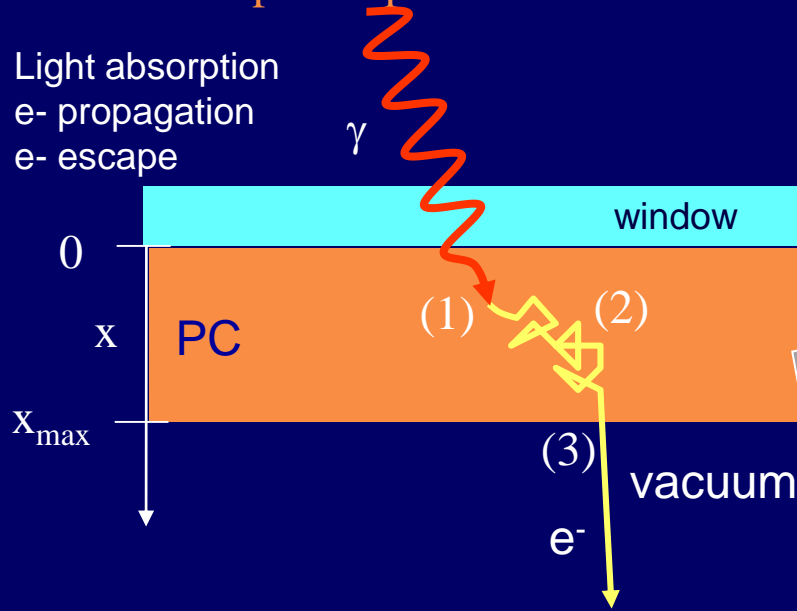
General Electric Research Laboratory, Box 1088, Schenectady, New York



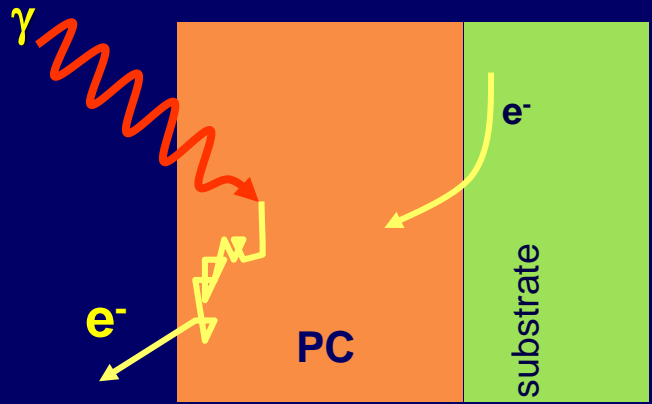
# Thin film photo cathodes

## Semitransparent photocathode:

- (1) Light absorption
- (2) e- propagation
- (3) e- escape



## Reflective photocathode:



# RD26: the technology of MWPCs + CsI

RD51



François Piuz

1992, F. Piuz et al. Development of large area advanced fast-RICH detector for particle identification at LHC operated with heavy ions

## TO ACHIEVE HIGH CsI QE:

### Substrate preparation:

Cu clad PCB coated by Ni (7  $\mu\text{m}$ ) and Au(0.5  $\mu\text{m}$ ), surface cleaning in ultrasonic bath, outgassing at 60  $^{\circ}\text{C}$  for 1 day

### Slow deposition of 300 nm CsI film:

1 nm/s (by thermal evaporation or e-gun) at a vacuum of  $\sim 10^{-7}$  mbar, monitoring of residual gas composition

### Thermal treatment:

after deposition at 60  $^{\circ}\text{C}$  for 8 h

### Careful Handling:

measurement of PC response, encapsulation under dry Ar, mounting by glove-box.

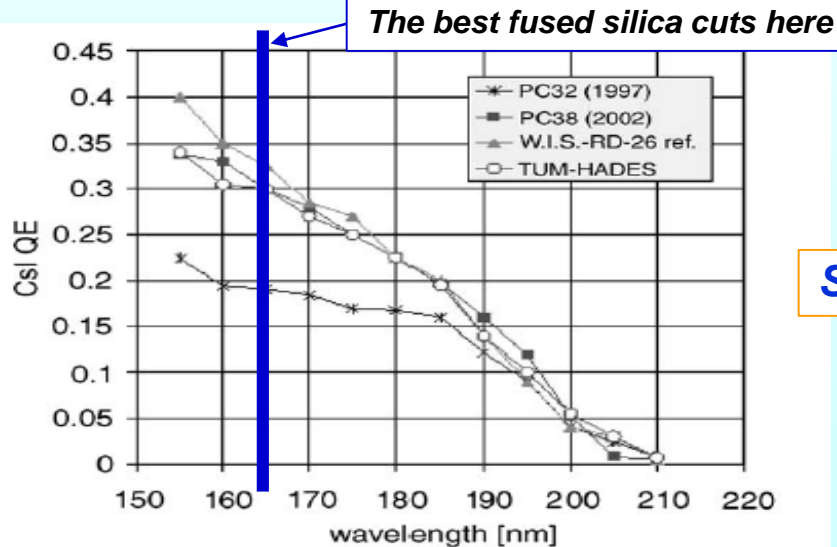
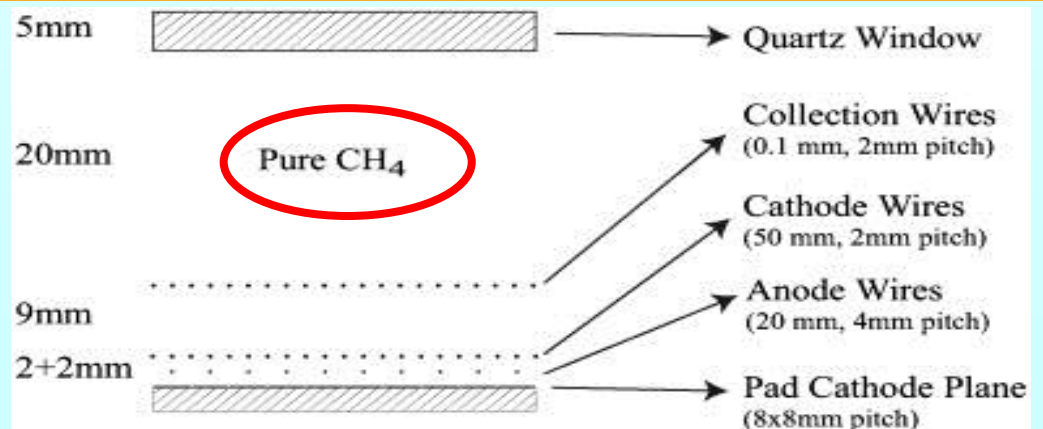


Fig. 1. The QE of CsI PCs produced at CERN for ALICE and at TUM for HADES, compared to that measured at the W.I.S. on small samples (reference for RD-26). PC32 is one of the four PCs equipping the ALICE-RICH prototype used in STAR at BNL.

A. Di Mauro, NIM A 525 (2004) 173.

## Schematic structure of the COMPASS Photon Detector:

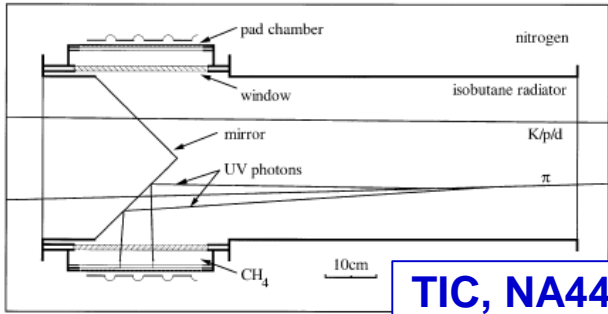




# RICH with large area gaseous PD's

## 2<sup>nd</sup> generation: MWPC's + CsI

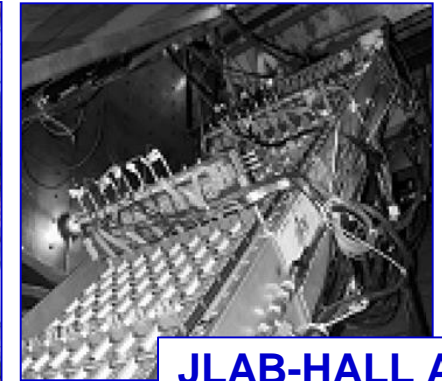
- MWPCs with solid state photocathode (the RD26 effort)



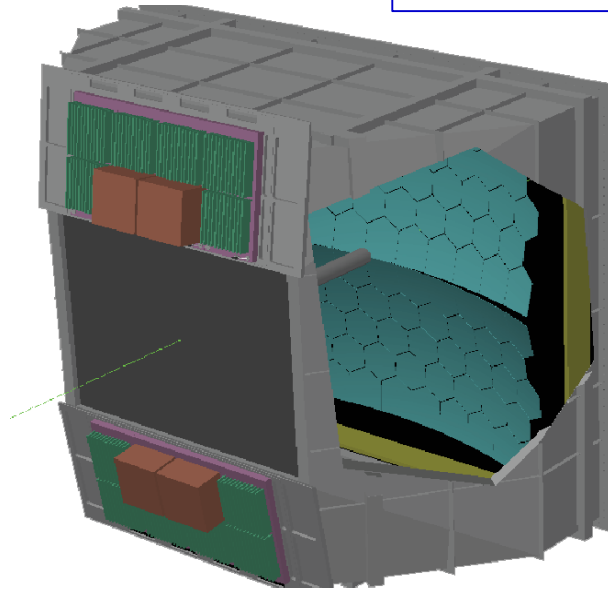
TIC, NA44



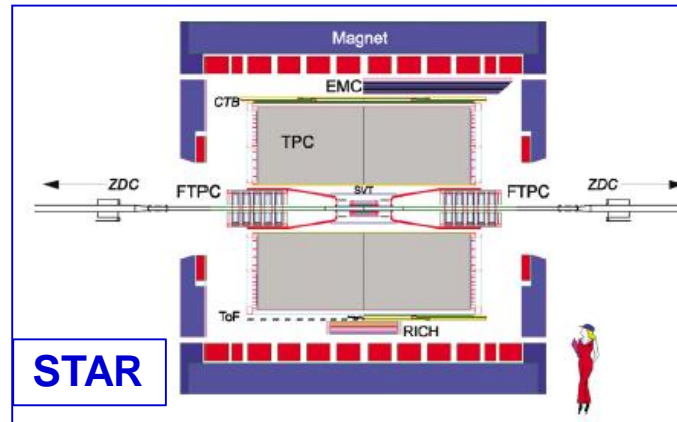
ALICE-HMPID  
2009  
CsI > 10 m<sup>2</sup>



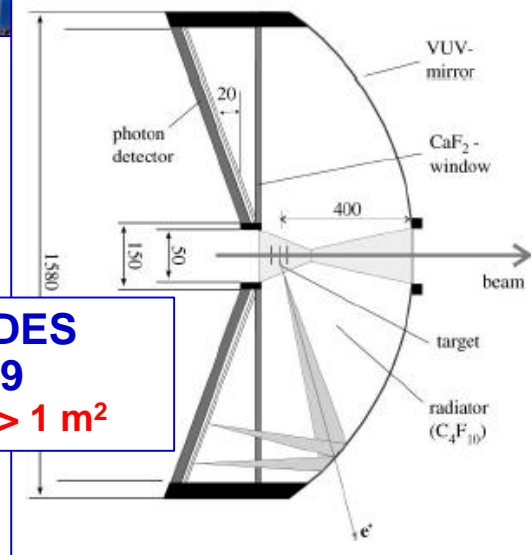
JLAB-HALL A



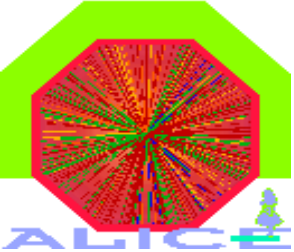
COMPASS RICH-1 2002  
CsI > 5 m<sup>2</sup>



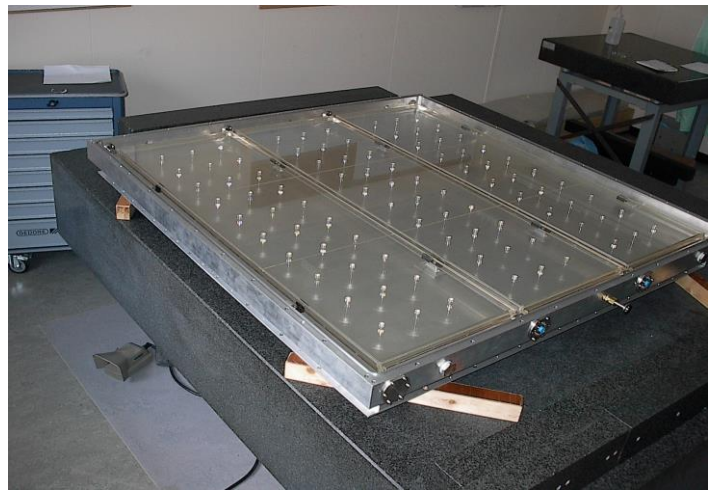
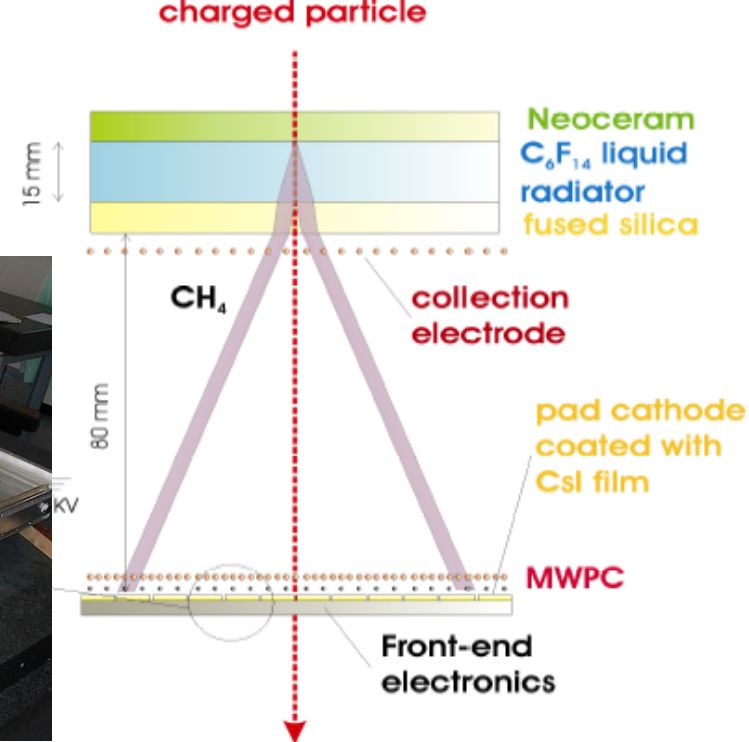
STAR



HADES  
1999  
CsI > 1 m<sup>2</sup>



# ALICE HMPID

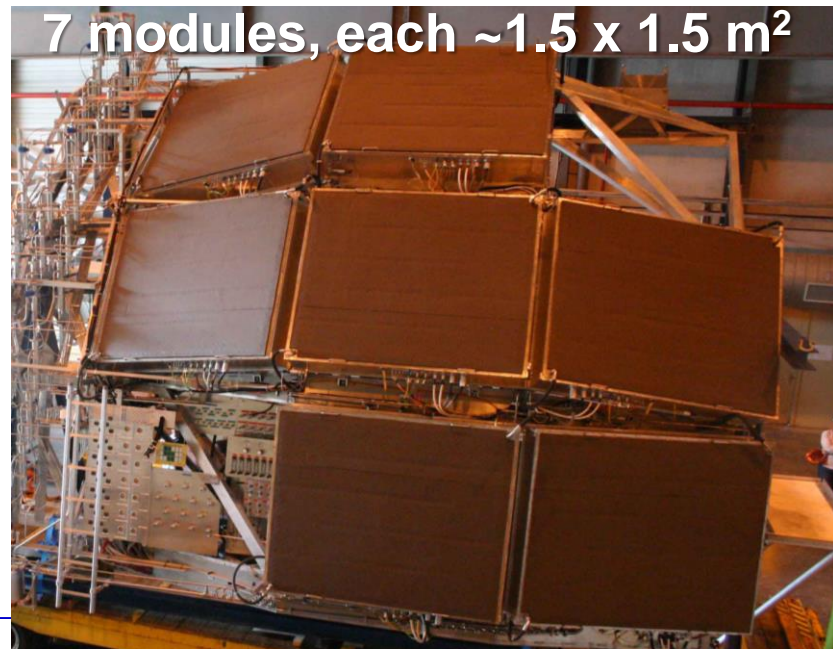


**RADIATOR:** 15 mm liquid C<sub>6</sub>F<sub>14</sub>,  
n ~ 1.2989 @ 175nm, β<sub>th</sub> = 0.77

**PHOTON CONVERTER:** Reflective layer of CsI (QE ~ 25% @ 175 nm)

**PHOTOELECTRON DETECTOR:** MWPC with CH<sub>4</sub> at atmospheric pressure (4 mm gap) HV = 2050 V.

- Analogue pad readout

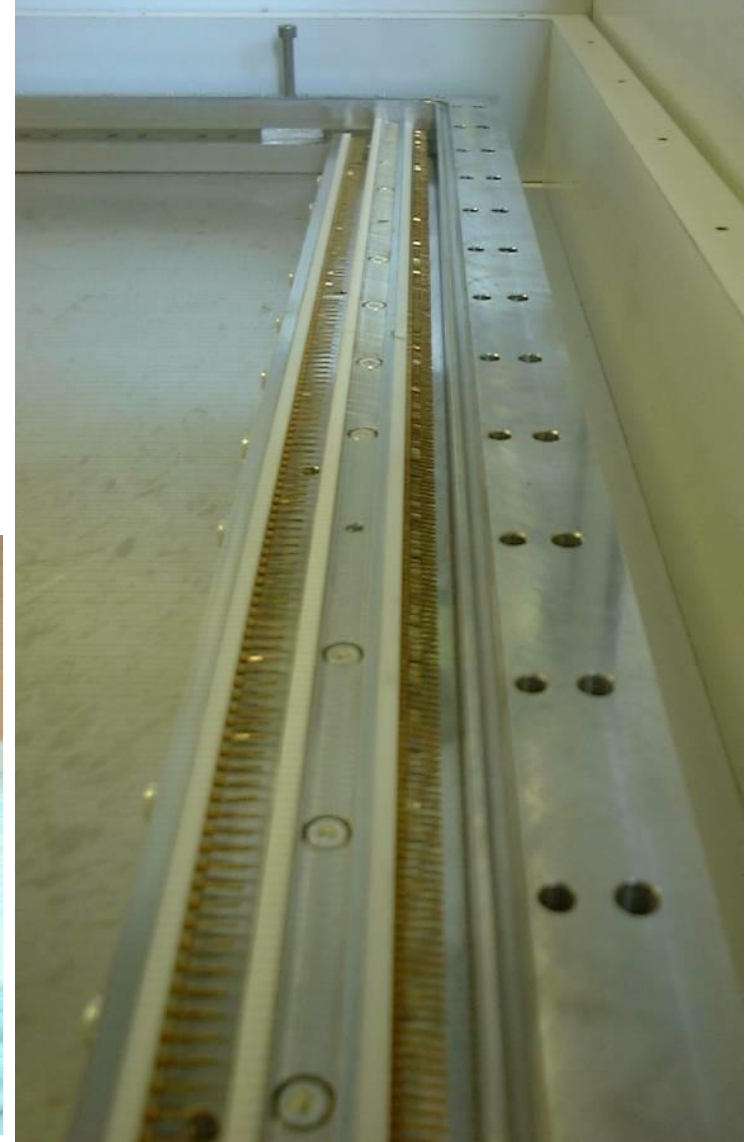
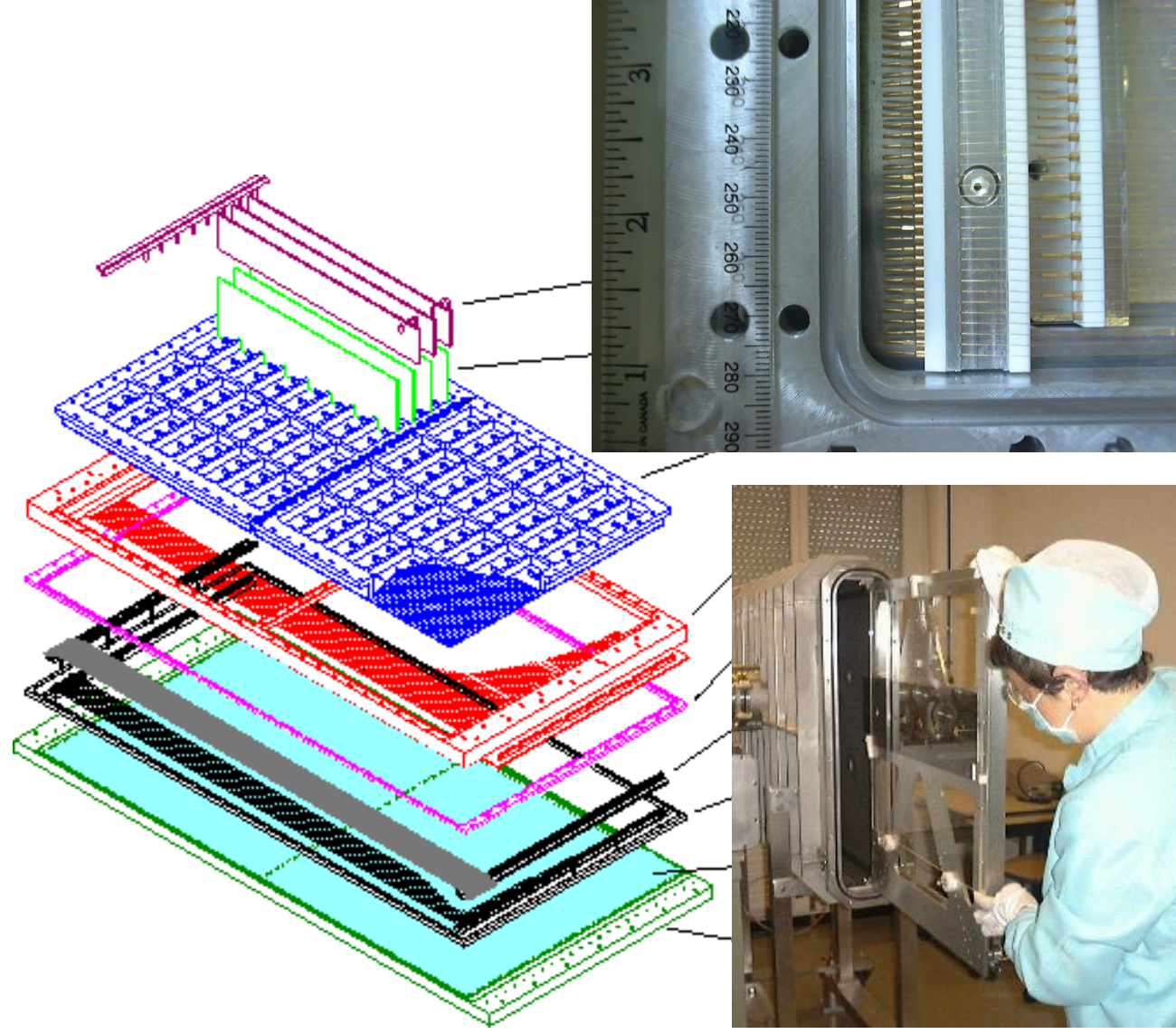


7 modules, each ~1.5 x 1.5 m<sup>2</sup>



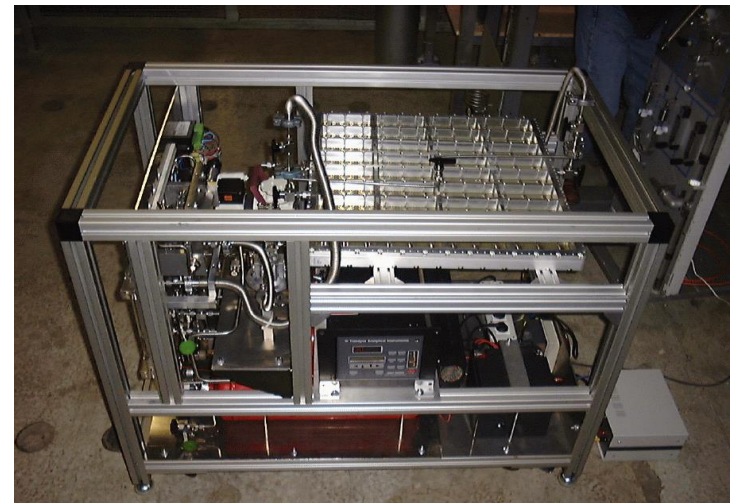
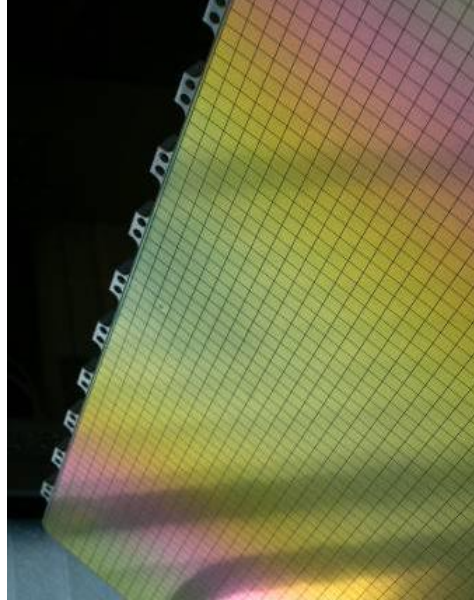
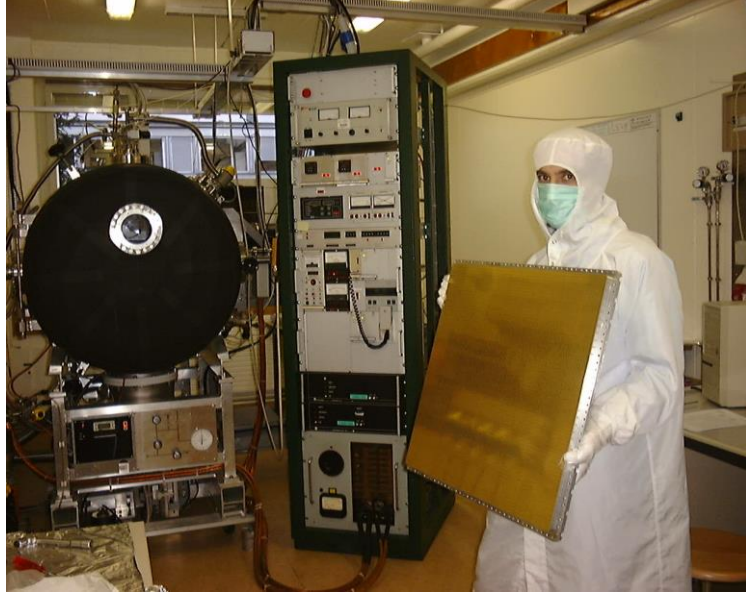


# COMPASS MWPC's with CsI





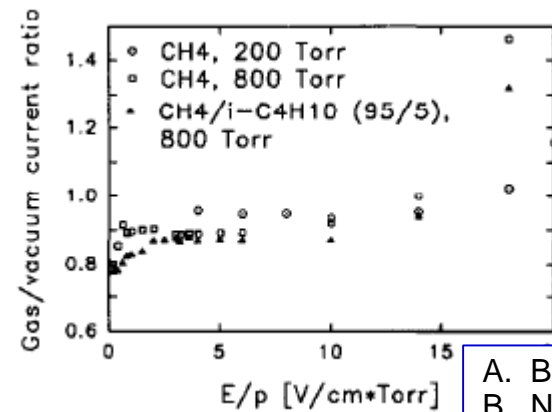
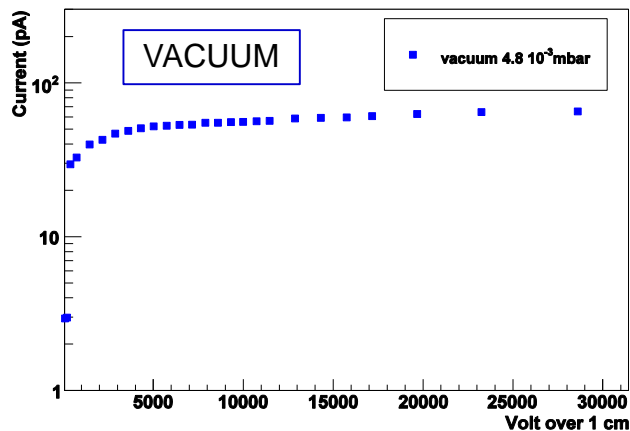
# COMPASS photocathodes



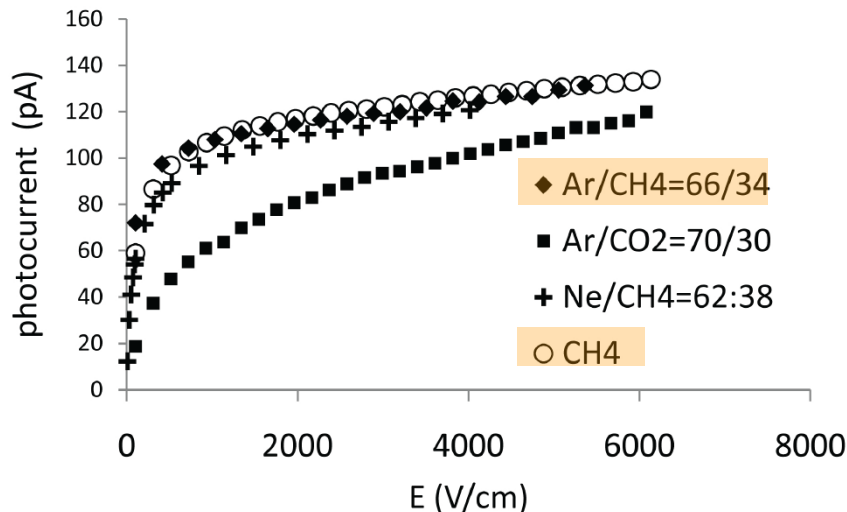


# PHOTOELECTRON EXTRACTION

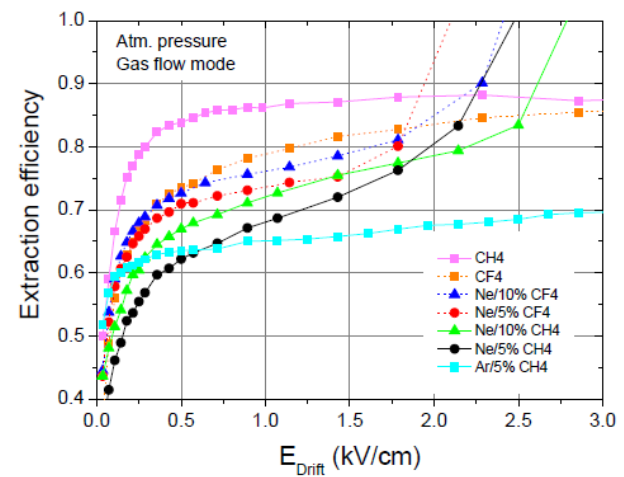
## Photoelectron extraction from a CsI film, the role of gas and E



A. Breskin et al.,  
B. NIM A 367 (1995) 342



M. Alexeev et al., NIM A (2010) in press



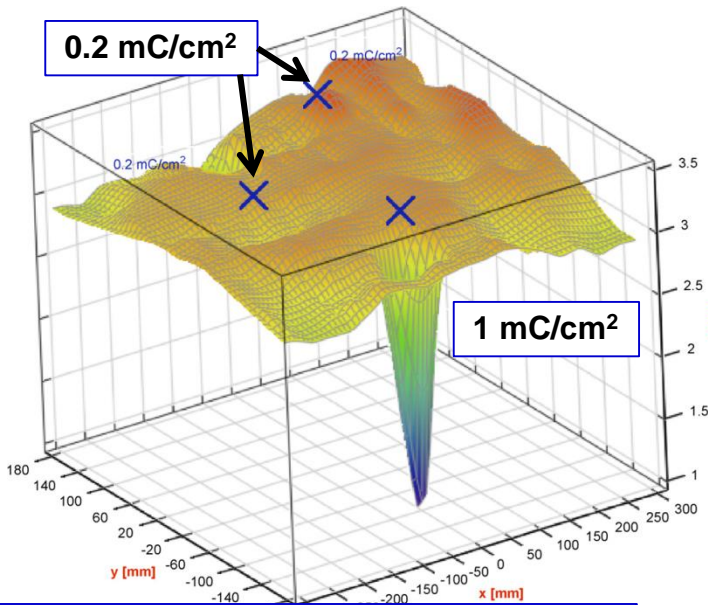
C. D. R. Azevedo et al., 2010 JINST 5 P01002



# MWPCs with CsI: the limits

- **Severe recovery time (~ 1 d) after detector trips**
  - Ion accumulation at the photocathode
- **Feedback pulses**
  - Ion and photons feedback from the multiplication process
- **Aging after integrating a few mC / cm<sup>2</sup>**
  - Ion bombardment of the photocathode

moderate gain:  $< 10^5$   
 (effective gain:  $< 1/2$ )  
 not fast



**MWPCs → slow signal formation**  
 + low gain → “slow” electronics (signal integration, low noise level)

- *Gassiplex FE* : integration time ~ 0.5 μs, time res > 1 μs
- *APV (COMPASS RICH-1 upgrade)* : resolution ~ 400 ns

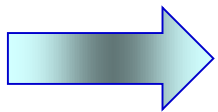
→ *Detector memory, i.e. not adequate for high rates*

H. Hoedlmoser et al., NIM A 574 (2007) 28.





- **Reduced photon feed-back and Ion BackFlow (IBF)**
  - Reduced ageing
  - High gain → high photoelectron detection efficiency
- **Intrinsically fast gaseous detectors**
  - Short integration time
  - High rate environments



**MICROPATTERN GASEOUS DETECTORS**

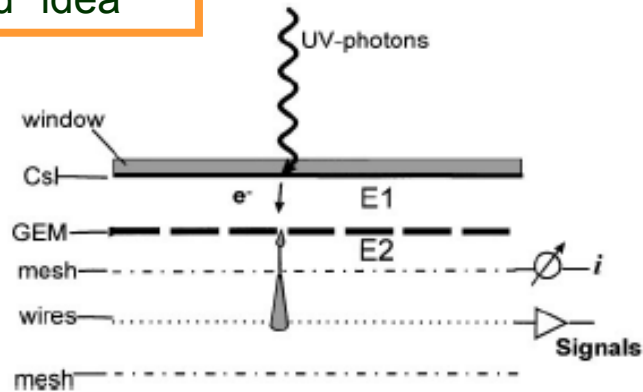


# ION & PHOTON BLOCKING GEOMETRIES

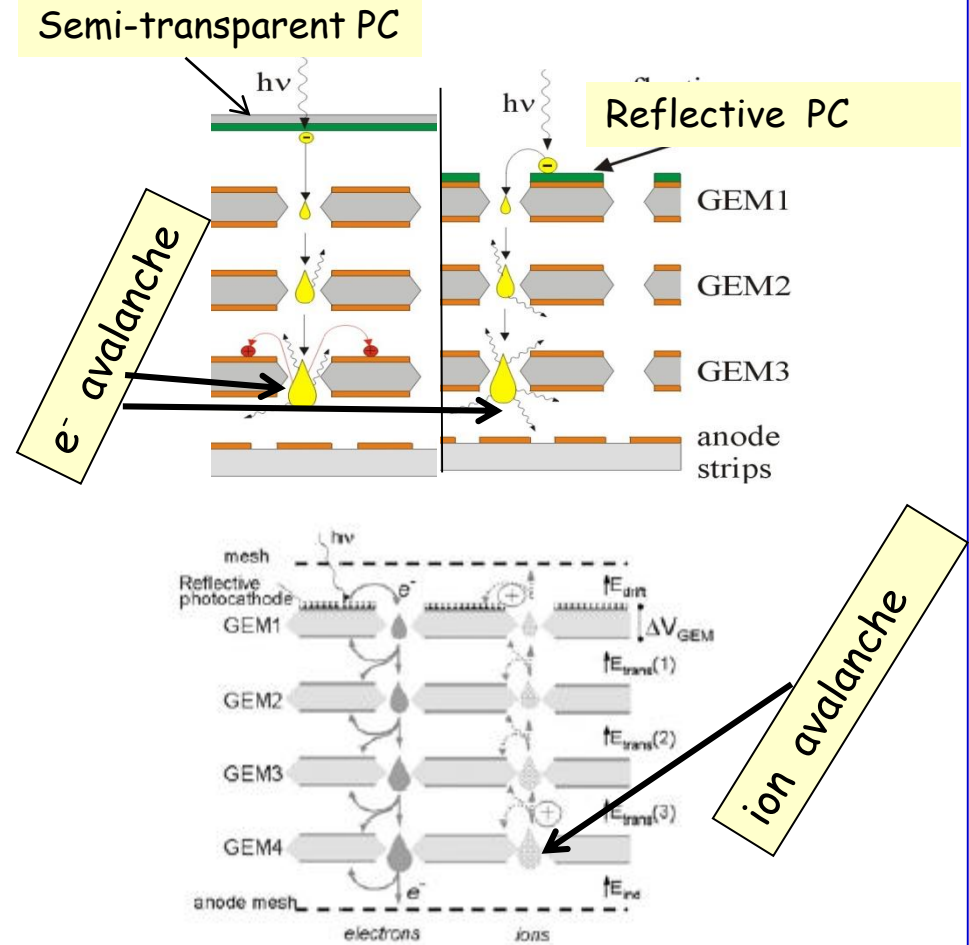
## GEM-based PDs

NO photon feedback  
Reduced ion feedback

### An "old" idea



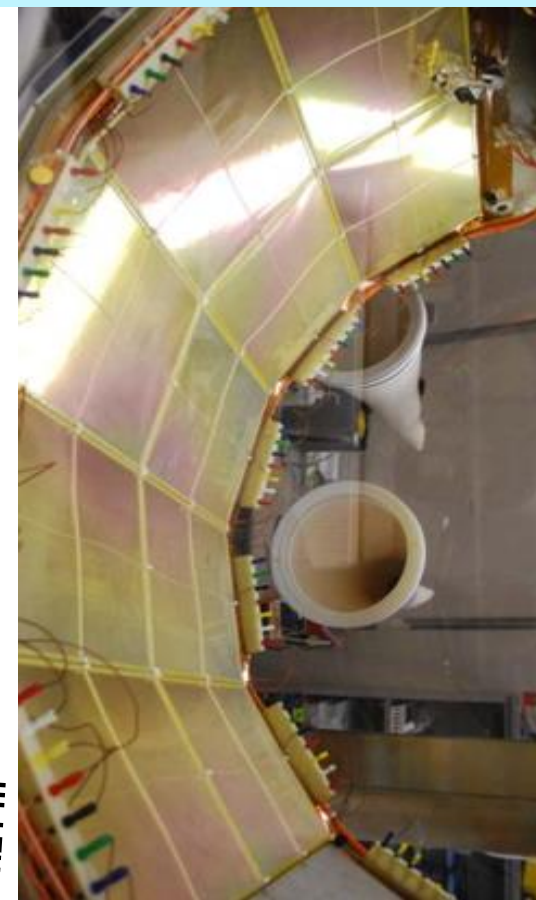
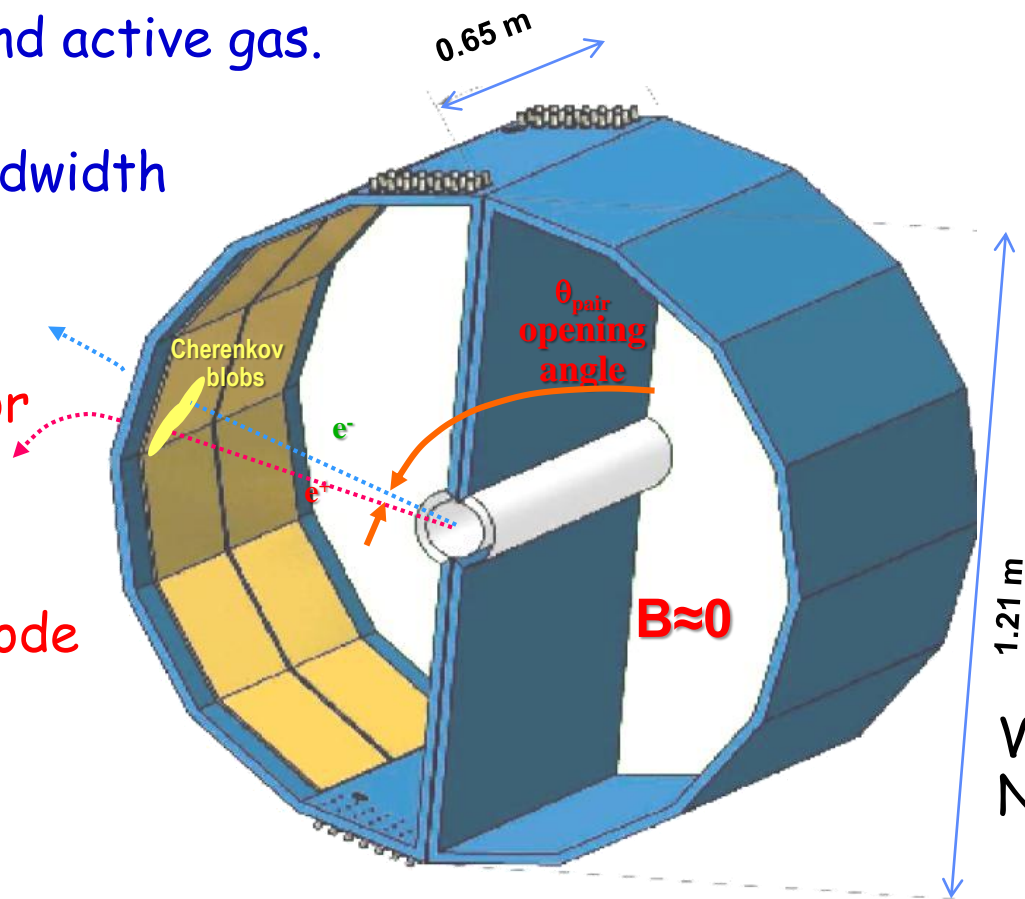
R. Chechik et al., NIM A 419 (1998) 423



A. Breskin and R. Chechik, NIM A 595 (2008) 116

# HBD- Cherenkov detector with GEMs +CsI

- ✓ Proximity focus configuration, no window, no mirror
- ✓  $\text{CF}_4$  radiator and active gas.  $L_{\text{rad}}=50$  cm
- ✓ Very large bandwidth 108 - 200 nm (6.2 - 11.5 eV)
- ✓ triple GEMs for signal multiplication
- ✓ CsI photocathode

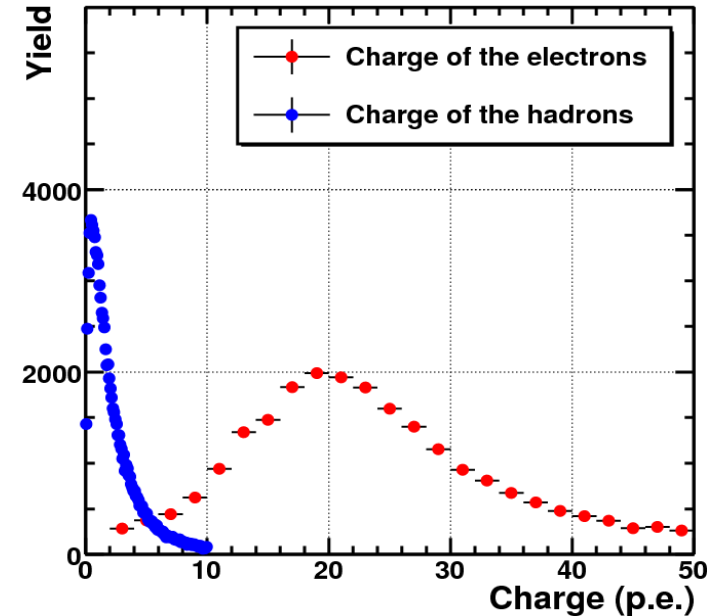
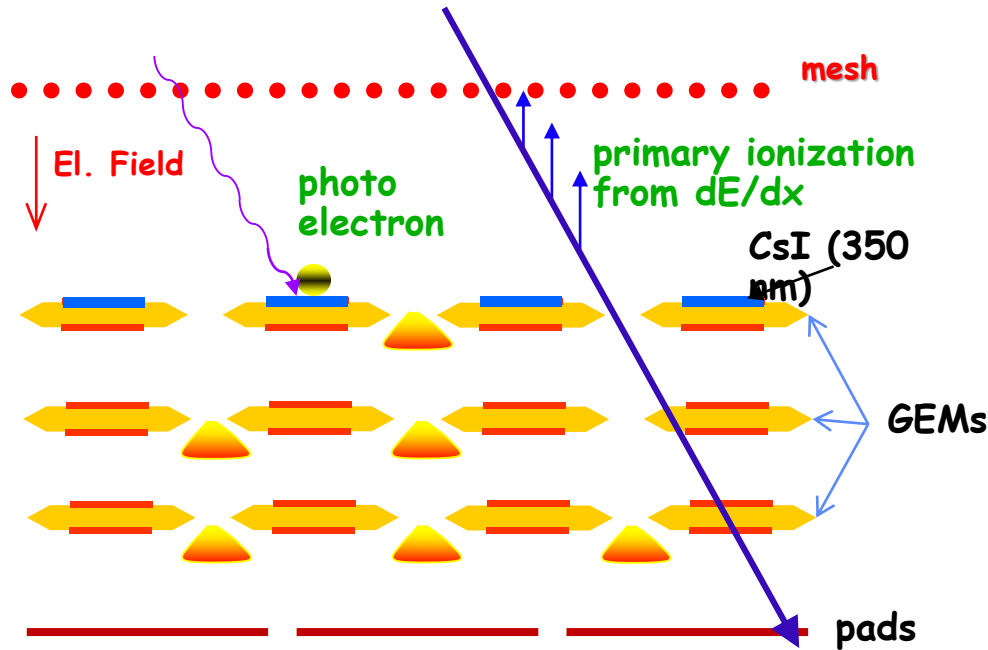


W. Anderson et al.,  
NIM A 646 (2011) 35



# HBD - hadron blindness

❖ Electron signals are relatively rare (compared to hadrons)



- a. Detector operated in reverse bias mode to repel the ionization charge from  $dE/dx$
- b. Cherenkov light is formed only by  $e^+$  or  $e^-$
- c. Successful operation at PHENIX since several years
- d. It is not a detector of single photons

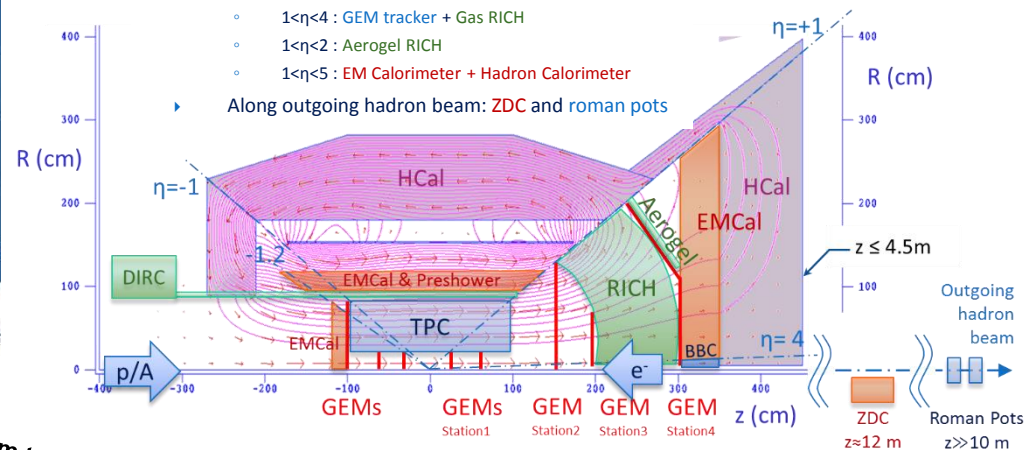


# a PID for EIC proposal

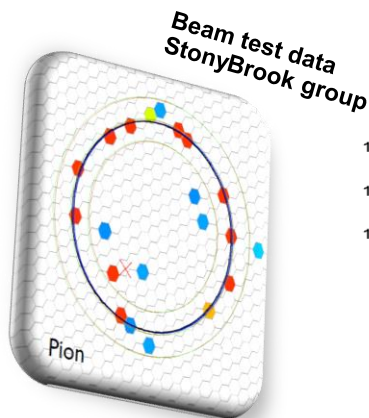
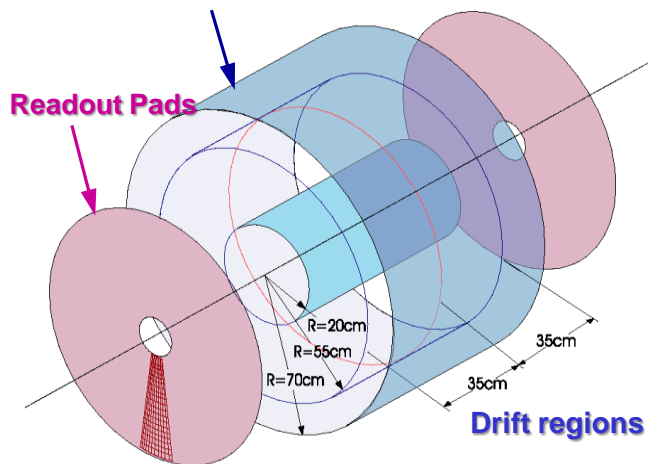
Current PHENIX	sPHENIX (+fsPHENIX)	An EIC detector
<ul style="list-style-type: none"> <li>15 years of operation</li> <li>Produced broad spectrum of studies of QGP and Hadron Physics</li> <li>140+ published papers to date</li> <li>Last run in 2016</li> </ul>	<ul style="list-style-type: none"> <li>Comprehensive upgrade based on the BaBar magnet</li> <li>Goal is to do a systematic study the QGP near <math>T_c</math> by measuring jets and heavy quarkonia</li> <li>Possible addition of a forward spectrometer (fsPHENIX) to continue study of Spin and CNM</li> </ul>	<ul style="list-style-type: none"> <li>Further upgrades of sPHENIX leads to a capable Day 1 detector for eRHIC</li> <li>Study polarized ep and eA physics</li> <li>Large coverage of tracking, calorimetry and PID</li> </ul>
~2000	2017→2020	~2025
RHIC: A+A, spin-polarized p+p, spin-polarized p+A		EIC: e+p, e+A

## sPHENIX → eRHIC Detector LOI: arXiv:1402.1209

- 1< $\eta$ <+1 (barrel) : sPHENIX + Compact-TPC + DIRC
- 4< $\eta$ <-1 (e-going) : High resolution EM calorimeter + GEM trackers
- +1< $\eta$ <+4 (h-going) :
  - 1< $\eta$ <4 : GEM tracker + Gas RICH
  - 1< $\eta$ <2 : Aerogel RICH
  - 1< $\eta$ <5 : EM Calorimeter + Hadron Calorimeter
- Along outgoing hadron beam: ZDC and roman pots

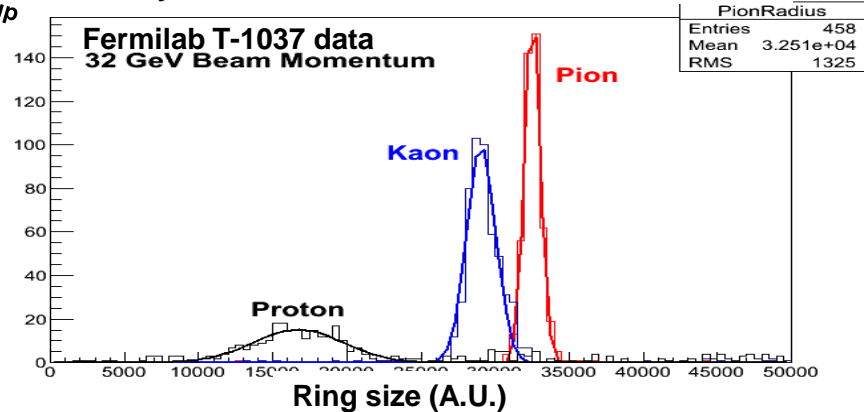


### GEMs with CsI Readout Plane



M.Blatnik et al. arXiv:1501.03530

### Courtesy : EIC RD6 TRACKING & PID CONSORTIUM





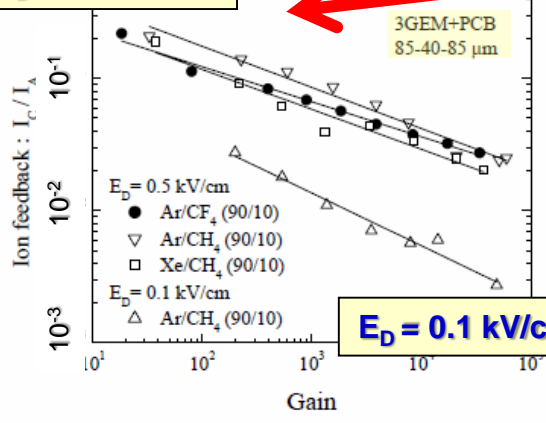
# GEM-based PDs and IBF

rich literature about IFB in GEM-based detectors  
 here examples with semi-transparent PC

- strong dependence from gain and  $E_{D\text{DRIFT}}$
- poor dependence from pressure and gas type

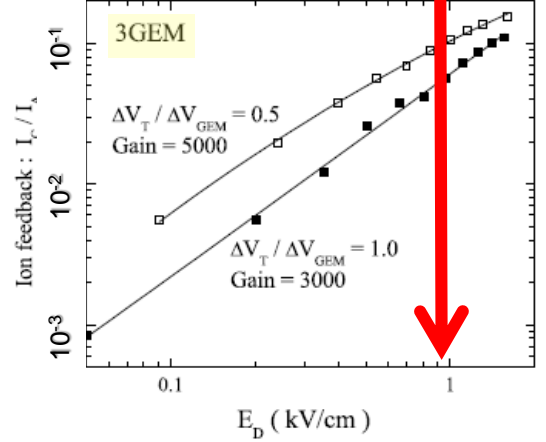
$E \sim 1 \text{ kV/cm}$  needed for good photoelectron extraction

$E_D = 0.5 \text{ kV/cm}$

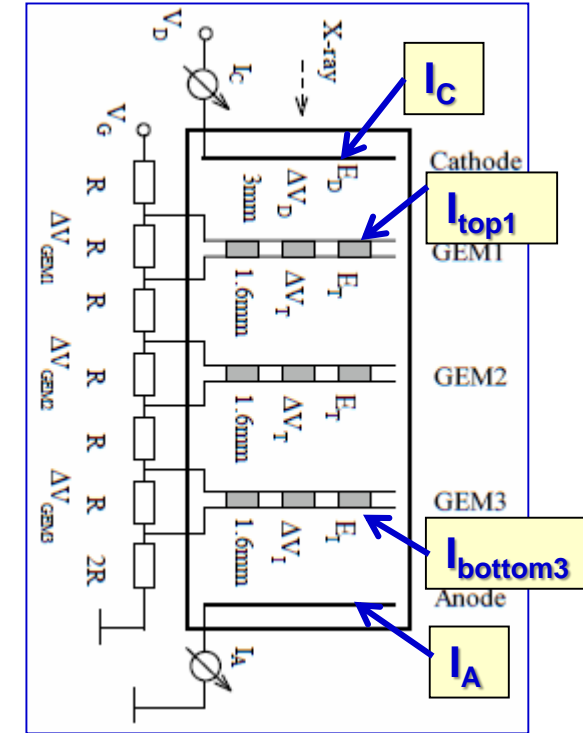


$E_D = 0.1 \text{ kV/cm}$

A. Bondar et al., NIMA 496 (2003) 325



A. Breskin et al., NIMA 478 (2002) 225d



The same for reflective PCs :  
 small and reversed  $E_D$  is needed

**IBF: a few % level in effective GEM-based photon detectors**



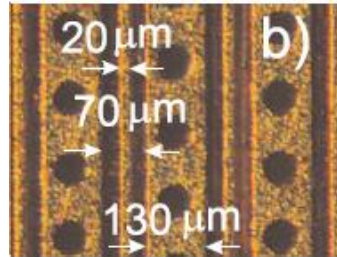
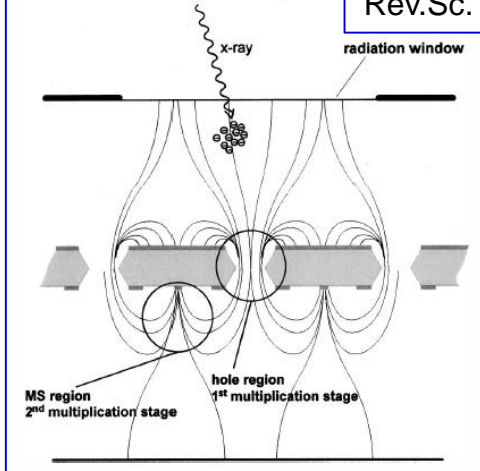
# OVERCOMING IBF

More complex geometries needed with extra electrodes to trap the ions:

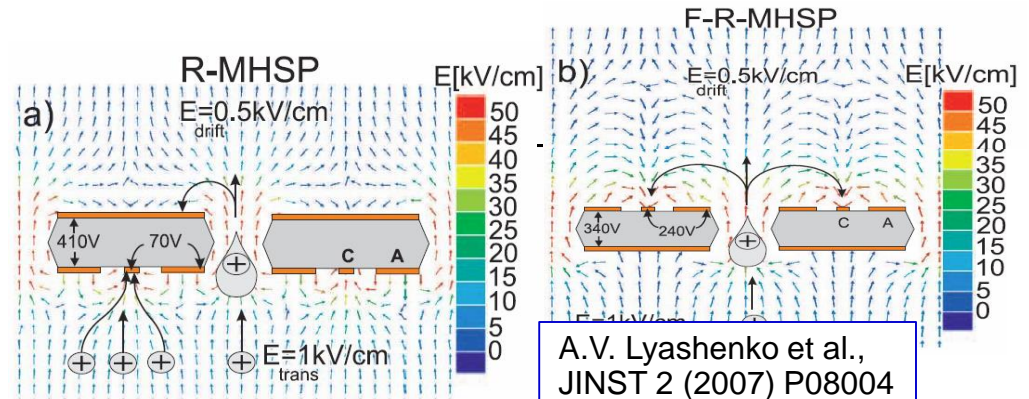
**MHSP**

X-Ray detector

J.F.C.A. Veloso et al.,  
Rev.Sc. Instr. 71 (2000) 2371

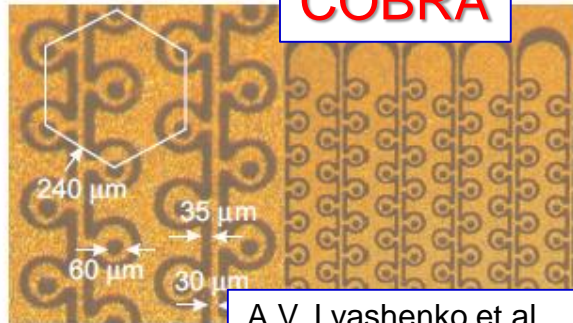


Micro-Hole & Strip Plate (**MHSP**), **COBRA**

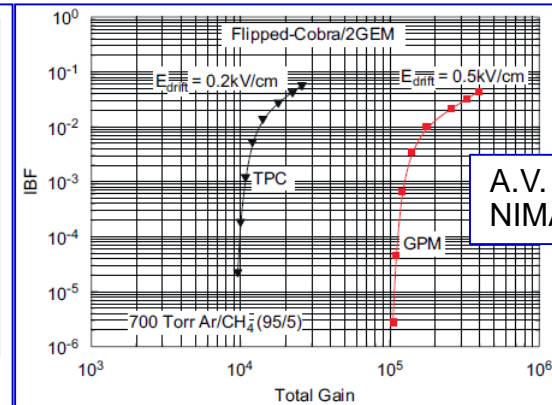
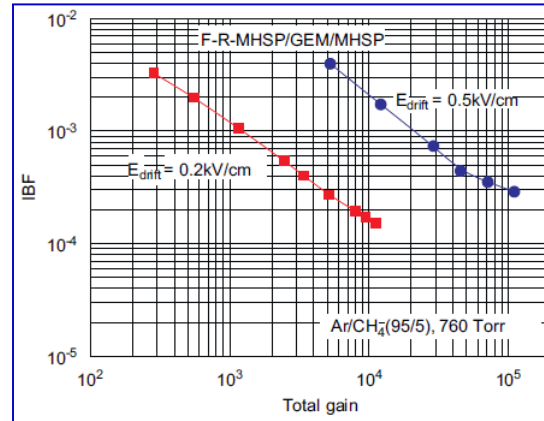


A.V. Lyashenko et al.,  
JINST 2 (2007) P08004

**COBRA**



A.V. Lyashenko et al.,  
NIMA 598 (2009) 116



A.V. Lyashenko et al.,  
NIMA 598 (2009) 116



# GEM-based PDs and GAIN

## LARGE GAIN RELEVANT FOR SINGLE PHOTON DETECTION

- **GEM-based PDs in laboratory studies**
    - for single photoelectron detection, they have been operated at gains  $> 10^5$  (see, for instance, the plots of the previous slides)
  
  - **GEM-based detectors in experiments**
    - Always a MIP flux and small rates of heavily ionizing fragments crossing the detectors (even when the detectors are used as photon detectors)
      - At COMPASS:  $G \sim 8000$  (B. Ketzer, private comm.)
      - At LHCb:  $G \sim 4000$  (M. Alfonsi NIMA 581 (2007) 283)
      - At TOTEM:  $G \sim 8000$  (G. Catanesi, private comm.)
      - Phenix HBD:  $G \sim 4000$  (W. Anderson et al., NIMA 646 (2011) 35)
- **In experiments, small chances**  
to operate GEM-based PDs at gains  $> 10^4$





# THGEM-based PDs, why ?

## PCB technology, thus:

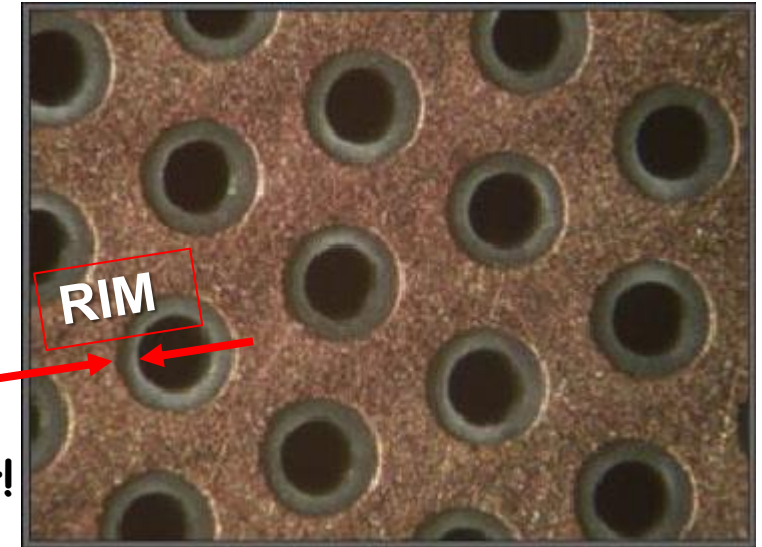
- robust
- mechanically self supporting
- industrial production of large size boards
- large gains have been immediately reported (**rim** !)

## Comparing to GEMs

- Geometrical dimensions  $\times \sim 10$ 
  - But  $e^-$  motion/multiplic. properties do not!
  - Larger holes:
    - dipole fields and external fields are strongly coupled
    - $e^-$  dispersion plays a minor role

## About PCB geometrical dimensions:

Hole diameter :	0.2 - 1 mm
Pitch :	0.5 - 5 mm
Thickness :	0.2 - 3 mm



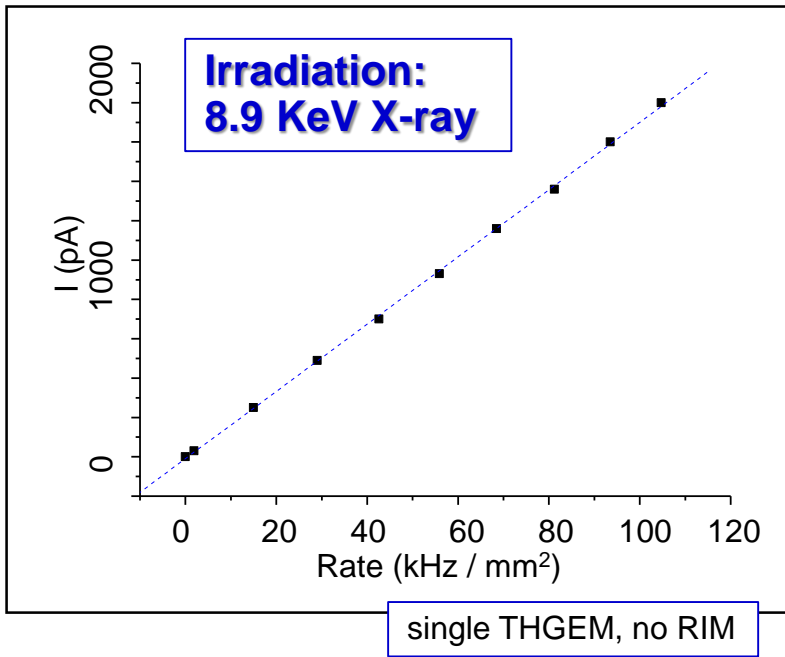
## introduced in // by different groups:

- L. Periale et al., NIM A478 (2002) 377.
- P. Jeanneret, PhD thesis, Neuchatel U., 2001.
- P.S. Barbeau et al, IEEE NS50 (2003) 1285
- R. Chechik et al, .NIMA 535 (2004) 303



# THGEM rate capability and IBF

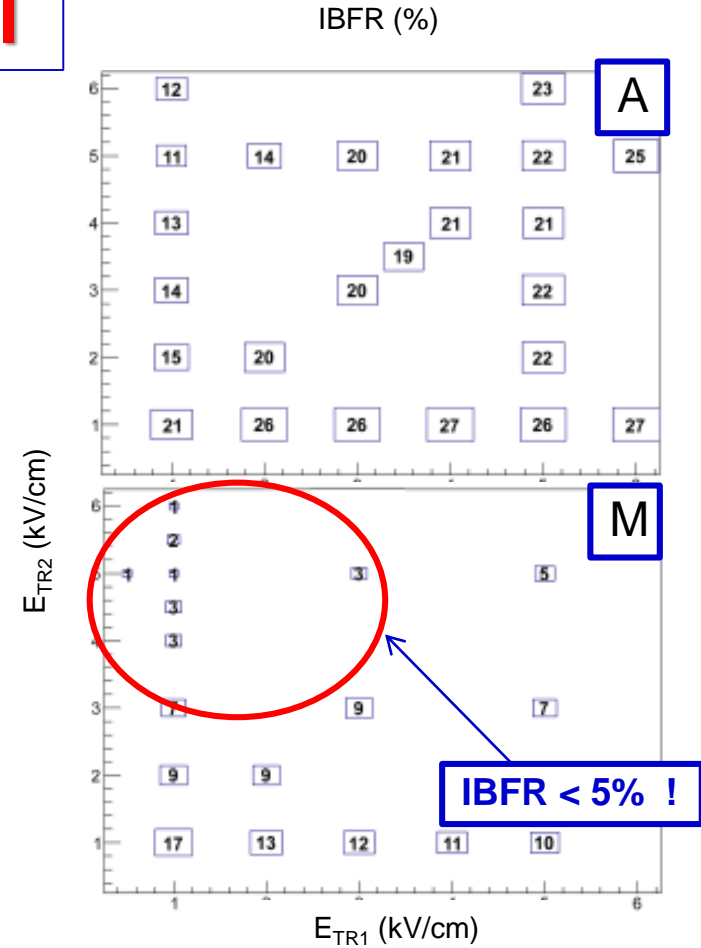
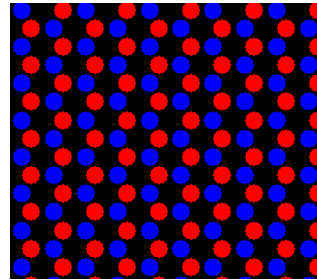
## High rate device



M. Alexeev et al. JINST 10 (2015) P03026  
The gain in Thick GEM multipliers and its time-evolution

## IBF control

Triple THGEM:  
Ion Back Flow  
reduction by  
staggering plates



M. Alexeev et al., JINST 7 (2012) C002014

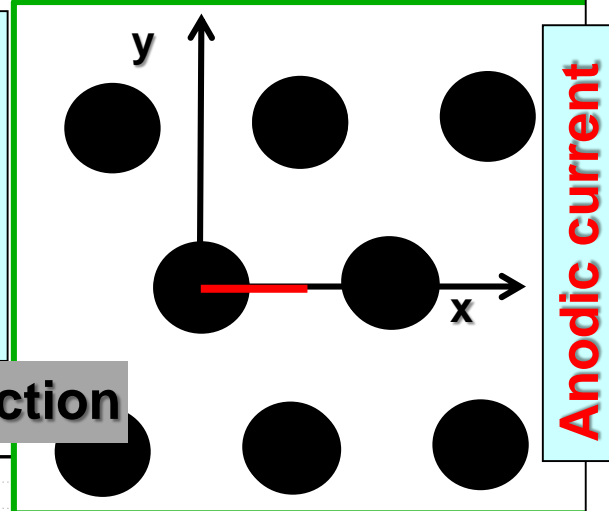


# PHOTOELECTRON EXTRACTION

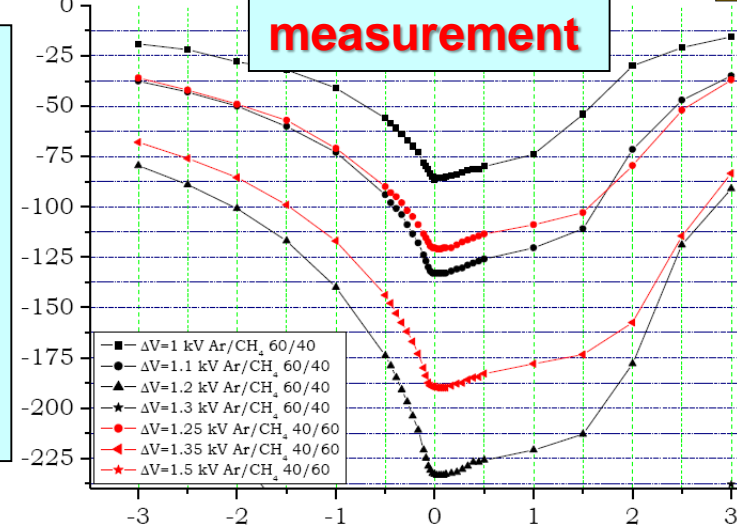
photoelectron trajectories from a THGEM photocathode, simulation, multiplication switched off

thickness 0.6 mm, diam. 0.4 mm, pitch: 0.8 mm,  $\Delta V = 1500$  V

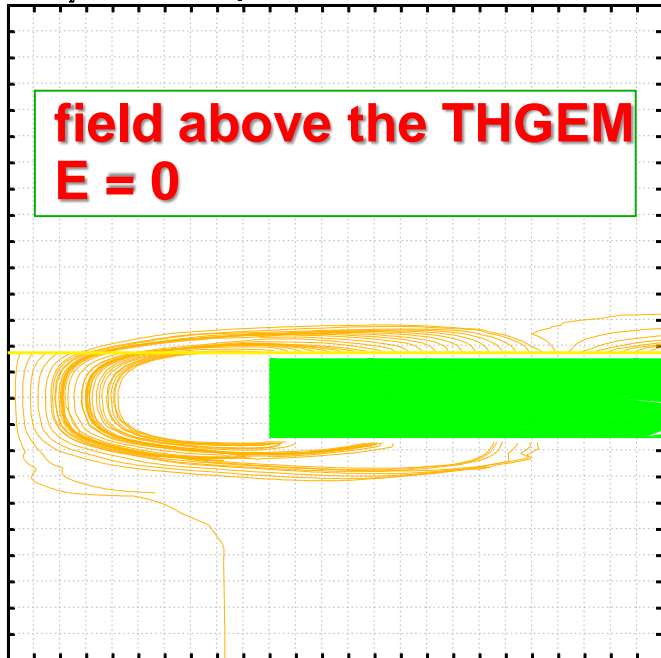
x cross-section



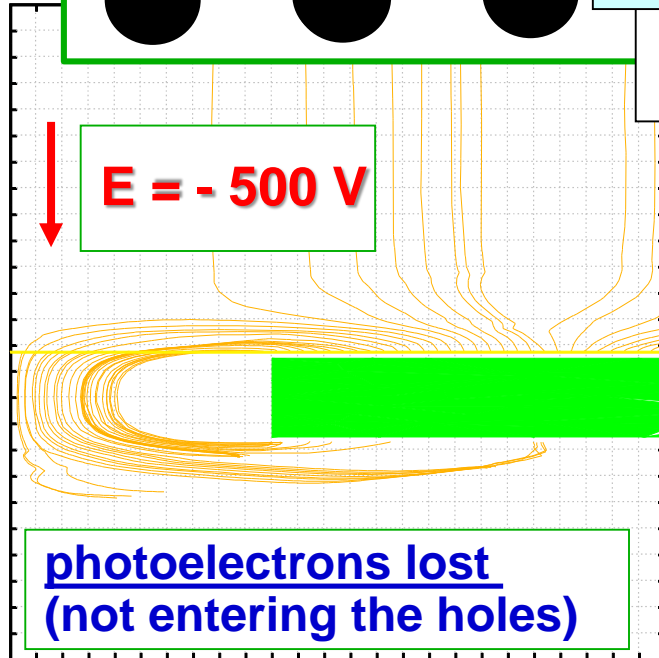
Anodic current



external electric field applied

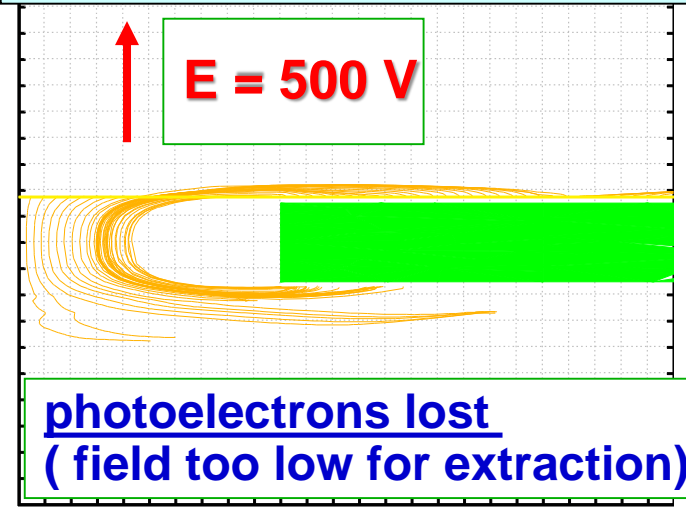


field above the THGEM  $E = 0$



$E = -500$  V

photoelectrons lost (not entering the holes)



$E = 500$  V

photoelectrons lost (field too low for extraction)



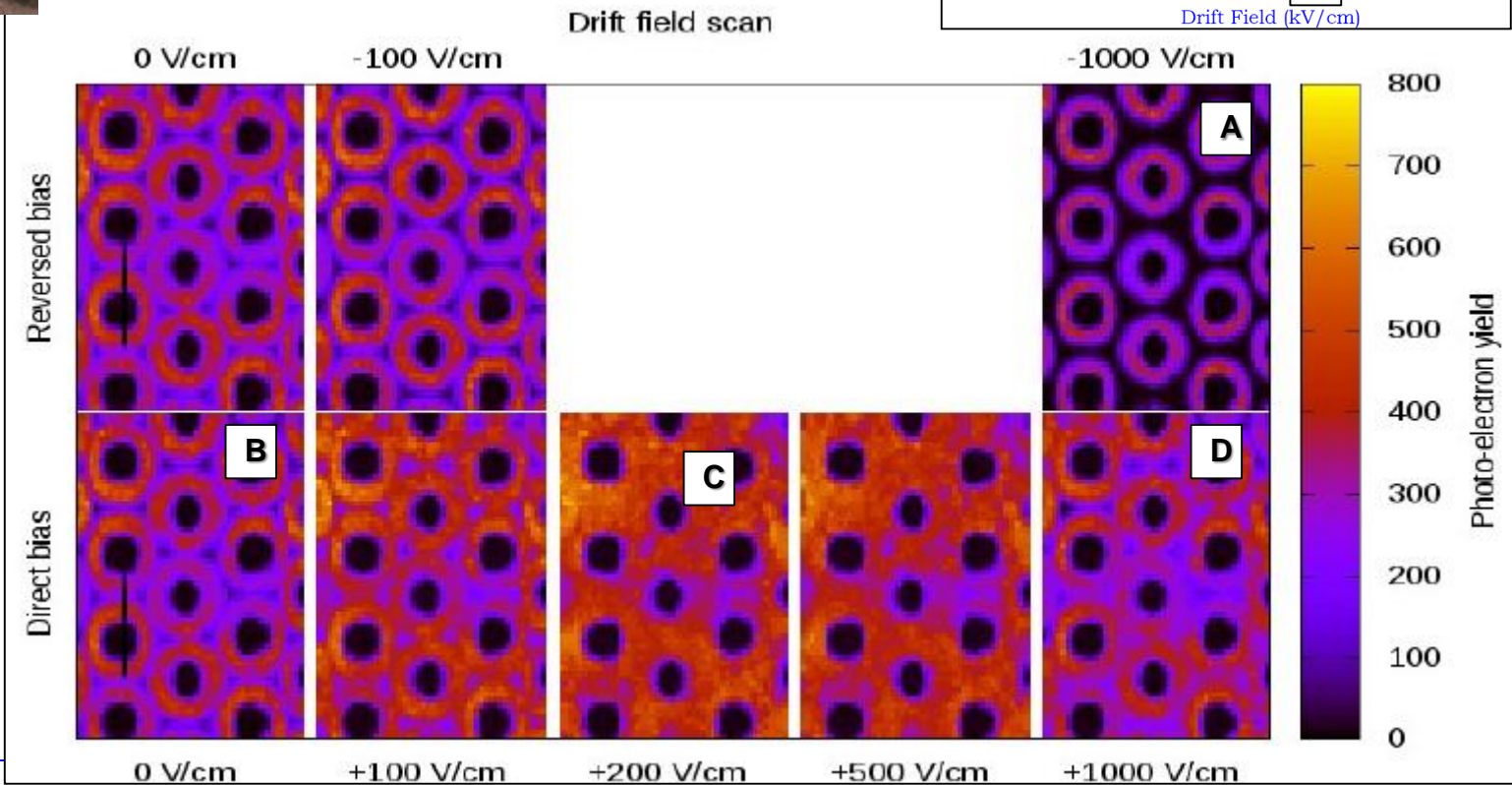
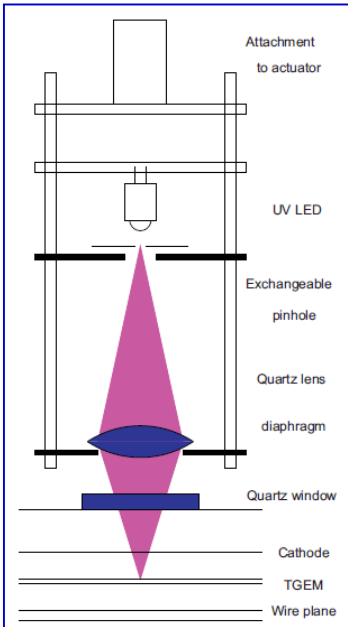
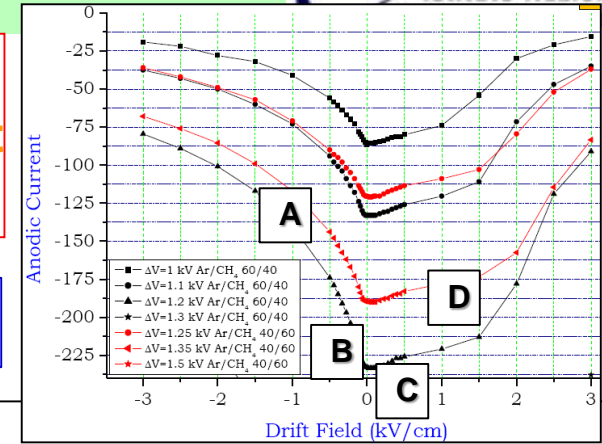
# LEOPARD



Photoelectron extraction from THGEM PC fully confirmed by direct observation with "Leopard"

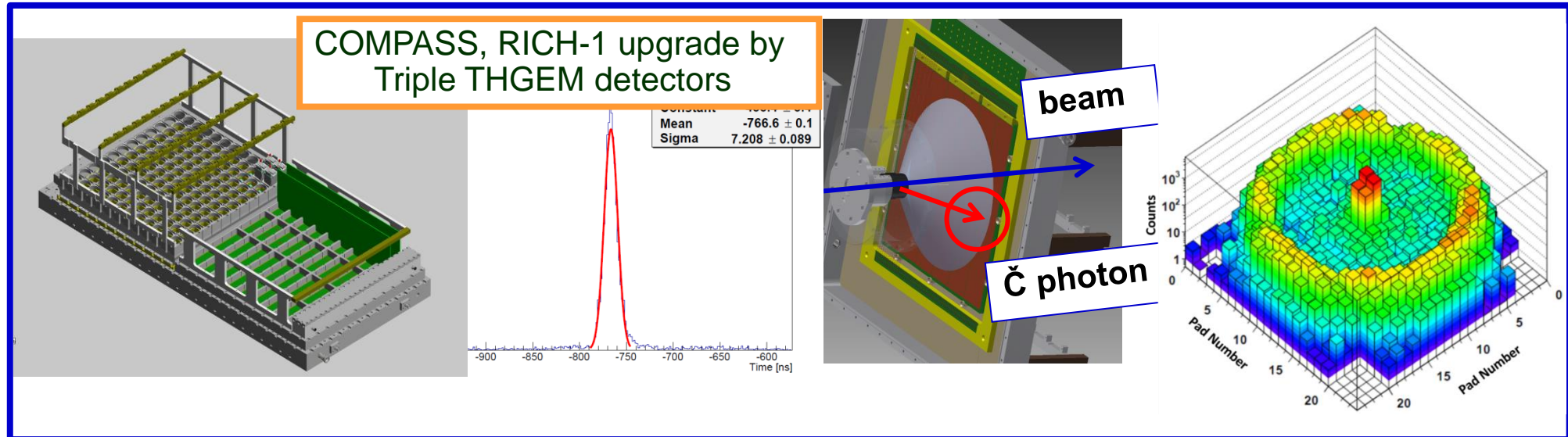
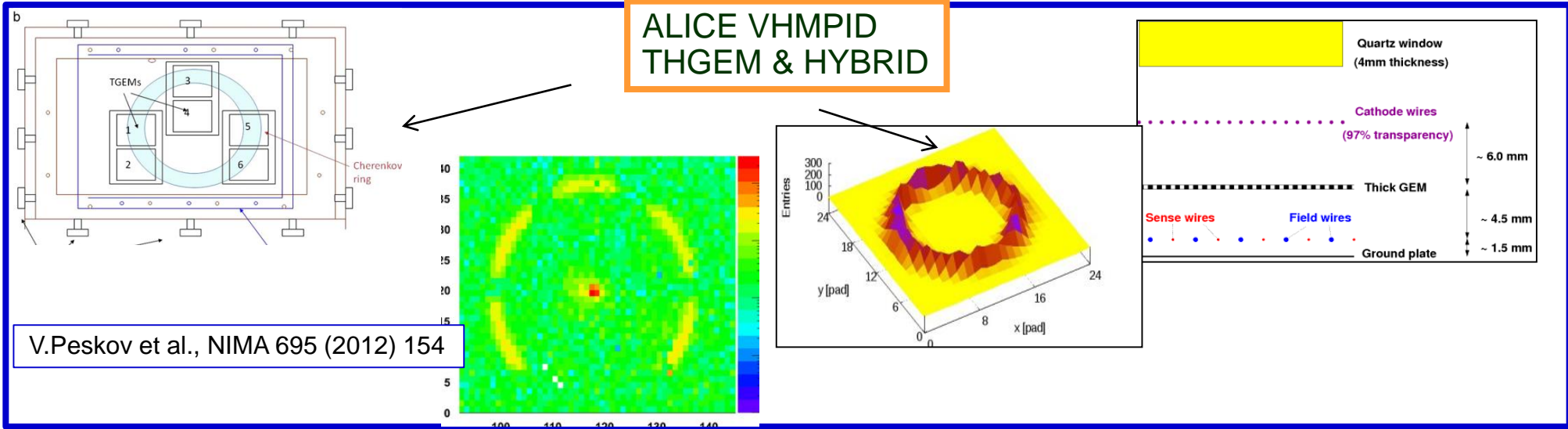
G.Hamar talk on Thursday

G.Hamar and D.Varga, NIMA 694(2012) 16





# THGEM R&D for RICHes





# NUMBER OF DETECTED PHOTONS

V.Peskov et al., NIMA 695 (2012) 154

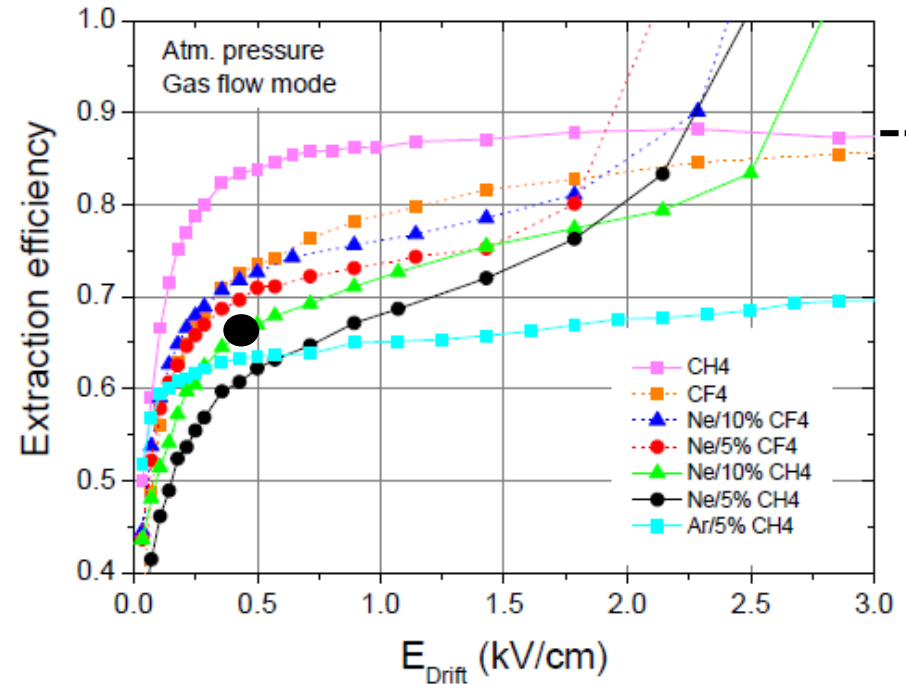
**N of detected photons is ~60-70% of MWPCs with CsI**

- **Ne+10%CH4, used with  $\Delta V$  at 650-750 V**

## 5. Conclusions and Outlook

We report the first successful implementation of a set of CsI-TGEMs with a liquid radiator where a Cherenkov ring has been observed. The results obtained are encouraging and suggest that the present performance could be improved in the future by optimizing elements of the design. We are launching now systematic studies on TGEM geometry optimization allowing increasing the value of  $\eta_{rel}$ ,  $\epsilon_{col}$  and  $A_{eff}$ . We also are planning to investigate

→ **Relative extraction efficiency**  
**Respect to pure methane at**  
 **$E \sim 7\text{kV/cm} \sim 75\%$**



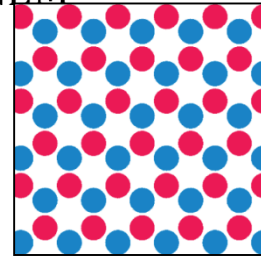
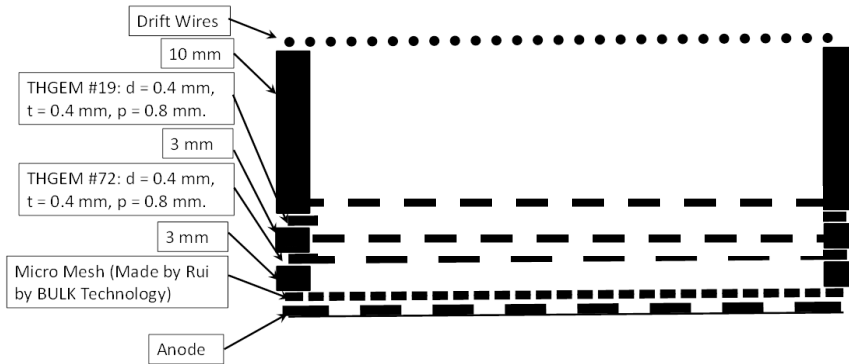
C. D. R. Azevedo et al., 2010 JINST 5 P01002



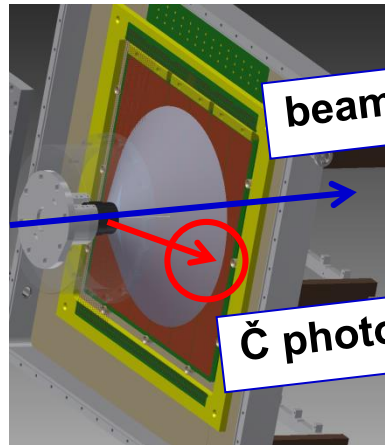
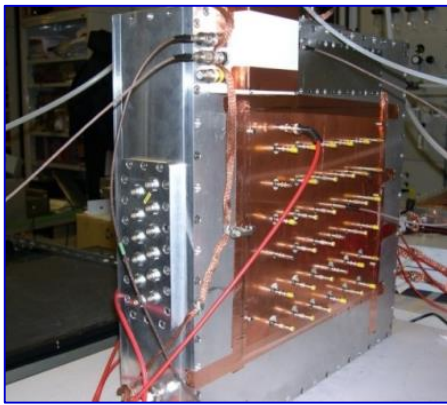
# HYBRID MPGD PDs (THGEM + MM)

- The 1<sup>st</sup> THGEM forms the PC
- The 2<sup>nd</sup> THGEM (staggered) forces the electron diffusion
- The MM provides large gain, made larger by the diffusing the impinging electron cloud

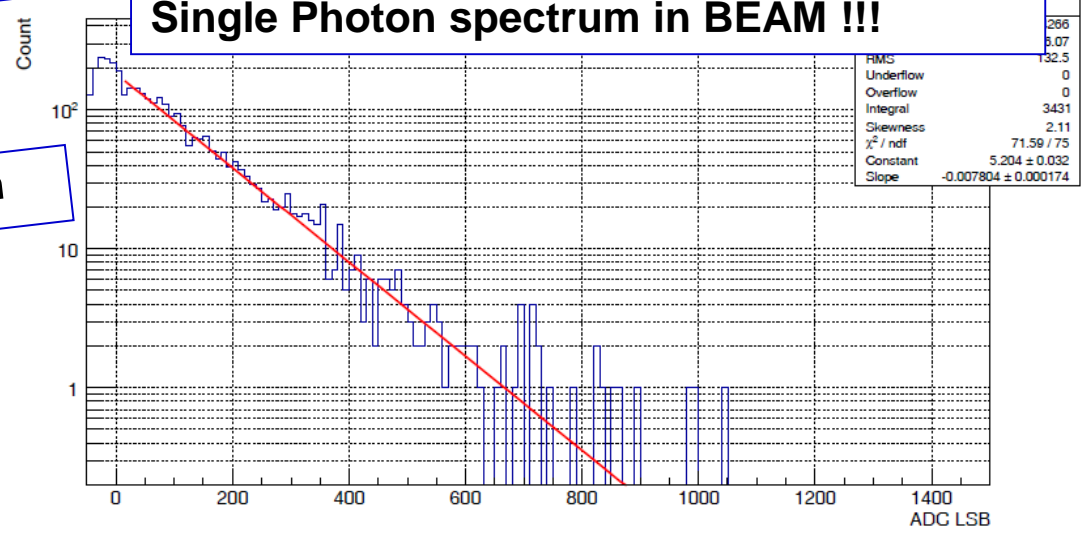
Setup Hybrid (300 X 300 mm<sup>2</sup>) with double THGEM



THGEM staggered !



Gain ~ 130K  
Single Photon spectrum in BEAM !!!

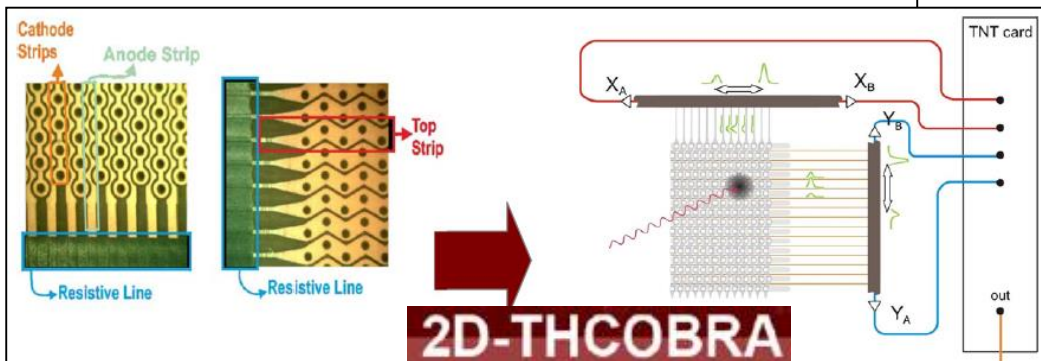
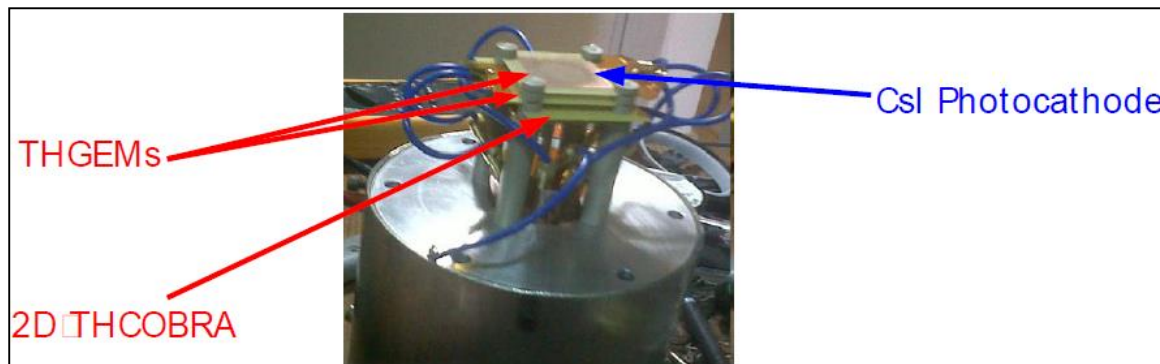


The same architecture independently studied in parallel as GPM for DM searches (see later)



# HYBRID MPGD PDs (THGEM + THCOBRA)

- 2 THGEMs
- a THCOBRA with 2 d R-O structure



Parameters			
Structure	Hole Diameter ( $\mu\text{m}$ )	Pitch ( $\mu\text{m}$ )	RIM ( $\mu\text{m}$ )
THGEM 1	400	800	5
THGEM 2	700	1300	100
2D THCOBRA	400	1000	80

## Gas Photomultiplier (GPM) : 2D-THCOBRA

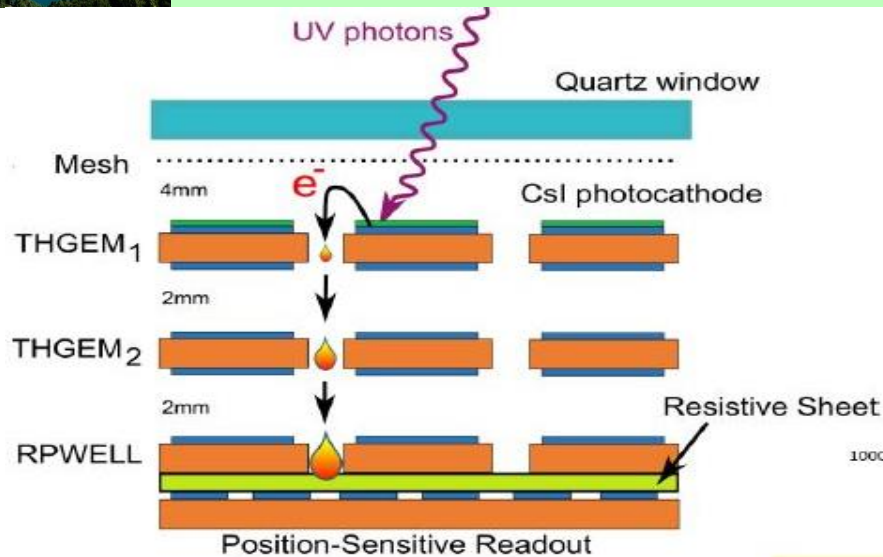
- Good Performance
  - Gain of  $10^6$
  - IBF values of about: 20%
- 2D THCOBRA adequate to obtain image
- Position Resolution: FWHM=  $300 \mu\text{m}$ ,  $\sigma = 128 \mu\text{m}$
- Count rate of 100kHz

T. Lopes 2013 JINST 8 P09002



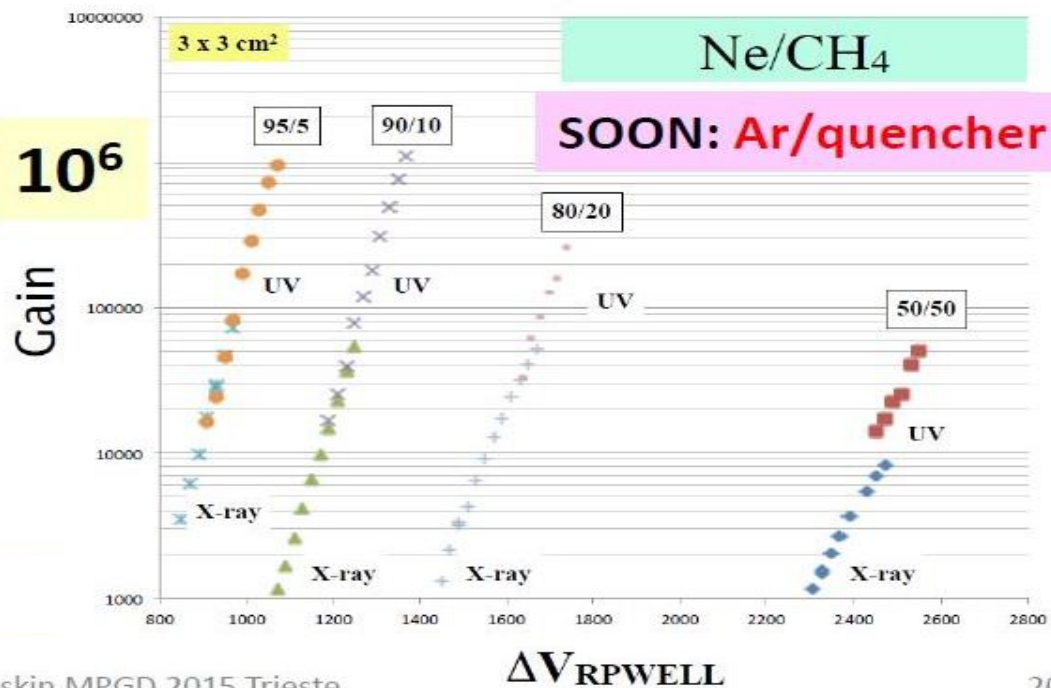


# RPWELL



## R&D @ Weizmann

- RT & noble-liquid T
- **RPWELL vs nTHGEMs+RPWELL**
- **Gases**
- **stability**



A Breskin MPGD 2015 Trieste

$\Delta V_{RPWELL}$

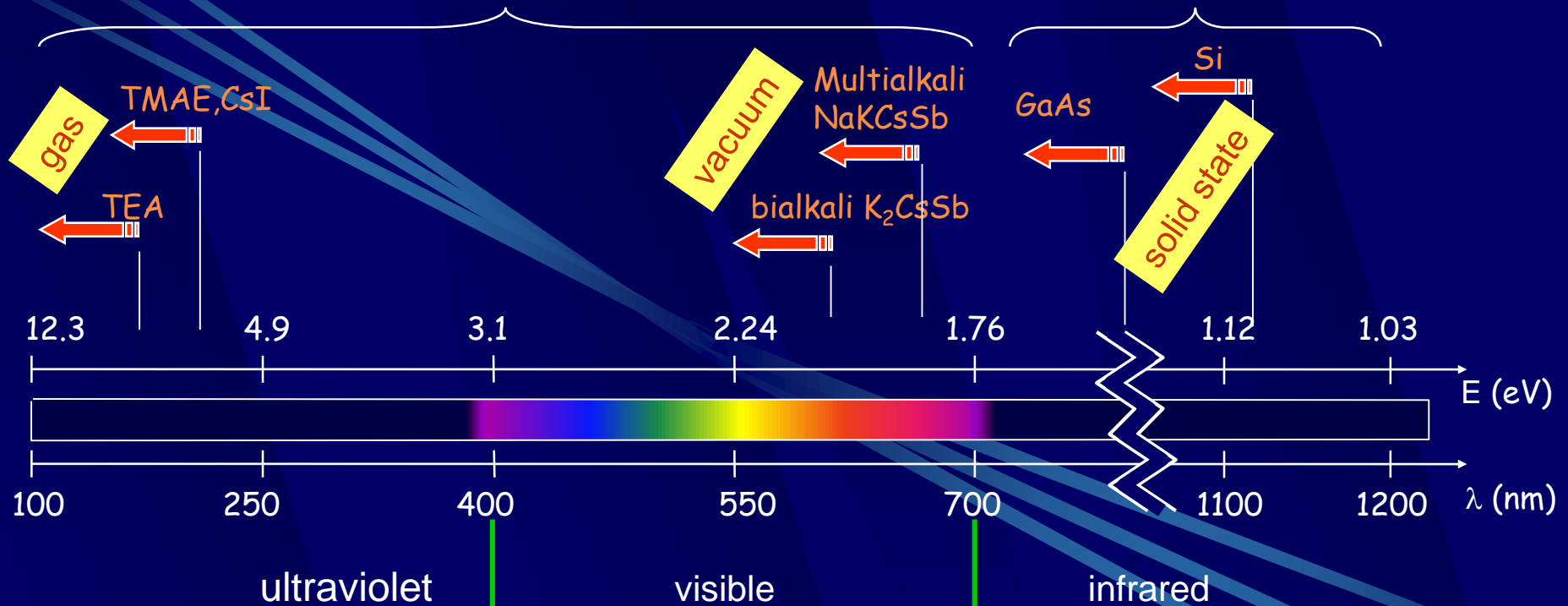
20

# Different photocathodes and their thresholds

$$\varepsilon_Q = \text{Q.E.} = N_e / N_\gamma$$

external photoeffect  
 $\varepsilon_Q < 50\%$

internal or external photoeffect  
 $\varepsilon_Q > 50\%$



- Photon detection involves often materials like K, Na, Rb, Cs (alkali metals). They have the smallest electronegativity  $\rightarrow$  highest tendency to release electrons.
- Most photocathodes are VERY reactive; Exceptions: Si and CsI.



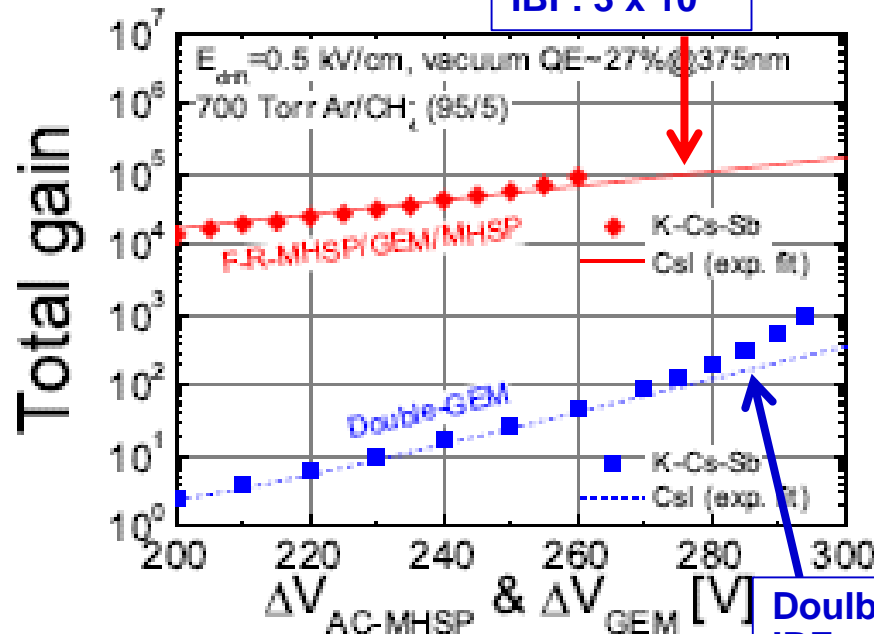
# PHOTOCATHODES FOR VISIBLE LIGHT

- Chemical reactivity (gas purity better than ppm level needed → UHV materials and sealed detectors)
- PC stability under ion bombardment - work function lower than CsI one
- **AGEING** CsI: -16% QE at  $25\mu\text{C}/\text{mm}^2$   
Bilkaly: -20% QE at  $0.4\mu\text{C}/\text{mm}^2$

F.Tokanai et al., NIMA 628 (2011) 190  
T.Moriya et al., NIMA 732 (2013) 263

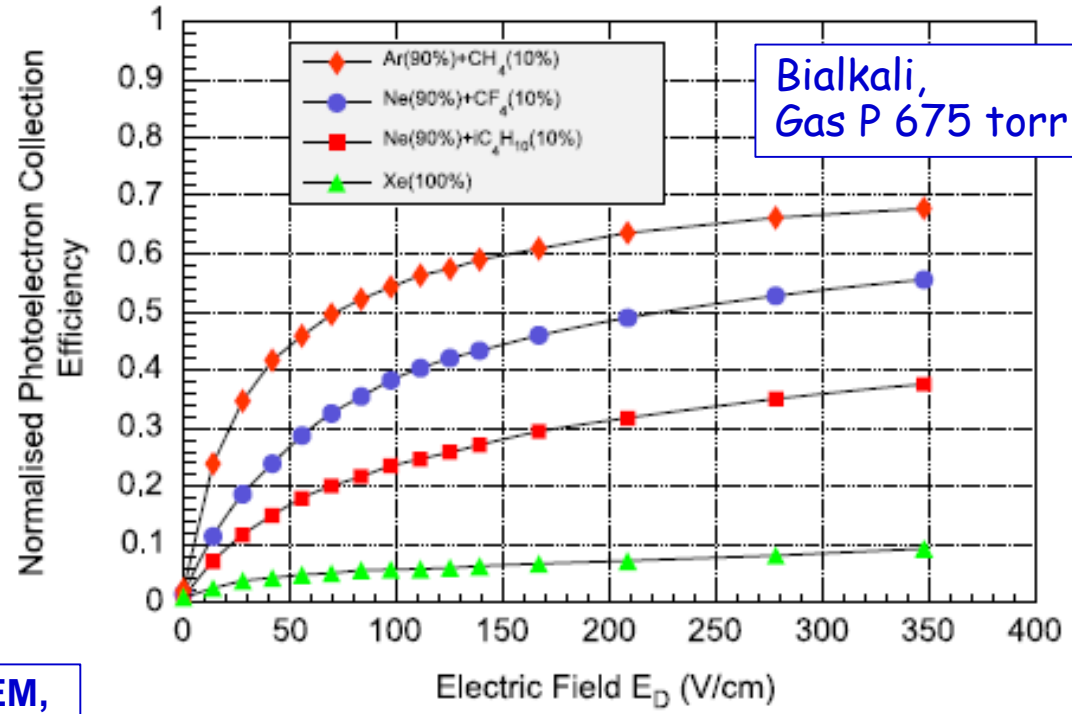
## K-Cs-Sb vs CsI

F-R-MHSP,  
IBF:  $3 \times 10^{-4}$



Double GEM,  
IBF:  $\sim 10^{-2}$

A.V.Lyashenko et al., 2009 JINST 4 P07005



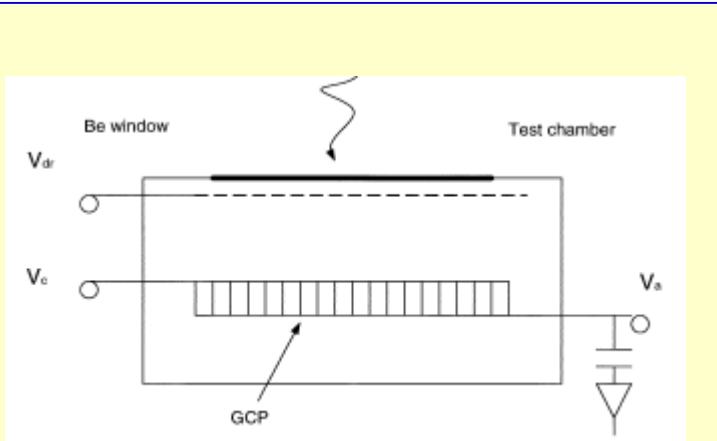
Bialkali,  
Gas P 675 torr

F. Tokanai et al., NIMA 610 (2009) 164

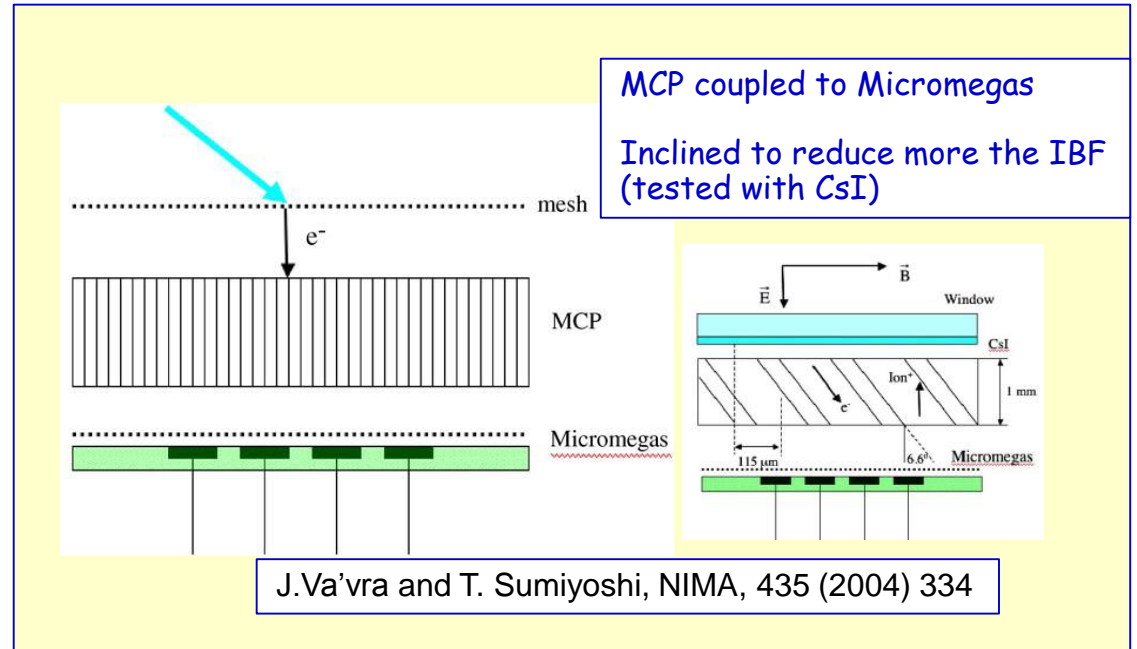


# FIRST GASEOUS DETECTORS FOR VISIBLE LIGHT

## The Capillary Plate (CP) approach



V. Peskov et al., NIMA 433 (1999) 492



MCP coupled to Micromegas

Inclined to reduce more the IBF (tested with CsI)

J.Va'vra and T. Sumiyoshi, NIMA, 435 (2004) 334



# GASEOUS DETECTORS FOR VISIBLE LIGHT

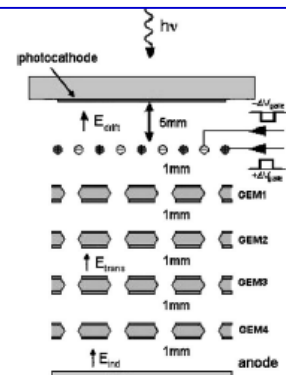
## the GEM approach

Multiple GEM sealed



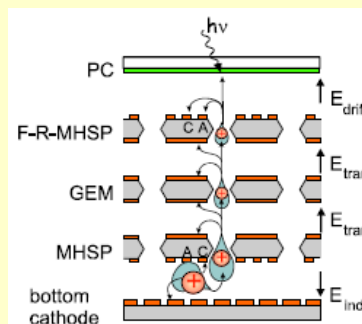
R.Chechik et al., NIMA 502 (2003) 195

Pulsed ion gating



A. Breskin et al.,  
B. NIMA 553 (2005) 46

K-Cs-Sb - Continuous mode, not a sealed PD



A.V.Lyashenko et al.,  
2009 JINST 4 P07005

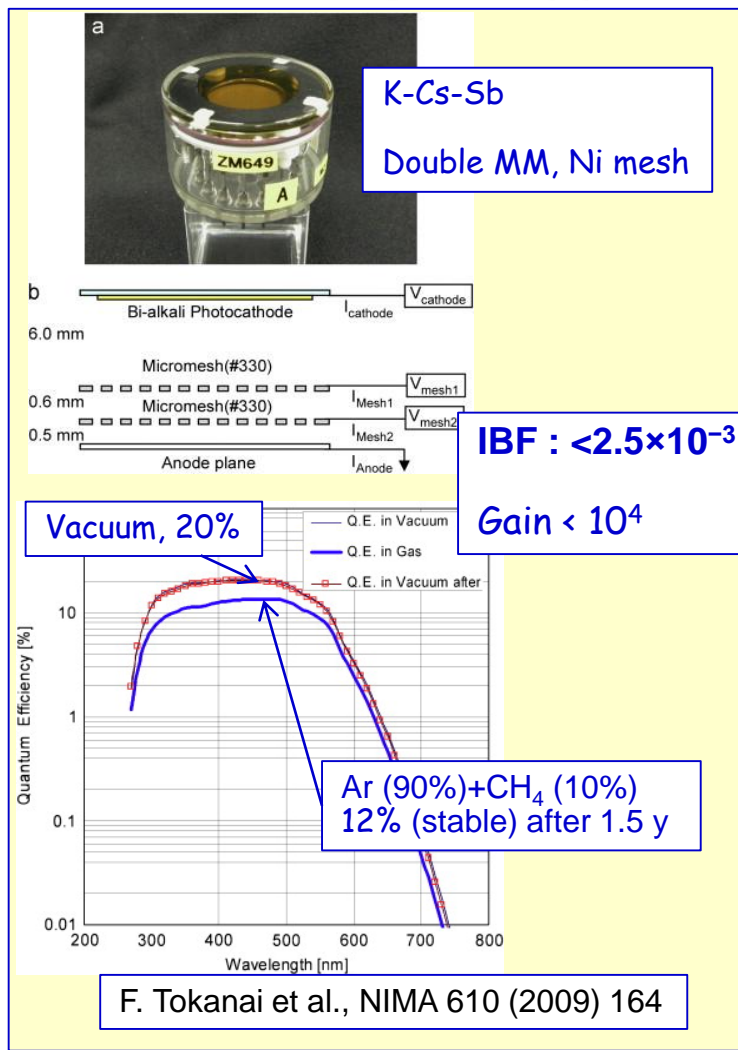
### Poor compatibility of alkali and GEM material ?

Extremely poor QE of the alkali PC:  
**the material of the GEM chemically reacts with the alkali metals**

F. Tokanai et al., NIMA 610 (2009) 164

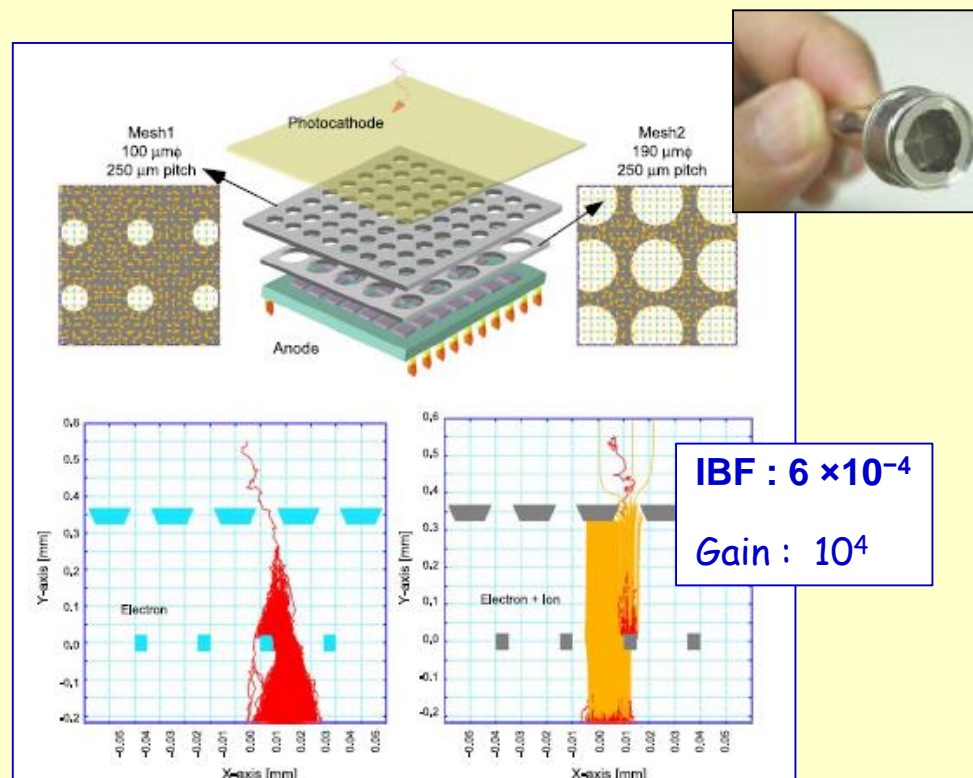
# GASEOUS DETECTORS FOR VISIBLE LIGHT

## the MicroMegas approach



2 staggered MM layers to enhance ion trapping

In collaboration with HAMAMATSU



F. Tokanai et al., NIMA 766 (2014) 176

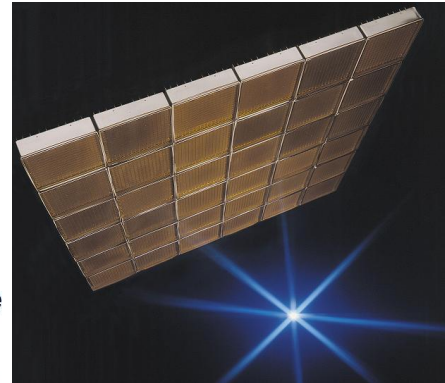
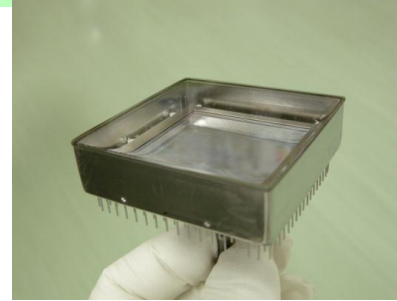


# GASEOUS DETECTORS FOR VISIBLE LIGHT

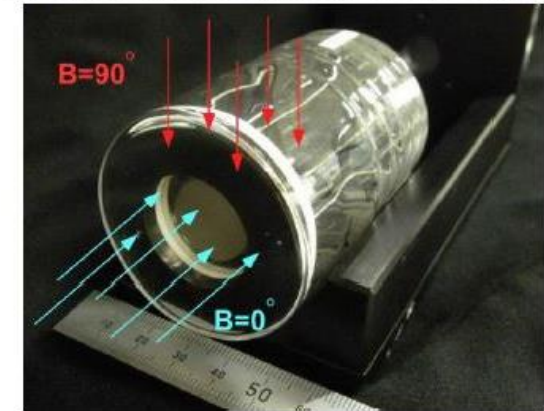
## Gaseous PMT

Yamagata U. TMU, HAMAMATSU

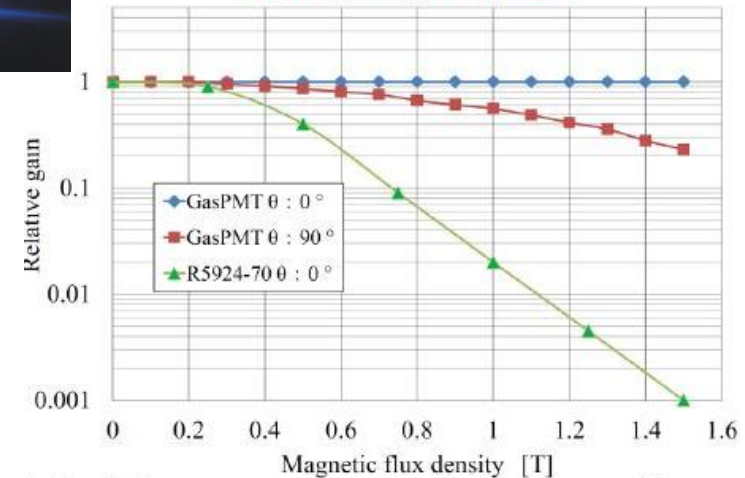
Sensor type	Sensitivity	Position Resolution	Timing Resolution	Uniformity	Price	Magnetic Field	Effective Area
Vacuum PMT	⊙	△	⊙	△	○	△	○
CCD / CMOS	△	⊙	×	⊙	△	⊙	×
Gaseous PMT	○	○	○	○	⊙	⊙	⊙



Operation in magnetic field environmen

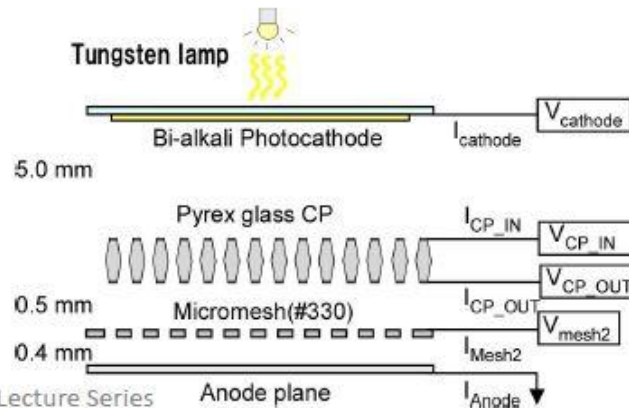


Ar(90%)+CH<sub>4</sub>(10%) 1気圧



H. Sugiyama et al., NIMA (2016) in press

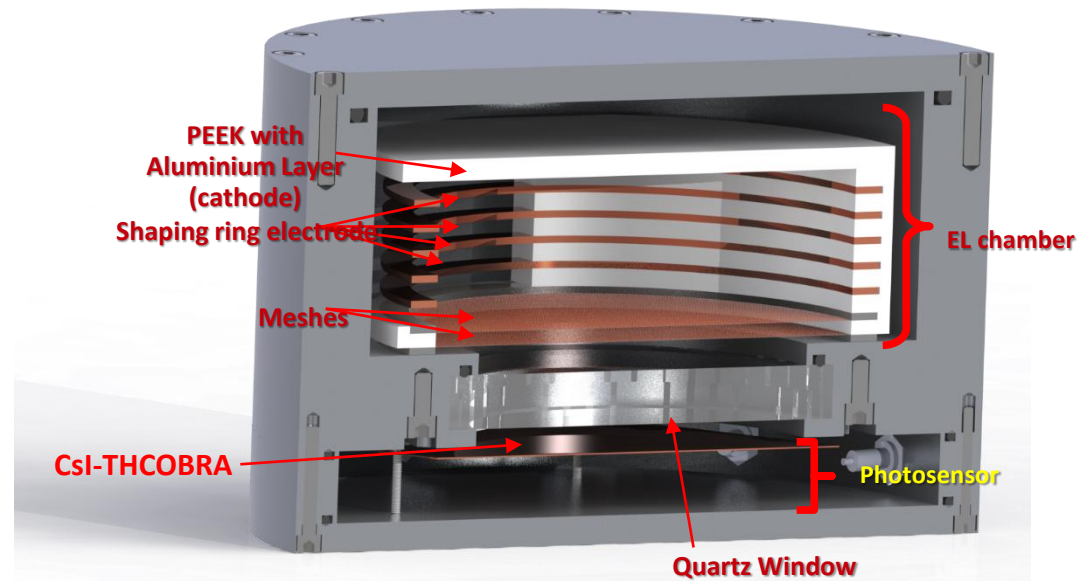
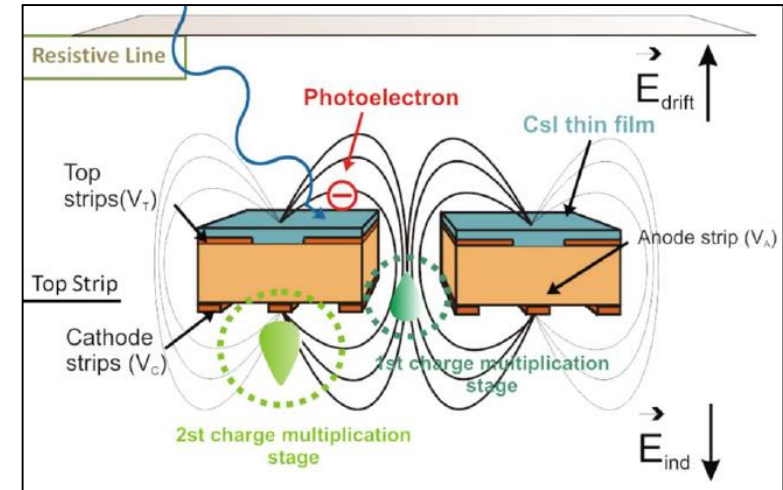
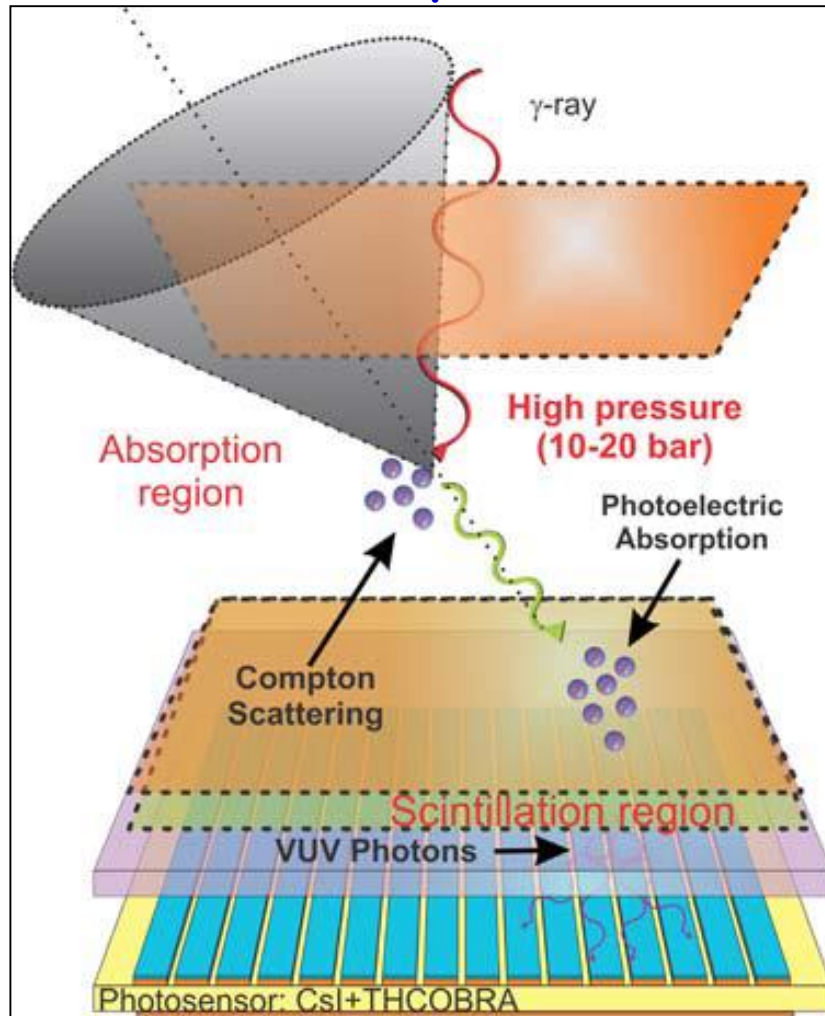
- The advantage of the **gaseous PMT**:
  - It can achieve a **very large effective area** with moderate **position** and **timing** resolutions.
  - It can be easily operated under a **very high magnetic field**.





# Gaseous Compton camera for medical imaging

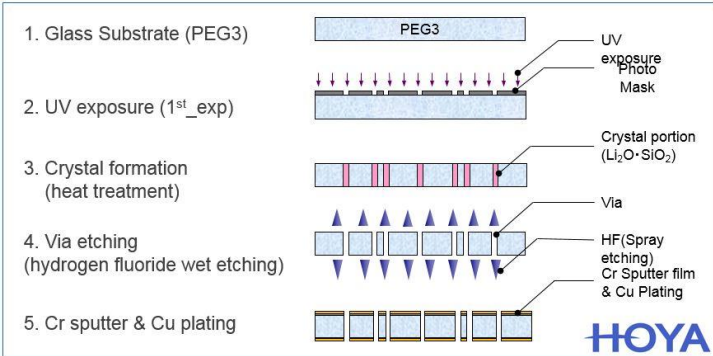
Electroluminescence light is detected by THCOBRA with 2D R-O  
 Drift time provides the third coordinate





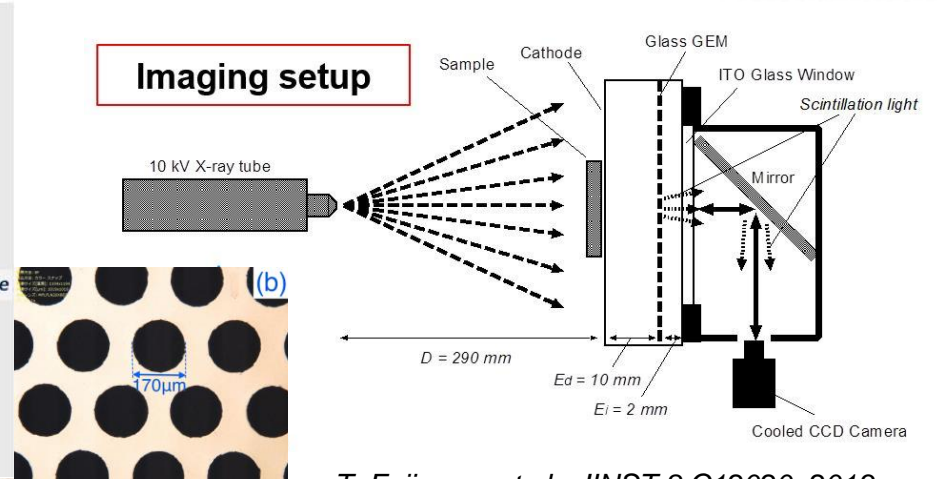
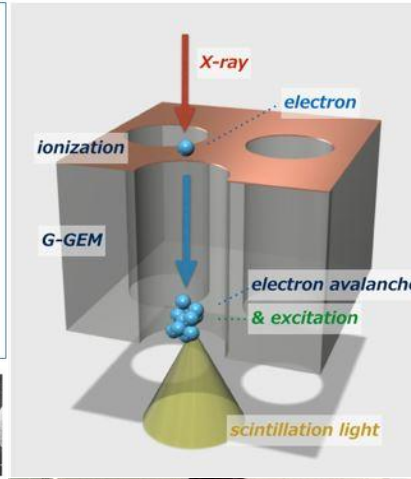


# Scintillating Glass-GEM imager



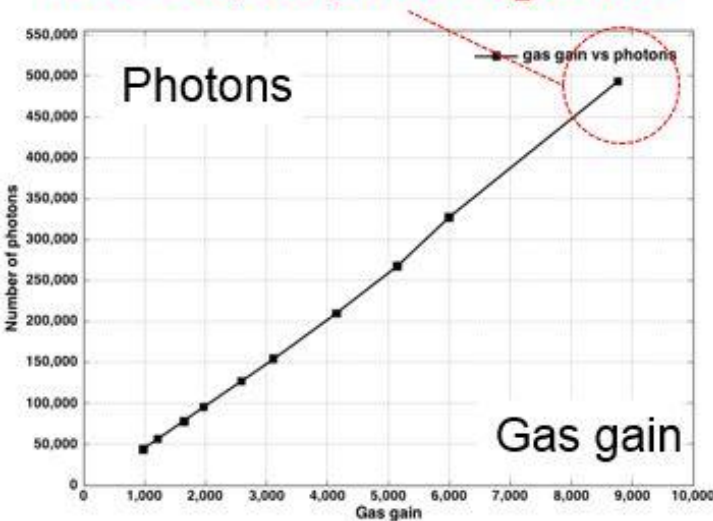
**PHOTO ETCHABLE GLASS 3 : PEG3**

- Promising technique for precise patterning
- Able to drill high aspect hole
- 680μm deep hole (ex. CERN GEM: 50μm)



T. Fujiwara, et al., JINST 8 C12020, 2013

**Max: 500,000 photons @5.9keV**



100mm

**Obtained image of leaves (2 sec integration time)<sup>[10]</sup>**

Excellent spatial resolution  $\approx 500\mu\text{m}$   
 Quick imaging of low Z material with low energy X-rays ( $\approx 7 \text{ keV}$ )

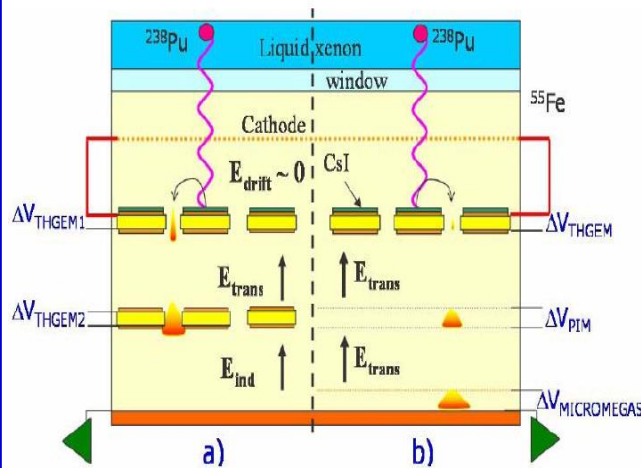


# CRYOGENIC MPGD-PDs

## Read-out elements of cryogenic noble liquid detectors

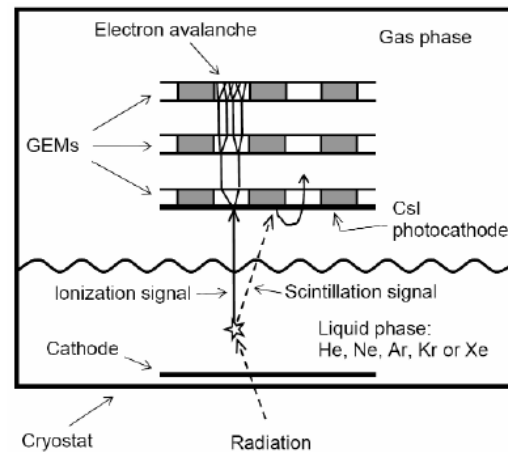
- Rear event detectors ( $\nu$ , DM)
- Detecting the scintillation light produced in the noble liquids
- Options of scintillator light and ionization charge detection by a same detector !

### with WINDOW



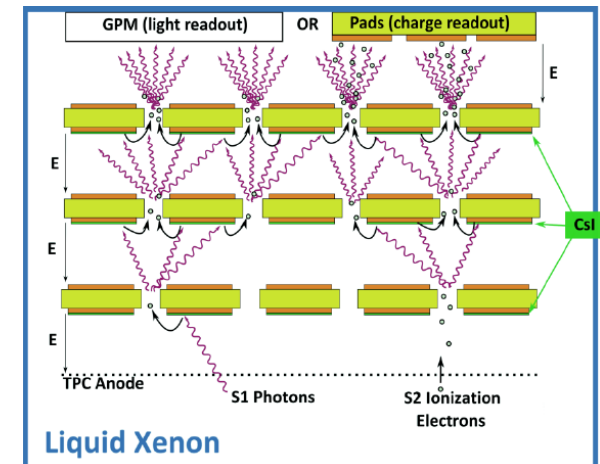
S.Duval et al., JINST 6 (2011) P04007

### WINDOWLESS (2-PHASES)



A. Bondar et al., NIMA 556 (2006) 273

### OPERATED IN THE CRYOGENIC LIQUID



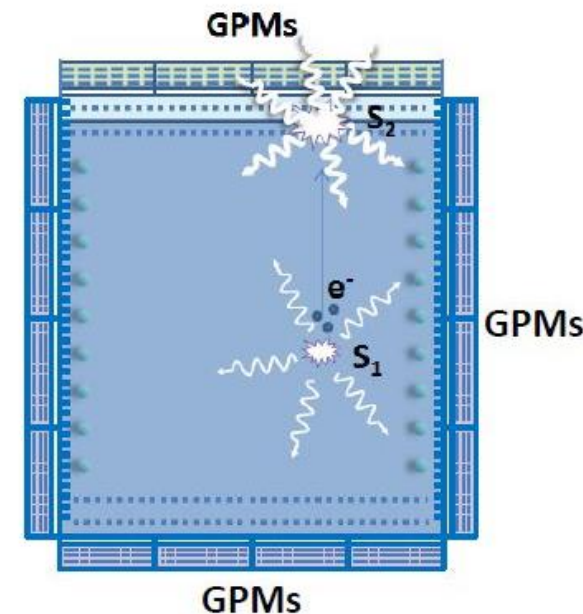
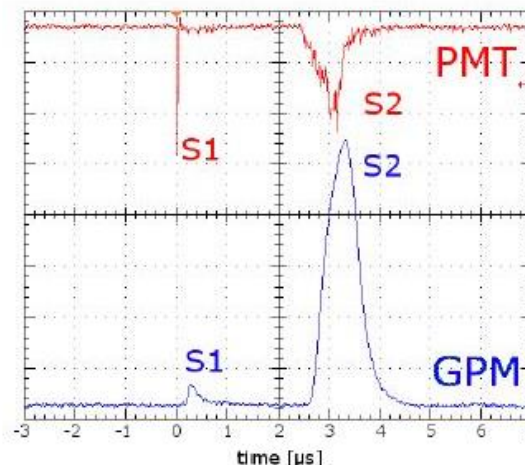
L.Arazi et al., JINST 8 (2013) C12004



# Triple THGEM Gaseous Photo-Multiplier for DM

- WIS R&D on **GPMs** for future multi-ton LXe TPCs for dark matter searches (within DARWIN)
- Aim for  **$4\pi$  coverage** – not practical with PMTs (cost, bulkiness) or SiPMs (dark count rate)
- Successful demonstration of 4" cryogenic **triple-THGEM GPM** with reflective CsI coupled to dual phase LXe TPC: ([arXiv:1509.02354](https://arxiv.org/abs/1509.02354))

- Stable gain  $\sim 10^5$
- Large dynamic range: 1 –  $O(10^3)$  photoelectrons
- 1 ns timing ( $\sim 200$  PEs)
- Expected PDE  $\sim 15\%$  after optimization



- Also: on-going R&D on **n/γ imaging** with pixilated readout ([arXiv:1501.00150](https://arxiv.org/abs/1501.00150))

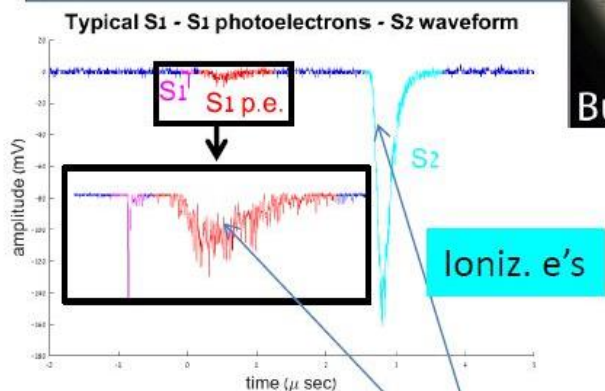
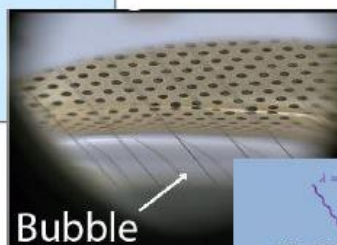
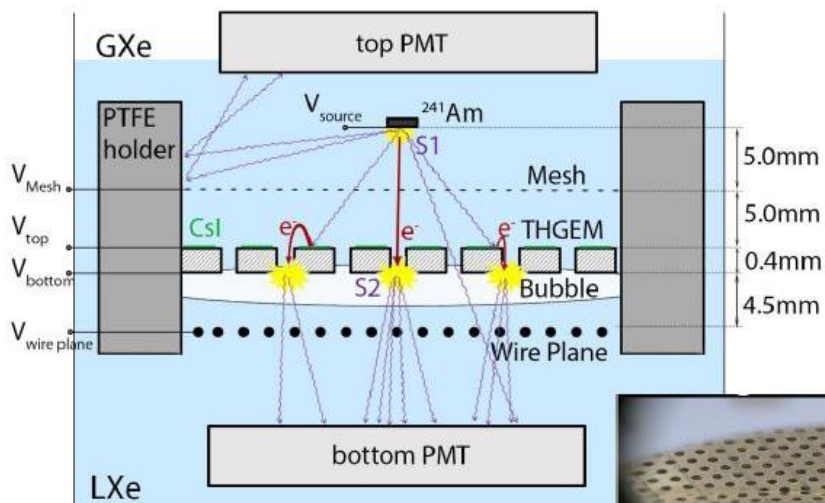


# Bubble-assisted electroluminescence in LXe

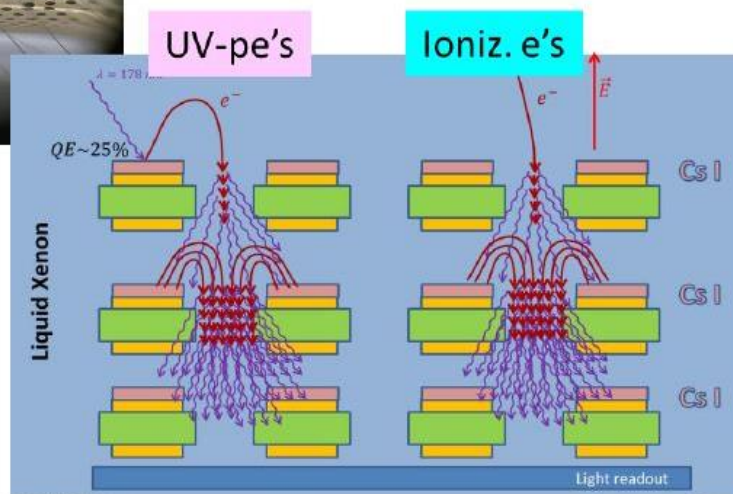
## A "local dual-phase" noble-liquid detector

TOWARDS LARGE-SCALE NOBLE-LIQU DETs

Energy resolution 5MeV alphas:  $\sigma/E=7.5\%$   
 Time resolution:  $\sigma=10\text{ns}$   
 Bubble (under THGEM, GEM) stable for days  
 CsI on THGEM: high pe extraction



UV-pe's  
 EL in bubble



A Breskin MPGD 2015 Trieste

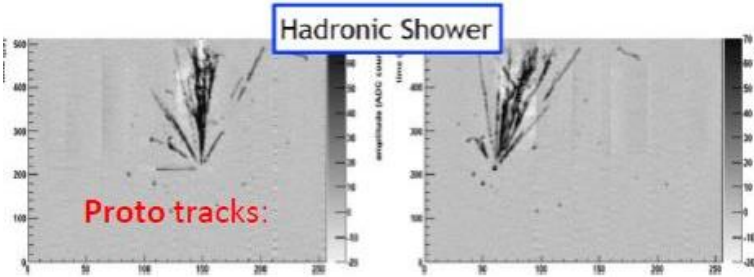
Breskin, *J. Phys. Conf. Ser.* 460(2013) 012020

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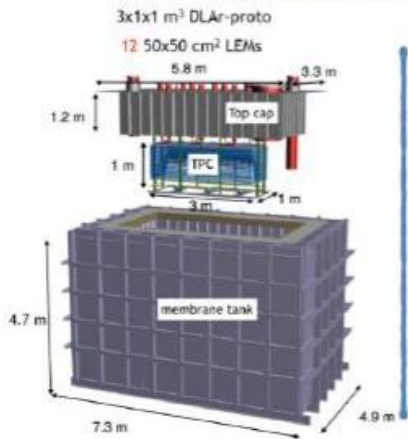
# Dual-phase LAr LEM TPC

Goal: Neutrino oscillation experiments: WA105 (on ground) and future (underground) DUNE.



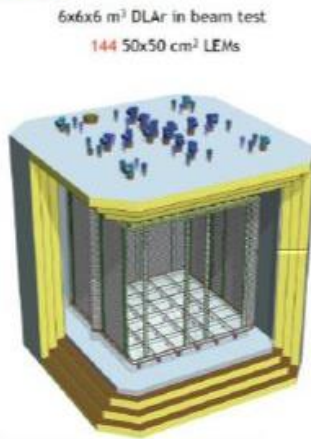
high ionization density in LAr → need low gain

## “demonstrators”

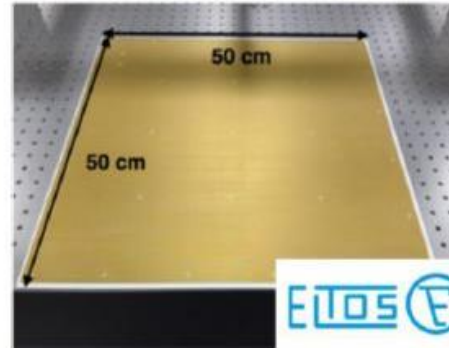
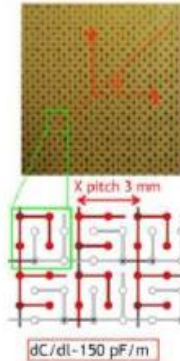


Timescale: 2015-2016

12 & 144 50x50cm<sup>2</sup> LEMs



Timescale: 2016-2019

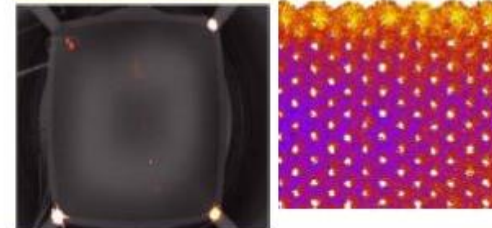


Optimised values

- 40 μm rim
- 1 mm FR4 thickness
- 500 μm diameter hole
- 800 μm hole pitch and hexagonal layout

DC: 5nA/LEM(50x50)  
Stable gain ~20 (fine)

C.Cantini et al.,  
JINST 10 P03017  
(2015)



~3500V; spark on edges  
(use COMPASS RICH solution?)  
Charging up of rims: gain stabilizes. OK

DUNE: ~3000 LEMs (50x50)

Ongoing R&D on RESISTIVE WELL concepts

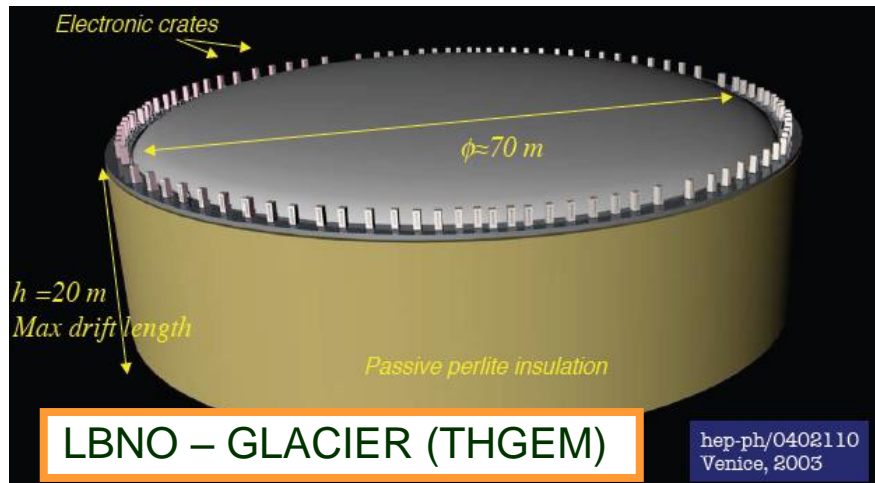


# LARGE SIZE PROJECTS



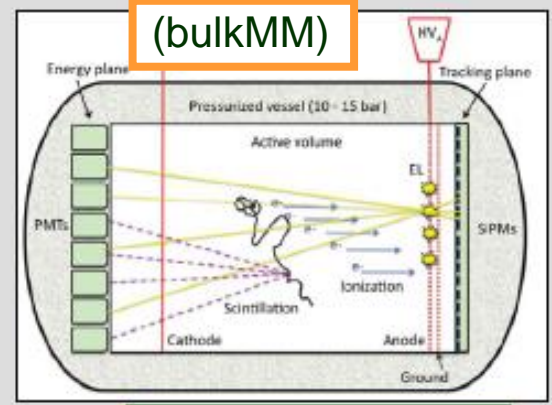
(THGEM)

XENON (dark matter)



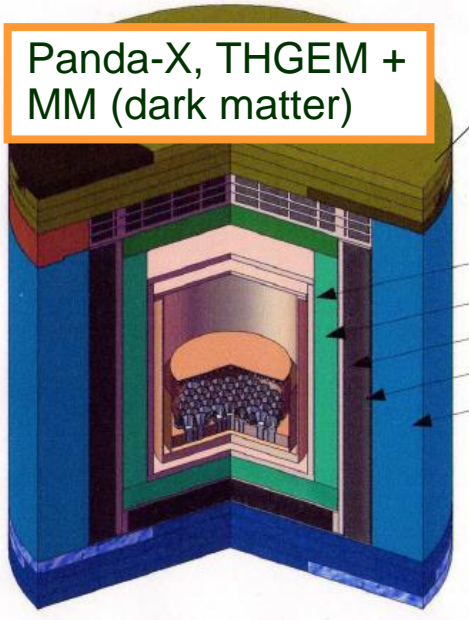
LBNO - GLACIER (THGEM)

hep-ph/0402110  
Venice, 2003

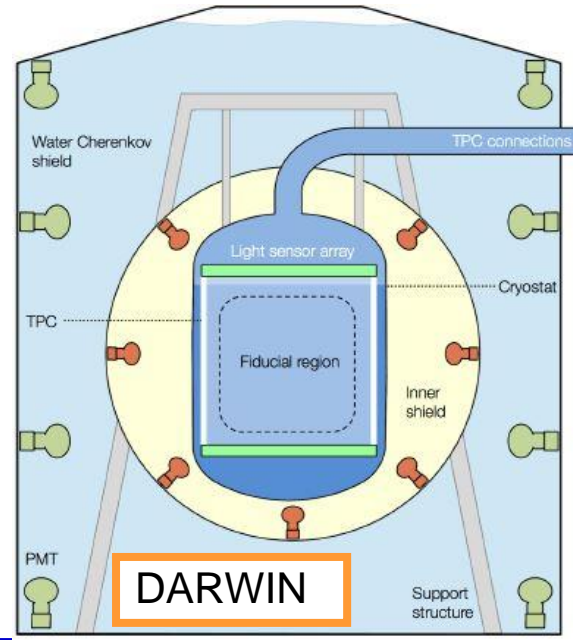


(bulkMM)

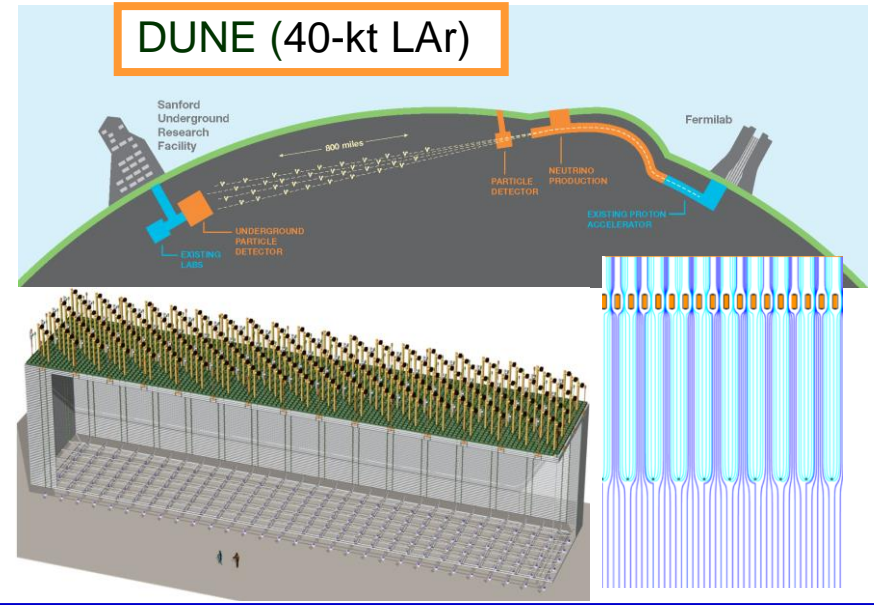
NEXT-100 (neutrino-less double beta decay)



Panda-X, THGEM + MM (dark matter)



DARWIN

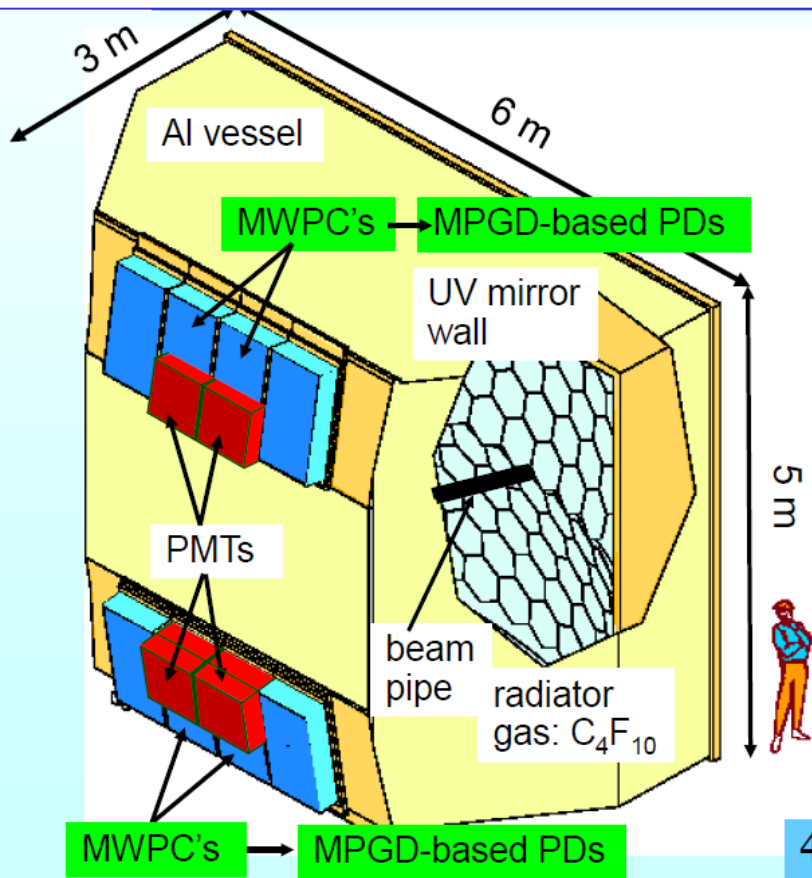


DUNE (40-kt LAr)

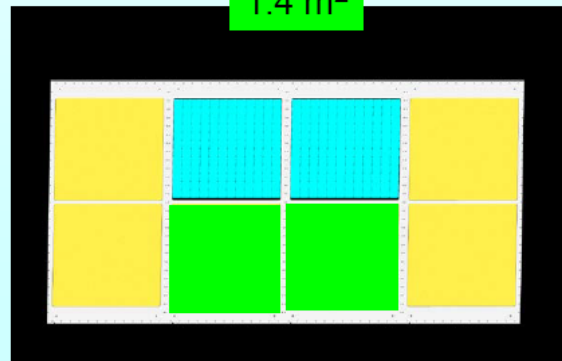
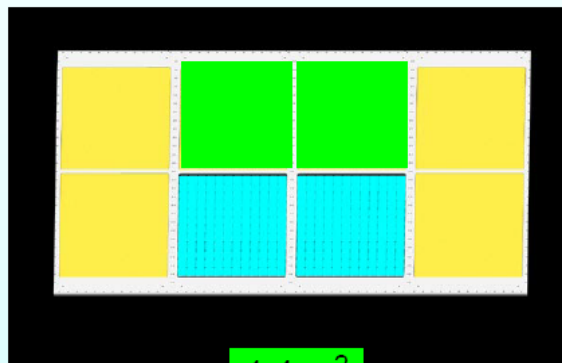


# FIRST STEP TO LARGE SIZE

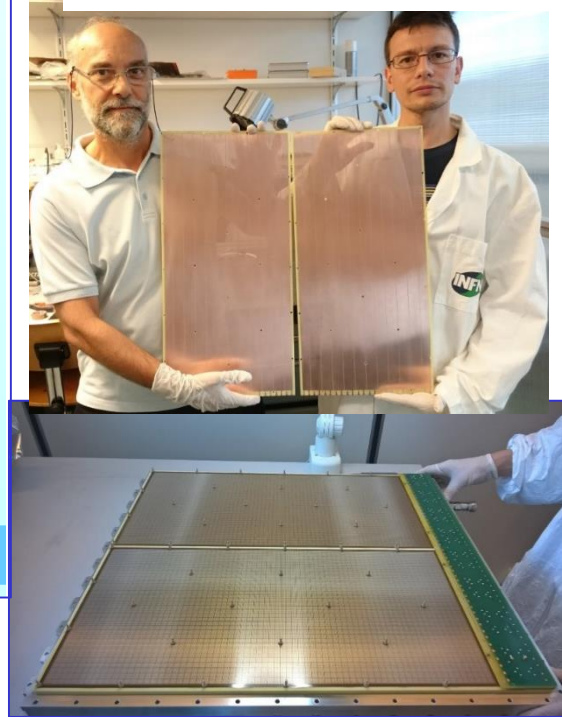
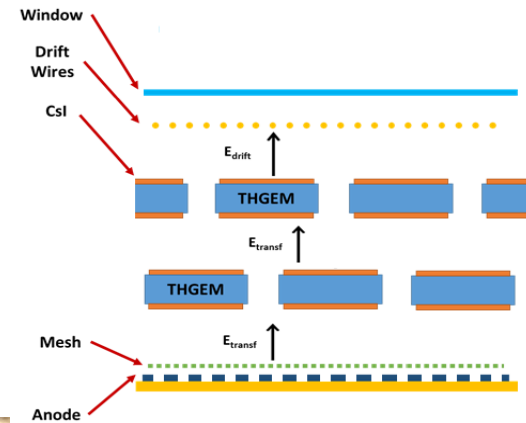
## COMPASS RICH UPGRADE



for COMPASS run 2016



4 new detectors of 600 mm x 600 mm



### presentation by Stefano Levorato



# SUMMARY / CONCLUSIONS

- **GASEOUS PHOTON DETECTORS**
  - Most effective approach to instrument large surfaces at affordable costs
  
- **MPGD-BASED PHOTON DETECTORS**
  - Allow to overcome the limitations of open geometry gaseous PDs
  - A wide effort to refine and consolidate the technology
  
- **MANY APPLICATIONS OF MPGD-BASED PHOTON DETECTORS**
  - From PID to  $\nu$ , DM, medical applications ...
  - First step toward large area: Hybrid THGEM+MM for COMPASS
  
- **BRIGHT FUTURE FOR:**
  - Inventions: new ideas, new techniques
  - Technology consolidation, new applications
  - Large scale projects